

# Stellarators

**Kenneth C. Hammond**

*Research Physicist*

*Princeton Plasma Physics Laboratory*

# About me

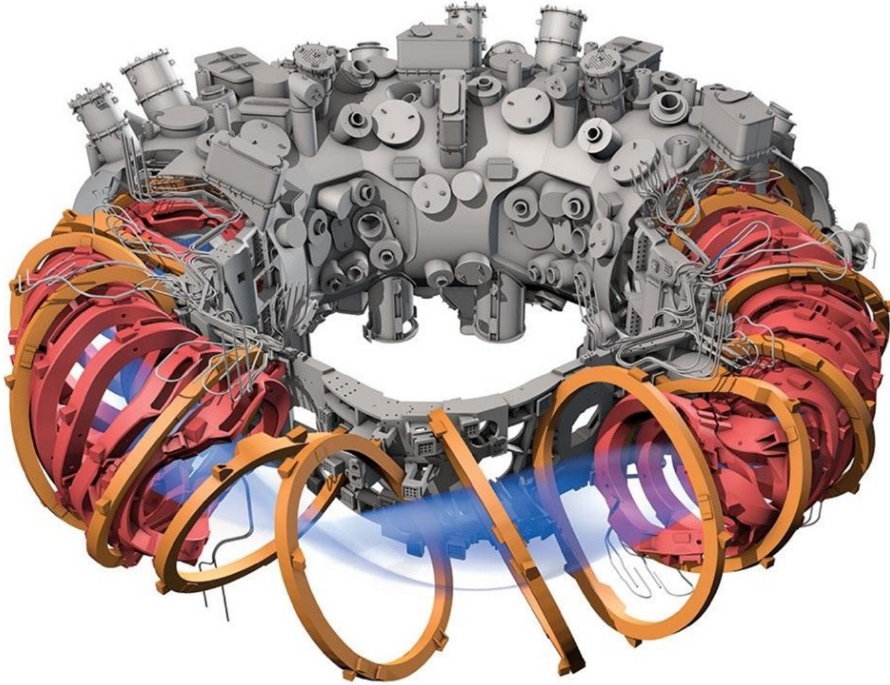
- Grew up in Saratoga Springs, NY
- Hobbies: track, cross-country, trombone
- Wanted (and still want) to do something about climate change and decarbonization
- 2008-2012: majored in physics at Harvard
  - 2011: SULI internship @ PPPL!
- 2012-2017: PhD in applied physics at Columbia University
- 2017-2019: postdoc at the Max Planck Inst. for Plasma Physics, Greifswald, Germany
- 2019-present: research physicist, PPPL
  - Still collaborating with IPP Greifswald
  - Also work on magnet design and optimization



# Introduction

# What is a stellarator?

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- Toroidal magnetic plasma confinement device
- Magnetic field forms nested, closed flux surfaces
- Magnetic field is generated primarily by magnets external to the plasma
  - Plasma current not required

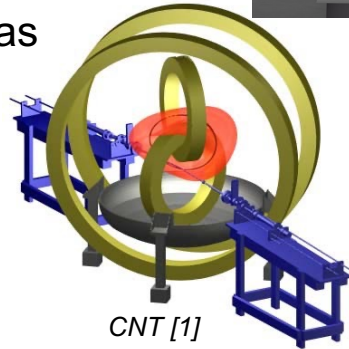
## The Wendelstein 7-X Stellarator

Max Planck Institute for Plasma Physics, Greifswald, Germany

Image: T. Klinger et al., Nucl. Fusion 2019

# Why build a stellarator?

- Generate electricity
  - Inherently steady-state operation
  - Low/no plasma current required
    - Lower vulnerability to disruptions
    - Lower recirculating power required
- Perform basic plasma research
  - Single-species plasmas
  - Pair (positron/electron) plasmas



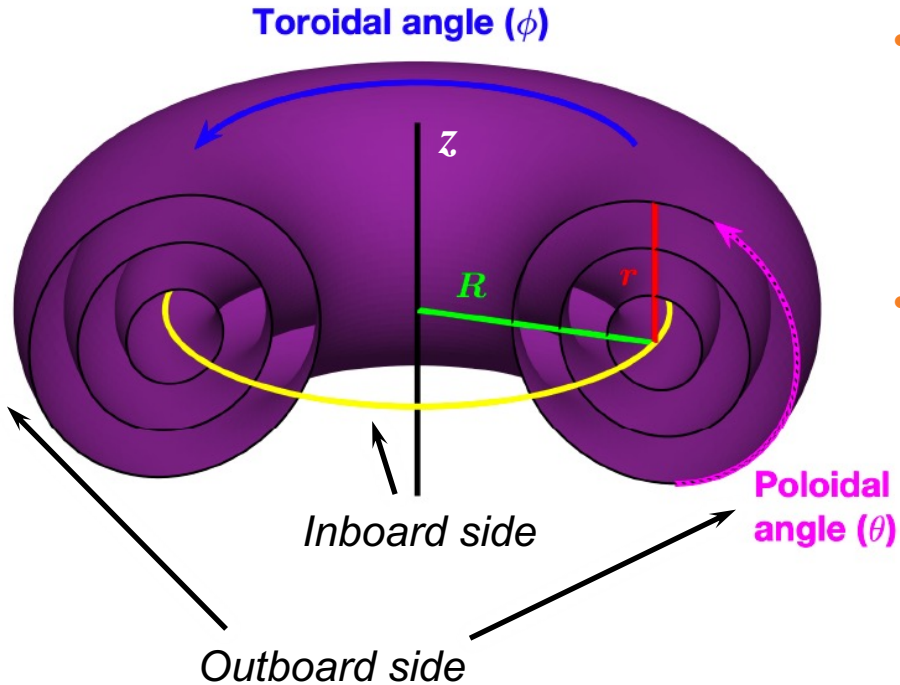
[1] J. P. Kremer et al., *Phys. Rev. Lett.* **97**, 095003 (2006)

CNT [1]

S. Woodruff, Woodruff Scientific

# Toroidal geometry and magnetic fields

# Basics of toroidal geometry

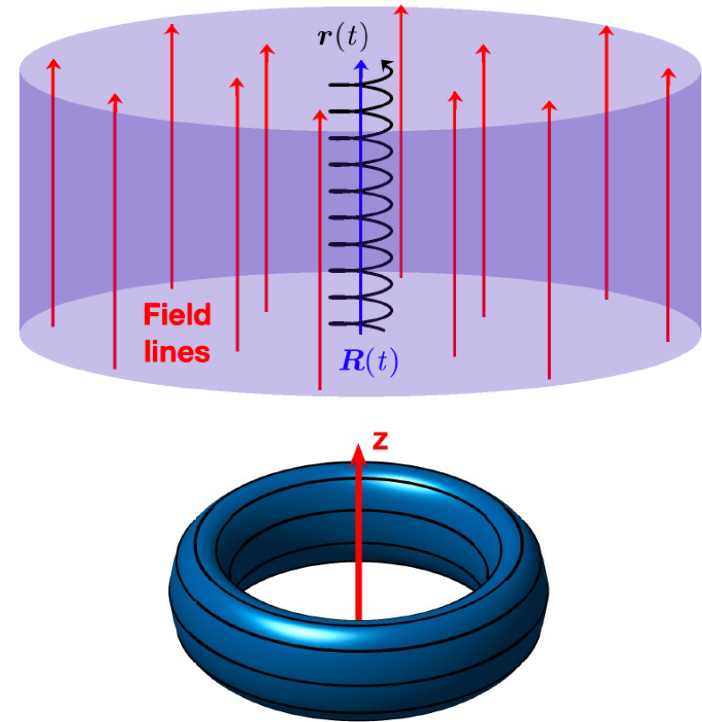


- Torus: a doughnut shape
- Cylindrical representation
  - $R$ : radial dimension; distance from  $z$  axis
  - $z$ : vertical dimension
  - $\phi$ : azimuthal/toroidal dimension
- Toroidal coordinates
  - $r$ : minor radius; distance from yellow curve
  - $\theta$ : poloidal angle; short way around the torus
  - $\phi$ : toroidal angle; long way around the torus

[2] L.-M. Imbert-Gérard et al., *Introduction to Stellarators*, SIAM (2025)

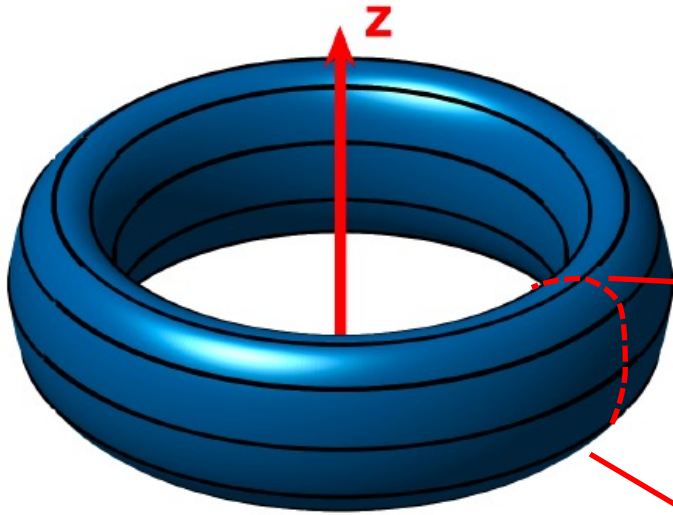
# Toroidal magnetic confinement: motivation

- Charged particles tend to gyrate around magnetic field lines  
→ confined in the direction perpendicular the magnetic field  $\mathbf{B}$
- Particles can stream freely along magnetic field lines  
→ not intrinsically confined parallel to  $\mathbf{B}$
- If the field lines wrap around form a torus
  - Particles stream quickly around the torus
  - Particles are slower to exit the torus
  - Enclosing wall is protected from the plasma (and vice versa)

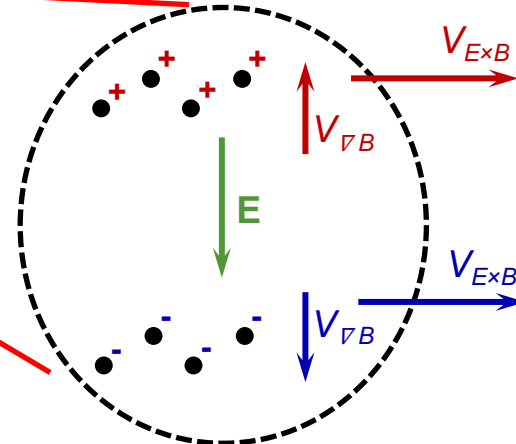


[2] L.-M. Imbert-Gérard et al., *Introduction to Stellarators*, SIAM (2025)

# A purely toroidal field is not sufficient to confine a quasineutral plasma

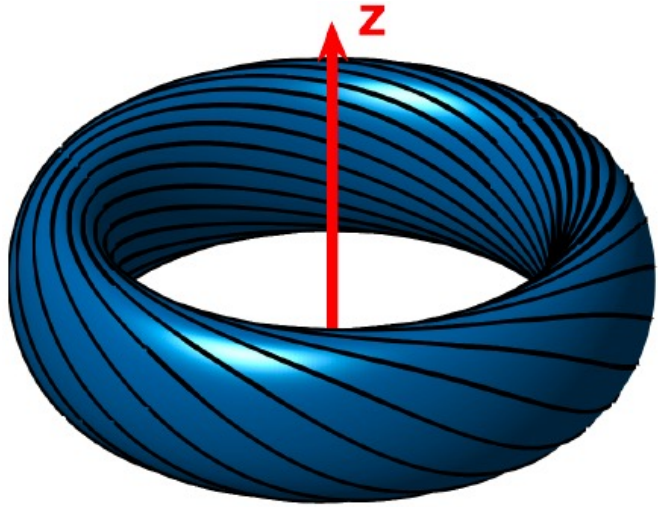


- Curvature and gradient in magnetic field cause electrons, ions to drift in opposite directions
- Charge separation creates a vertical electric field
- All particles drift outward in  $E \times B$  direction



[2] L.-M. Imbert-Gérard et al., *Introduction to Stellarators*, SIAM (2025)

# To confine a plasma in a toroidal field, the field lines must twist helically



L.-M. Imbert-Gérard et al. [2]

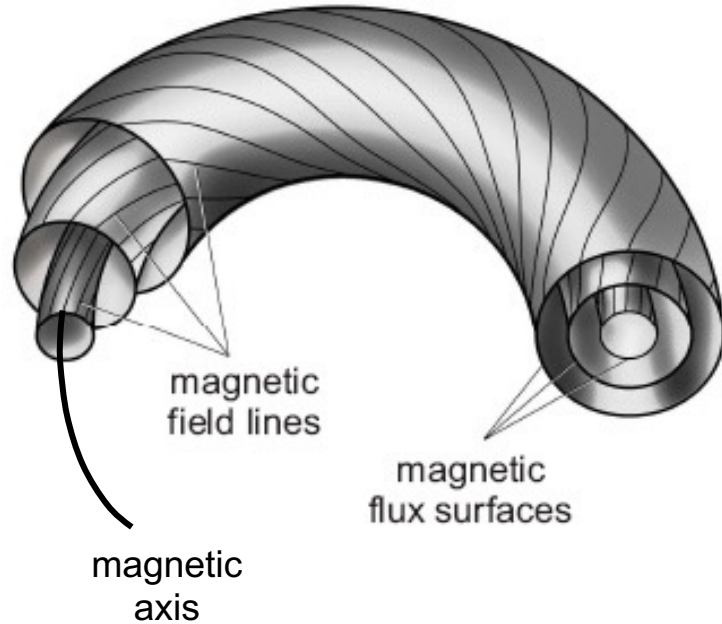


iStockPhoto

- Helical field lines sweep particles quickly from the top to the bottom of the torus and back
- Vertical drift motion persists, but alternates between inwards and outwards
- Sending the particles on helical paths is analogous to turning a honey dipper
  - Holding the dipper still: honey drips off
  - Rotating the dipper: honey remains confined

[2] L.-M. Imbert-Gérard et al., *Introduction to Stellarators*, SIAM (2025)

# Flux surfaces and rotational transform



- For good confinement, magnetic field lines should form nested *flux surfaces*
  - Surfaces surround a central curve: *magnetic axis*
- *Rotational transform*,  $\iota$  is a measure of helical twistiness of the field on each flux surface
  - In tokamaks: the *safety factor*  $q$  is  $1/\iota$
- Generating rotational transform
  - Tokamaks use a toroidal plasma current
  - Stellarators use external coils/magnets

M. Bernert [3]

[3] M. Bernert, PhD Thesis, LMU München (2013)

# How to generate a stellarator magnetic field

# Origins of the stellarator

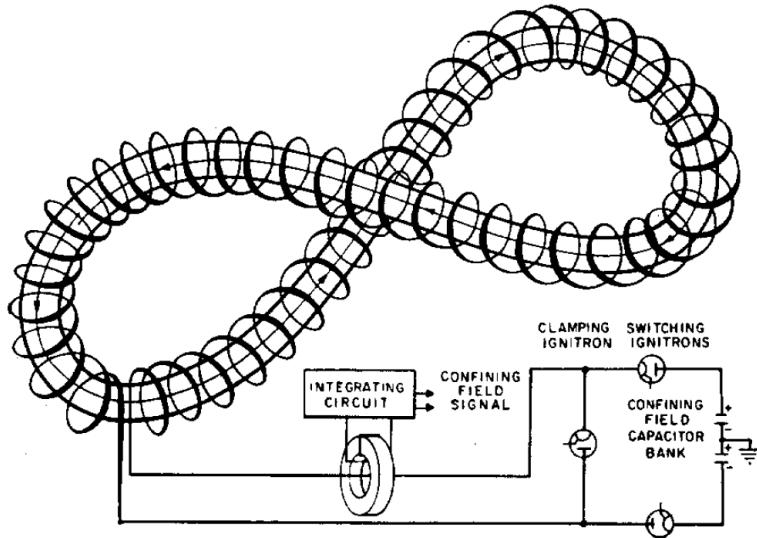
- Invented in the 1950s by Princeton astronomy professor Lyman Spitzer, Jr.
- Developed as part of Project Matterhorn, which eventually became the Princeton Plasma Physics Laboratory



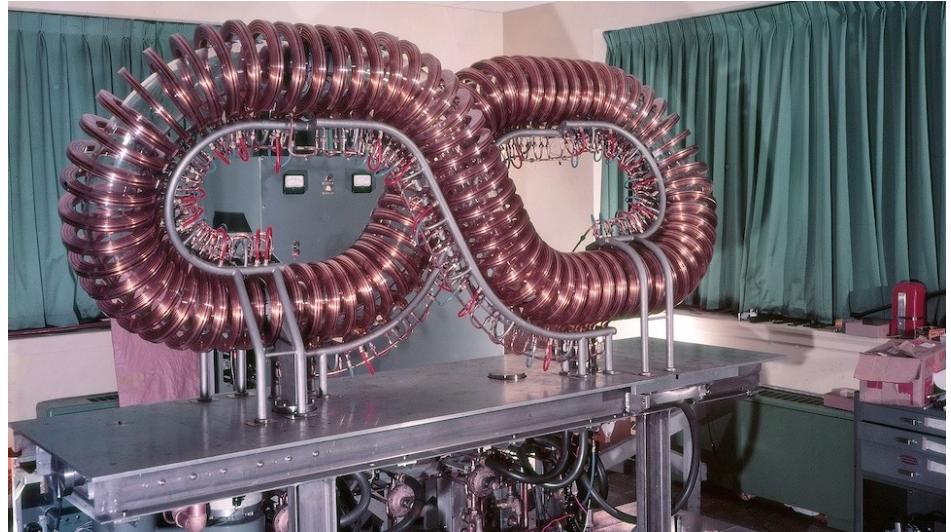
Lyman Spitzer with the Model A stellarator  
*Image: PPPL*

# How can external coils generate rotational transform?

- Spitzer's approach: solenoidal tube twisted into a figure-8
  - Crucial feature: *torsion*, or the departure of the tube's axis from a single plane



T. Coor et al. [4]

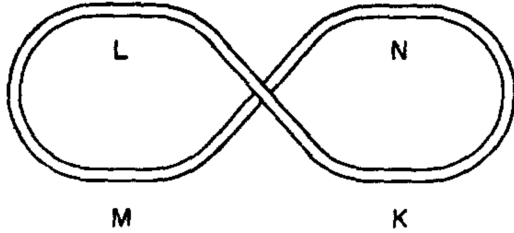


PPPL file photo

[4] T. Coor et al., *Phys. Fluids* 1, 411 (1958)

# Rotational transform in a figure-8 solenoid [5]

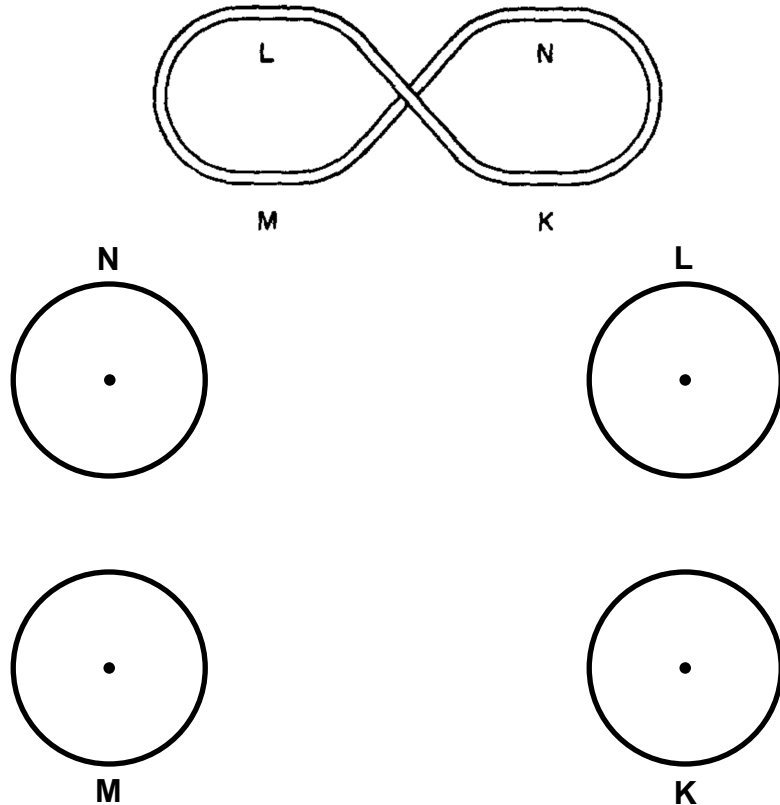
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- Magnetic field in the figure-8 is approximately *solenoidal*, i.e. everywhere parallel to the central axis of the tube

[5] L. Spitzer, *Phys. Fluids* **1**, 253 (1958)

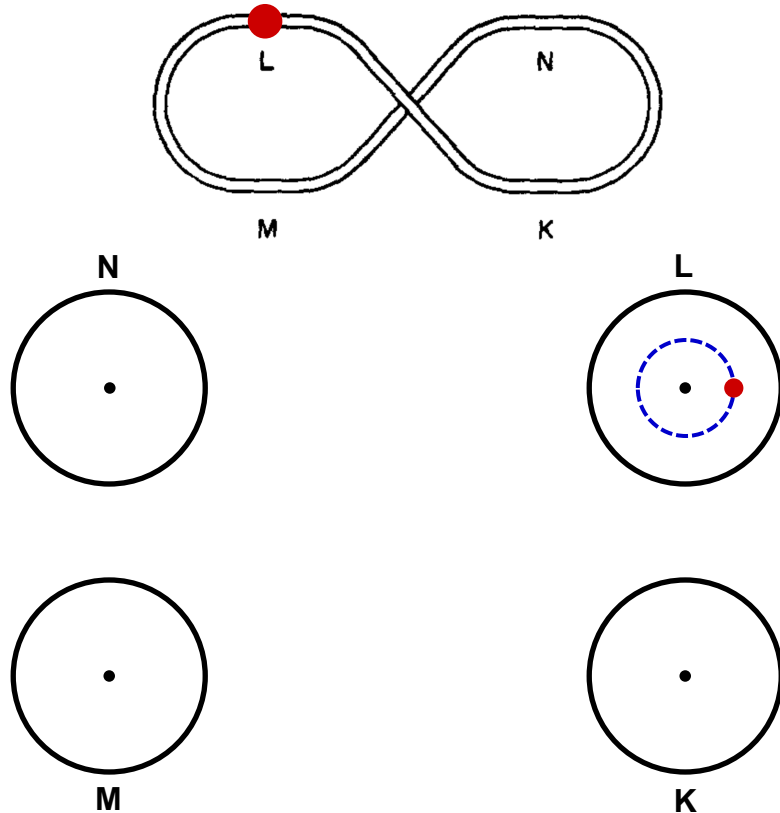
# Rotational transform in a figure-8 solenoid



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[5] L. Spitzer, *Phys. Fluids* **1**, 253 (1958)

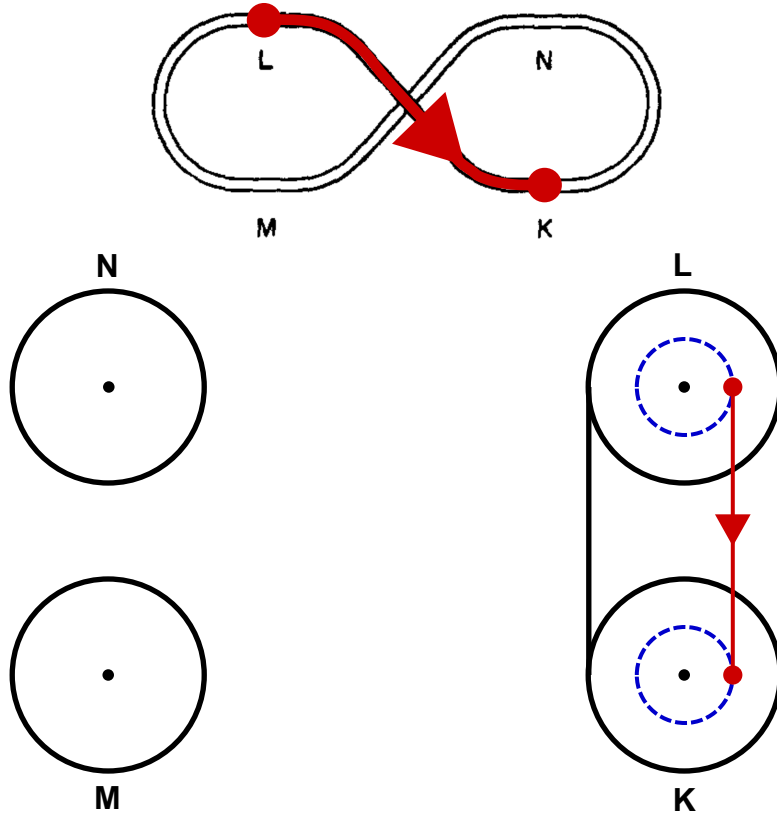
# Tracing the field line around the track



- Initialize a field line at a point in cross-section  $L$  a distance  $r$  from the central axis

[5] L. Spitzer, *Phys. Fluids* **1**, 253 (1958)

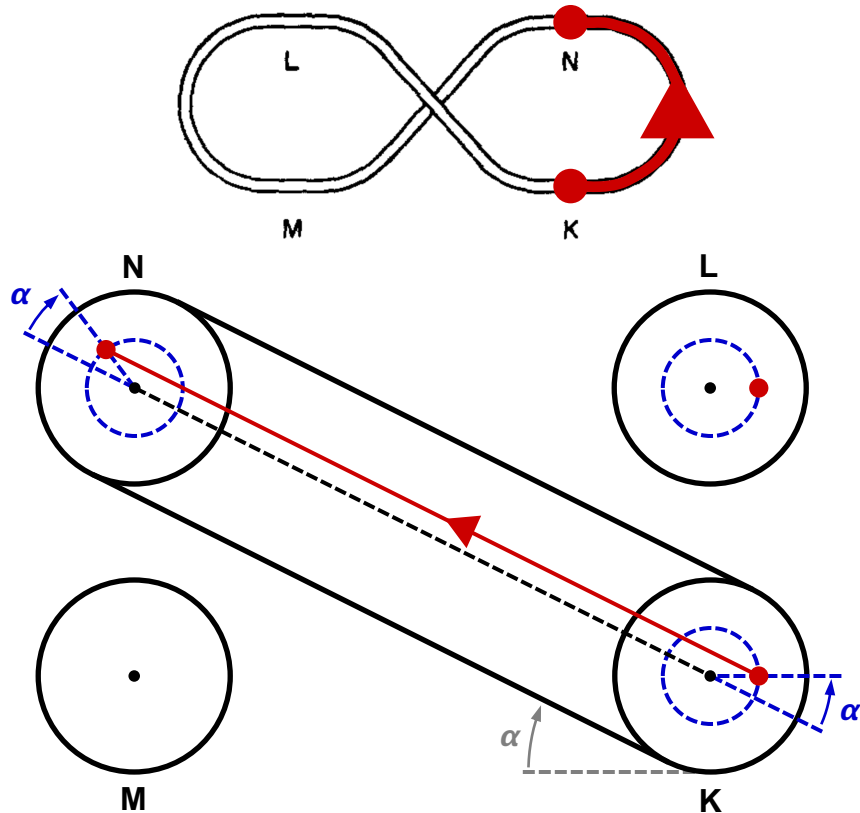
# First “downhill” ( $L-K$ )



- Field line trajectory
  - Remains on vertical plane, at “3:00” with respect to magnetic axis

[5] L. Spitzer, *Phys. Fluids* **1**, 253 (1958)

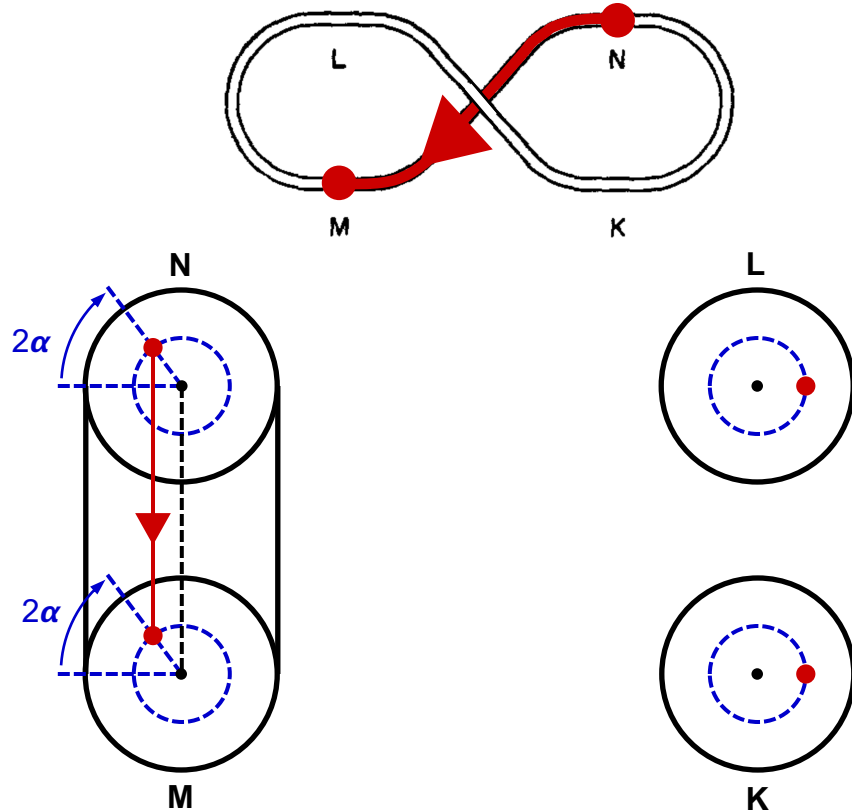
# First upward semicircle bend ( $K-N$ )



- Field line trajectory
  - Travels parallel to the plane of curve (tilted an angle  $\alpha$  to the horizontal)
  - Remains at an angle  $\alpha$  above the plane of the curve relative to magnetic axis

[5] L. Spitzer, *Phys. Fluids* **1**, 253 (1958)

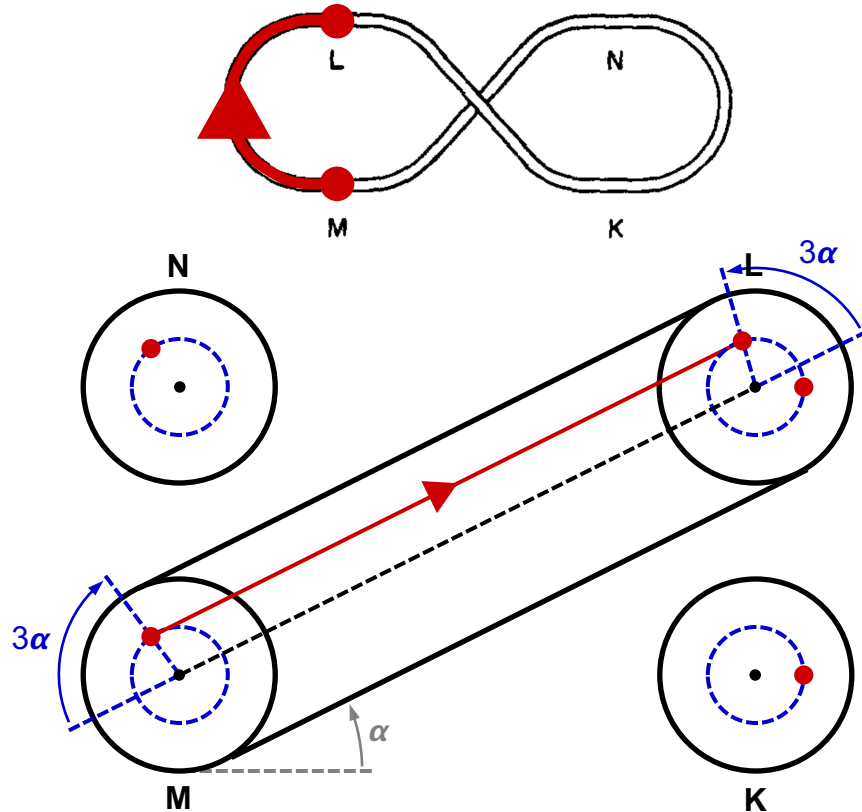
# Second “downhill” ( $N-M$ )



- Field line trajectory
  - Remains on vertical plane at an angle  $2\alpha$  above the horizontal plane about the magnetic axis

[5] L. Spitzer, *Phys. Fluids* **1**, 253 (1958)

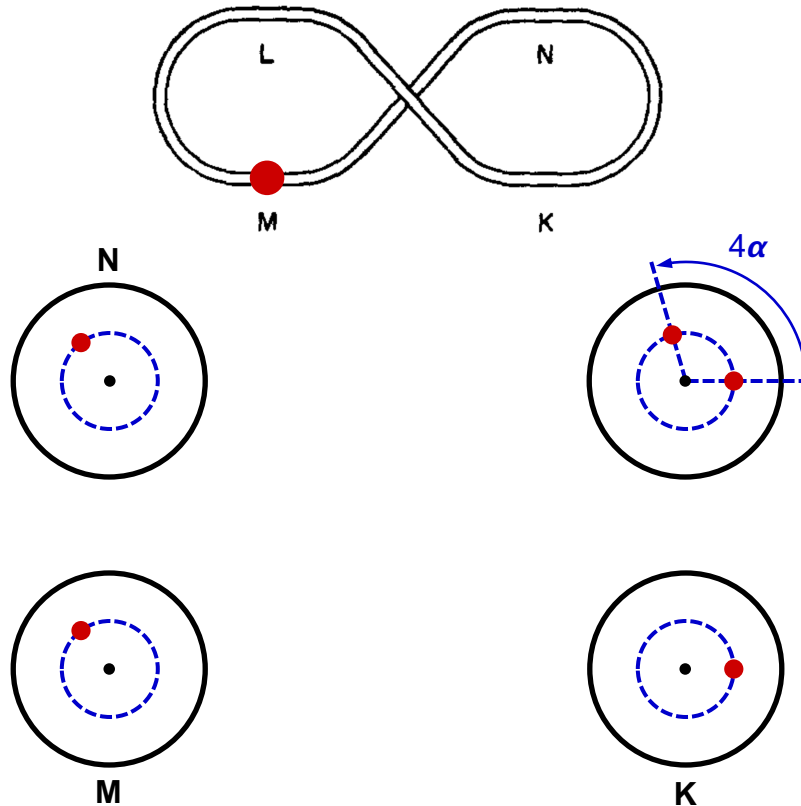
# Second upward semicircular bend ( $M-L$ )



- Field line trajectory
  - Travels parallel to the plane of curve (tilted an angle  $\alpha$  to the horizontal)
  - Remains at an angle  $3\alpha$  above the plane of the curve relative to the magnetic axis

[5] L. Spitzer, *Phys. Fluids* **1**, 253 (1958)

# Field line has rotated poloidally upon its return to the starting plane!

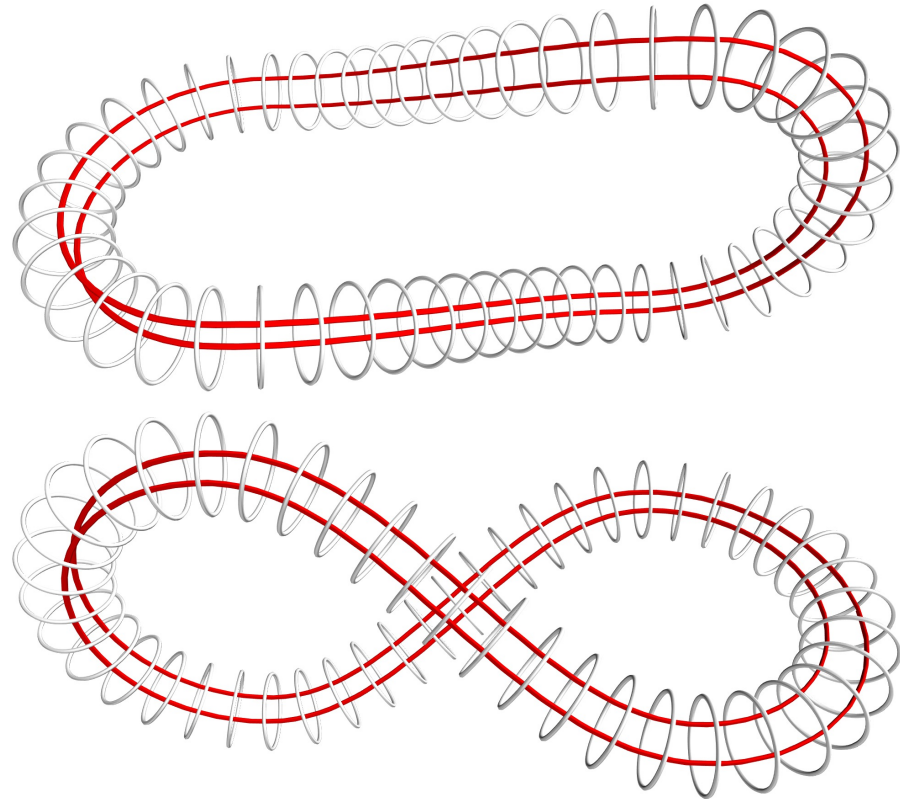
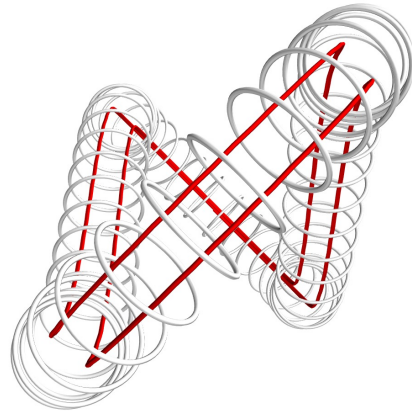


- “Puncture point” has rotated an angle of  $4\alpha$  from its starting location with respect to the central axis
- This corresponds to a *rotational transform*  $\iota = 4\alpha/(2\pi)$
- Same argument applies for any radial distance from the central axis → every flux surface exhibits the same rotational transform

[5] L. Spitzer, *Phys. Fluids* 1, 253 (1958)

# Numerical model of figure-8 stellarator agrees very well with simple geometric predictions

- Model set up in SIMSOPT [6,7]
- Magnetic field calculated from 50 ideal circular coils in figure-8 pattern
- ODE solver used to trace field lines

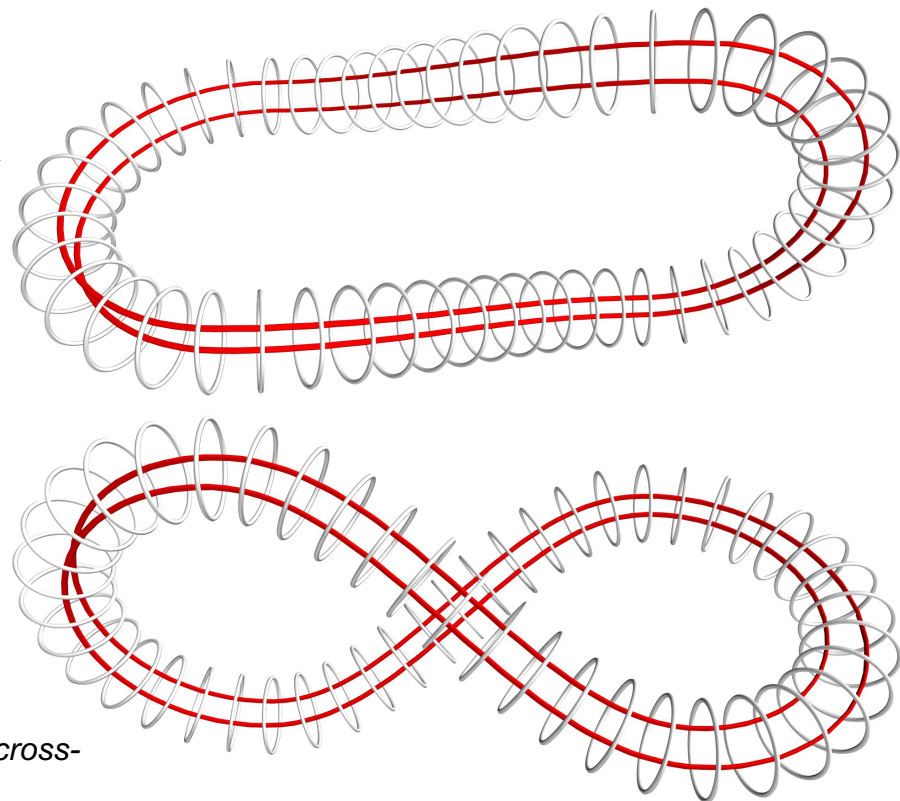
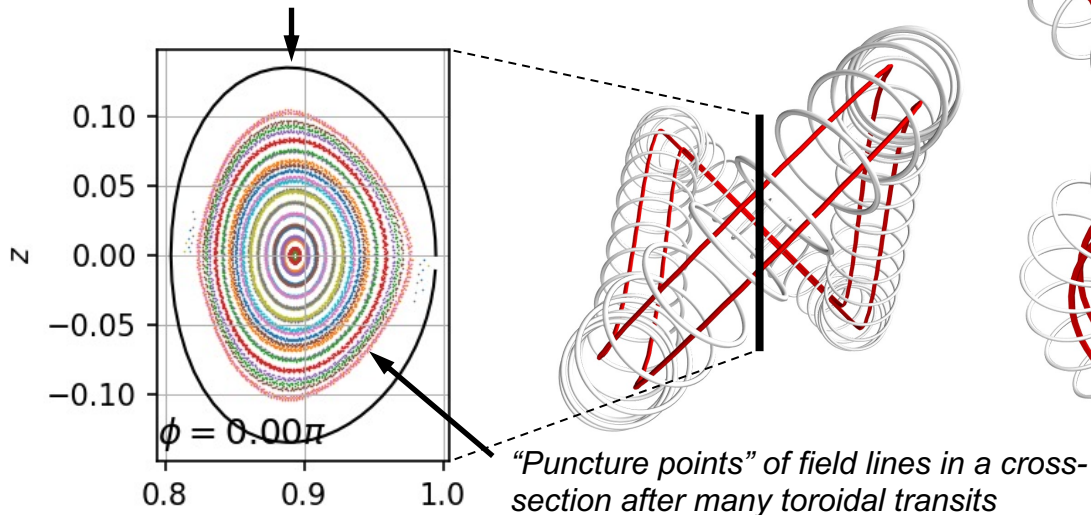


[6] M. Landreman et al., *J. Open Source Soft.* **6**, 3525 (2021)

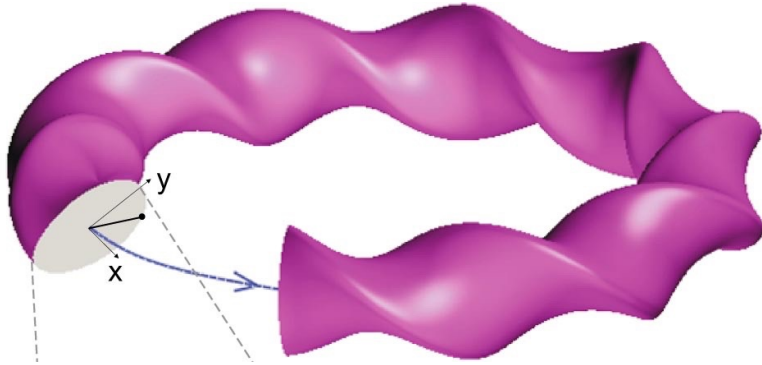
[7] <https://github.com/hiddenSymmetries/simsopt>

# Numerical model of figure-8 stellarator agrees very well with simple geometric predictions

- $\alpha = \pi/4 \rightarrow$  prediction:  $\iota = 1/2$ 
  - 3D field line visualization: shows that field lines “bite their tails” after two transits
  - *Poincaré plot*: shows that field lines are confined to flux surfaces

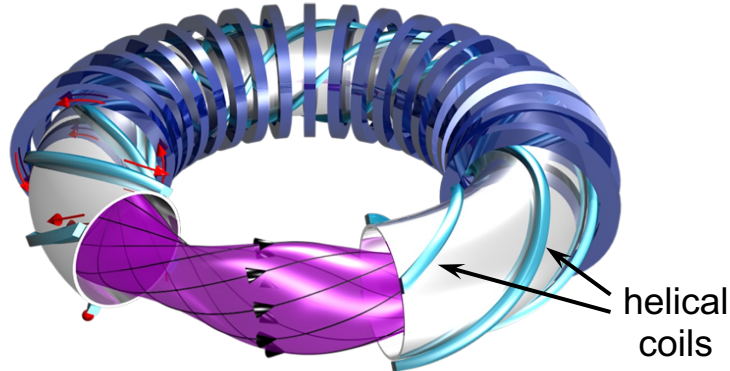


# Rotational transform can also arise from a helically rotating poloidal cross-section

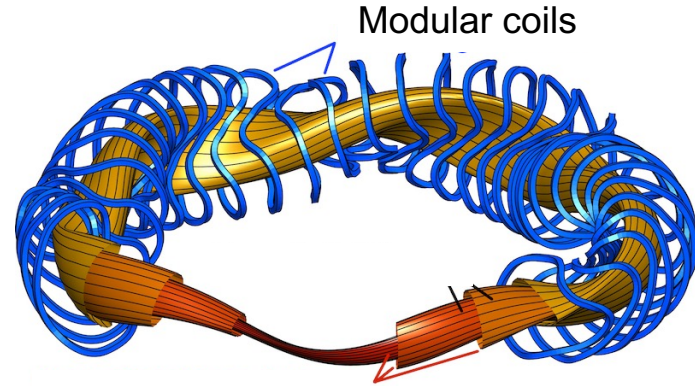


*P. Helander, Rep. Prog. Mod. Phys., 2014*

- This effect can be achieved with suitable coil shaping
  - Helical windings
  - Modular coils with non-planar shapes



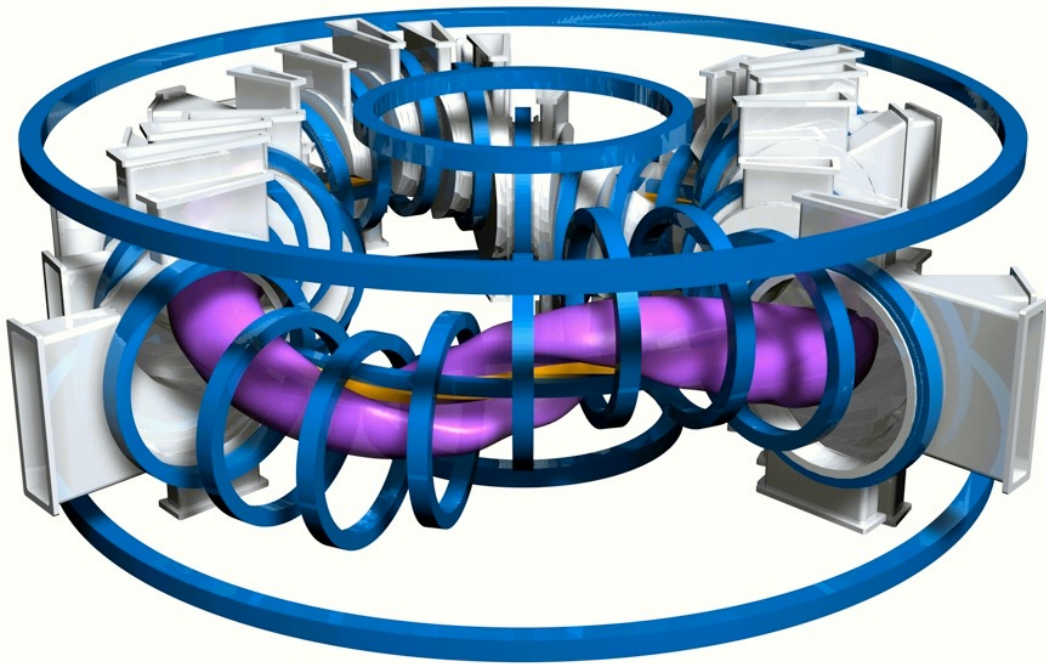
*Max Planck Institute for Plasma Physics*



*M. Landreman*

# Stellarators can take many forms

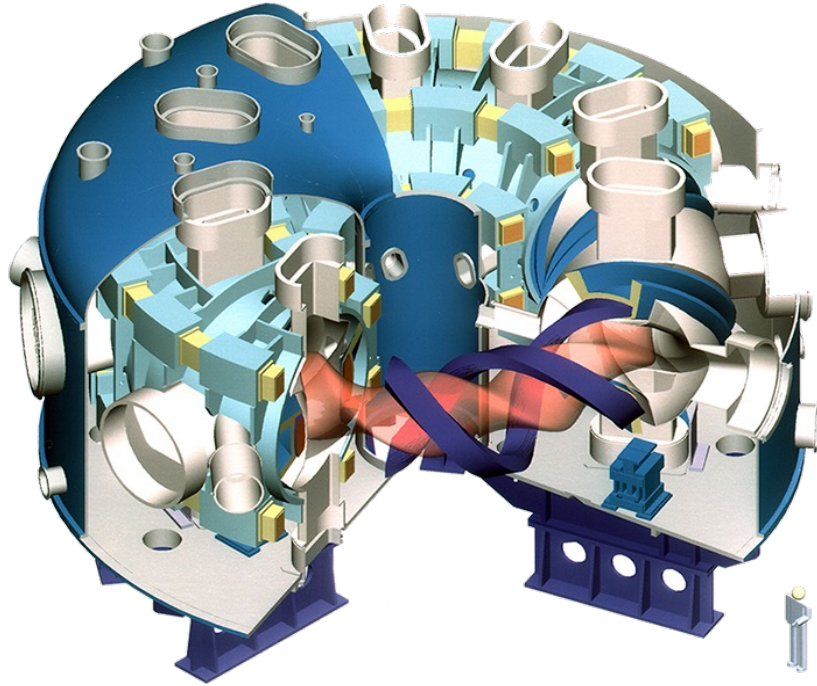
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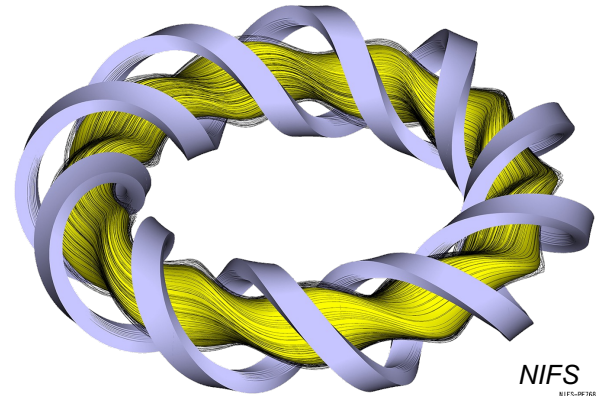
*CIEMAT*

- TJ-II, CIEMAT, Madrid, Spain
  - “Helic” configuration
  - Most coils are circular and planar

# Stellarators can take many forms



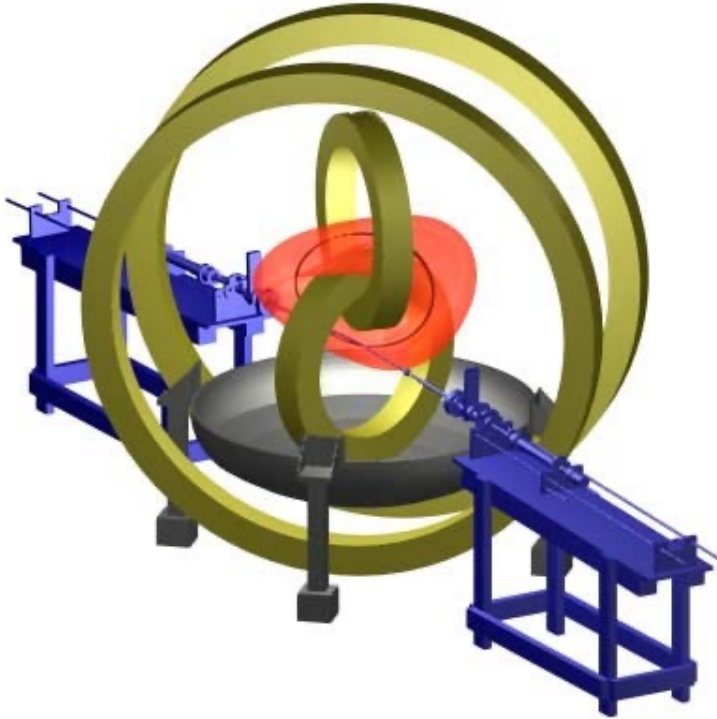
- Large Helical Device (LHD), National Institute for Fusion Science, Toki, Japan
  - Two superconducting helical coils provide most of the magnetic field
  - Concluded final campaign in December 2025



NIFS  
NIFS-PE768

# Stellarators can take many forms

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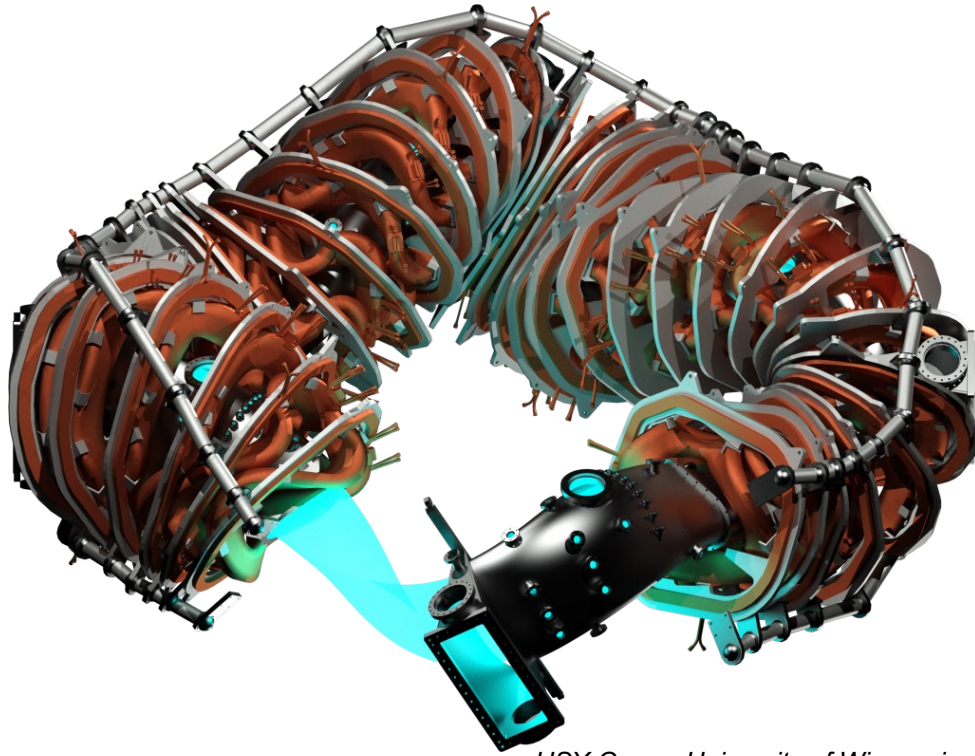


- CNT, Columbia University, New York, NY, USA
  - Four circular planar coils
  - Two coils are interlocked and tilted

*J. P. Kremer et al., Phys. Rev. Lett. 2006*

# Stellarators can take many forms

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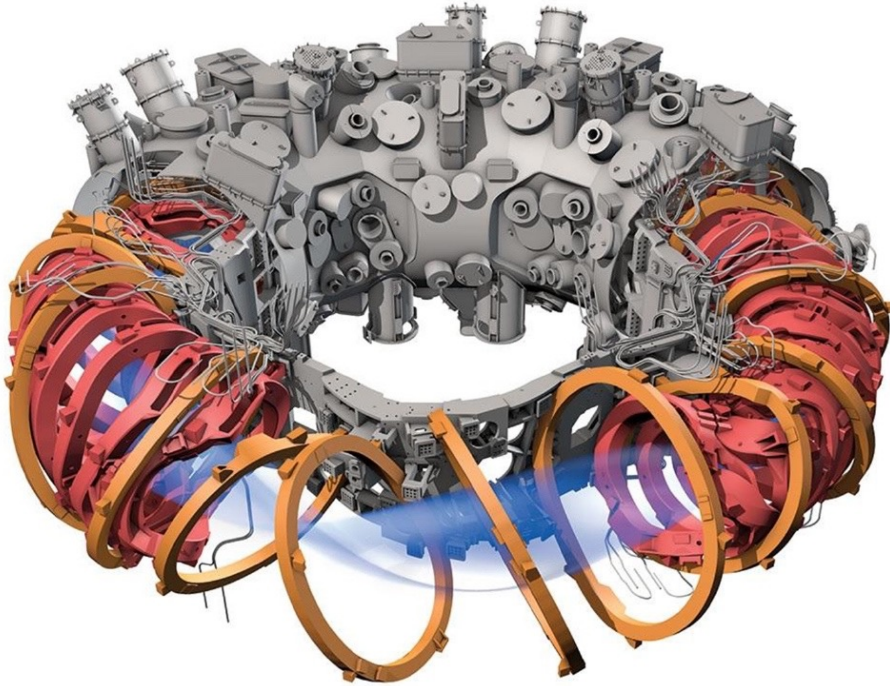


*HSX Group, University of Wisconsin*

- Helicly Symmetric Experiment (HSX), University of Wisconsin, Madison, WI, USA
  - Combination of Modular, “wiggly” coils and planar coils supply the main field
  - Modular coils optimized for good confinement

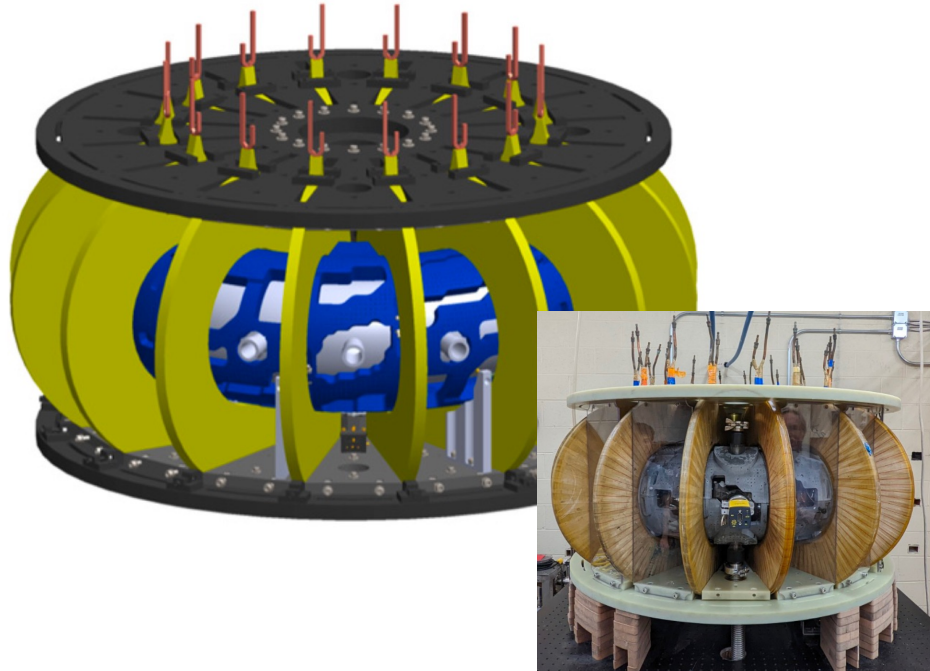
# Stellarators can take many forms

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- Wendelstein 7-X (W7-X), Max Planck Institute for Plasma Physics, Greifswald, Germany
  - Most advanced stellarator built to date
  - Modular coils optimized for good confinement
  - Superconducting coils

# Stellarators can take many forms



- MUSE, PPPL
  - First stellarator that uses permanent magnets to produce 3D field shaping
    - Nearly 10,000 magnets altogether
  - Planar coils supply toroidal field
  - Assembled primarily by students
  - Quasi-axisymmetric

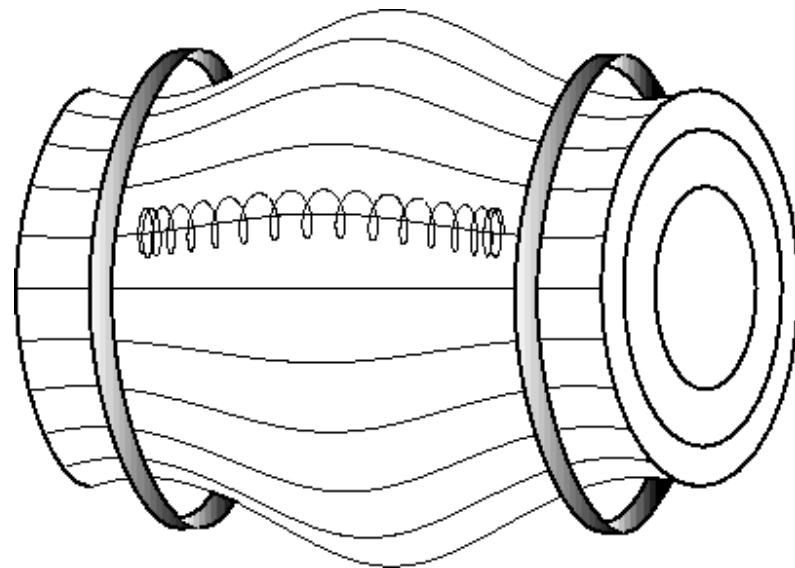
*T. Qian et al., Nucl. Fusion 2022*

*T. Qian et al., Nucl. Fusion 2023*

# Keeping particle orbits confined

# Rotational transform is necessary but not sufficient for good confinement

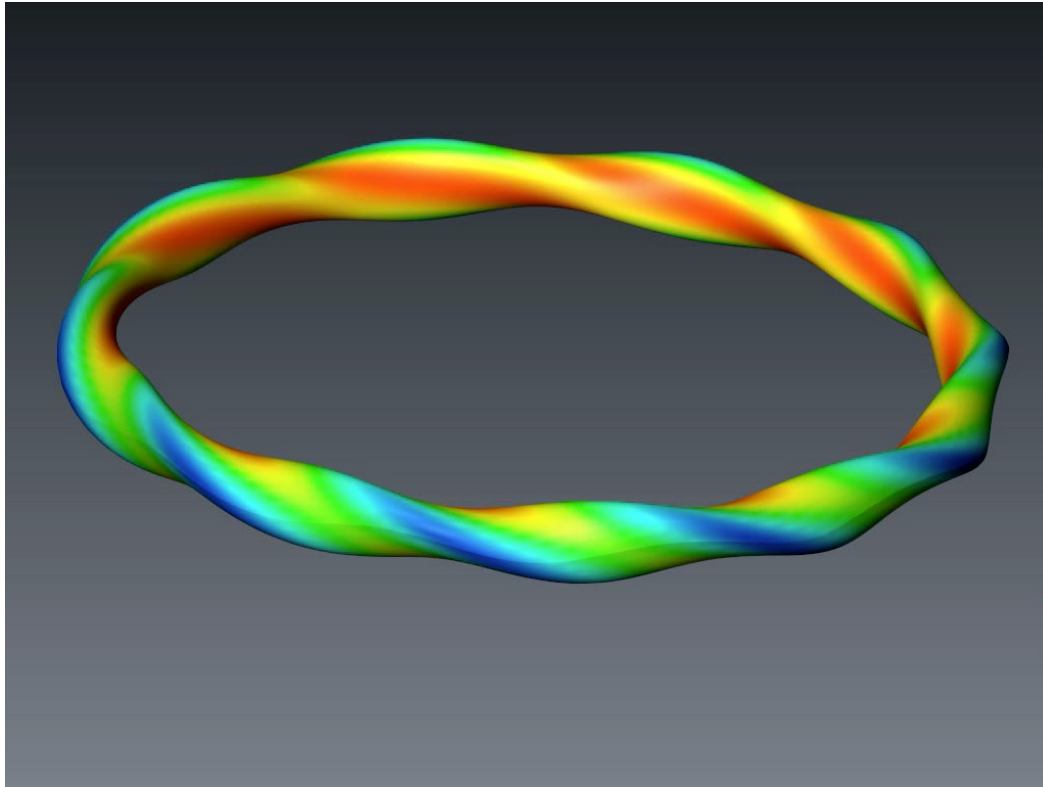
- Early stellarators exhibited poor energy and particle confinement despite producing flux surfaces with rotational transform
- Helical field lines generally confine *passing* particles well, but not necessarily *trapped* particles
  - *Trapped* particles bounce back and forth between points with high field strength due to magnetic mirror effect
  - *Passing* particles have a high enough  $v_{\parallel}/v_{\perp}$  orbit toroidally without bouncing



R. Fitzpatrick

# Orbits in a stellarator with poor trapped-particle confinement

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# Particles trapped in a magnetic mirror exhibit constants of motion

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- First adiabatic invariant: magnetic moment  $\mu$

$$\mu = \frac{mv_{\perp}^2}{2B} \quad v_{\perp} = \textit{perpendicular velocity}$$

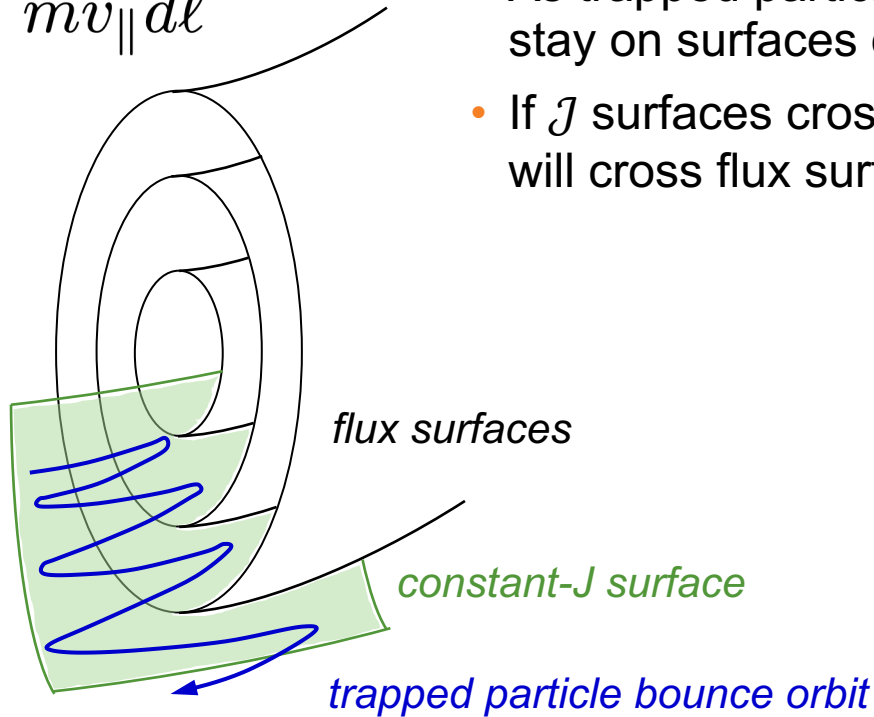
- Second adiabatic invariant:  $\mathcal{J}$

$$\mathcal{J} = \int_{\ell_1}^{\ell_2} mv_{\parallel} d\ell \quad v_{\parallel} = \textit{parallel velocity}$$

- For a trapped particle in a magnetic mirror,  $\mathcal{J}$  tends to remain constant for each particle as it bounces
- When drifting (due to curvature,  $\nabla B$ , etc.), particles follow trajectories that conserve  $\mathcal{J}$

# Conservation of $\mathcal{J}$ impacts how trapped particles drift among flux surfaces

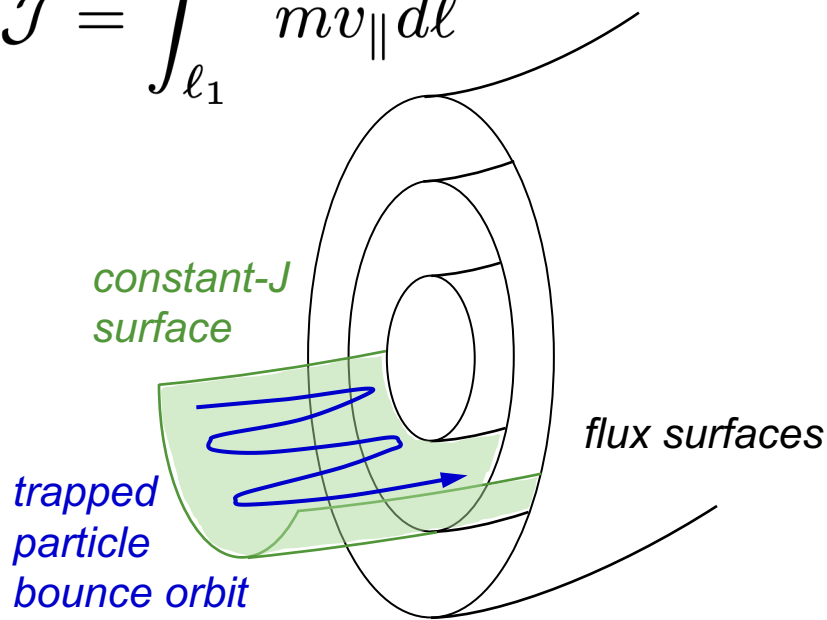
$$\mathcal{J} = \int_{l_1}^{l_2} m v_{\parallel} dl$$



- As trapped particles drift across field lines, they stay on surfaces of constant  $\mathcal{J}$
- If  $\mathcal{J}$  surfaces cross flux surfaces, trapped particles will cross flux surfaces (and eventually escape)

# If constant- $\mathcal{J}$ surfaces coincide with flux surfaces, trapped particles remain confined

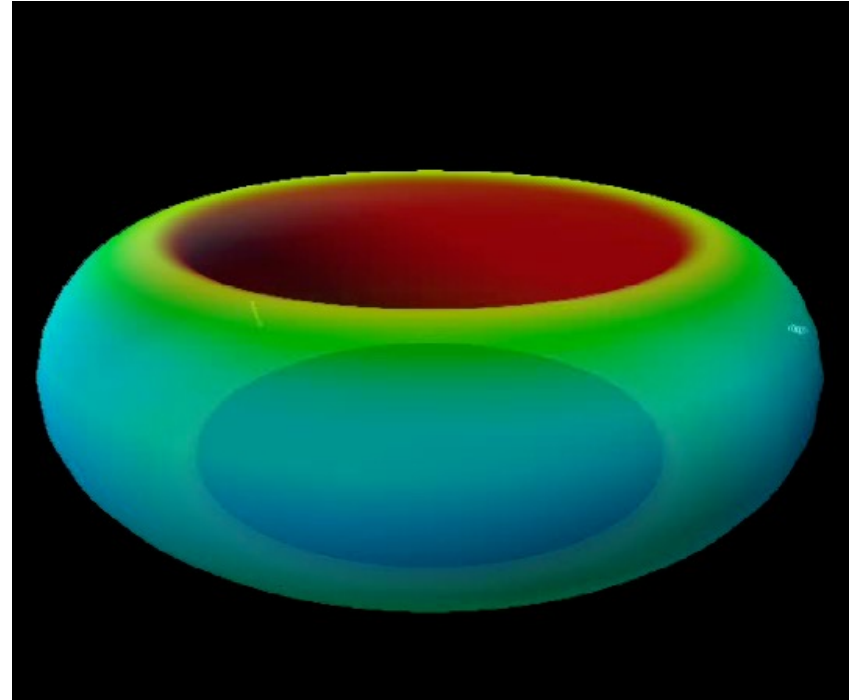
$$\mathcal{J} = \int_{l_1}^{l_2} m v_{\parallel} dl$$



- Stellarators exhibiting this property are referred to as *omnigenous*
- Equivalent definition: an omnigenous field is one with zero time-averaged radial magnetic drift

# Axisymmetric devices are automatically omnigenous

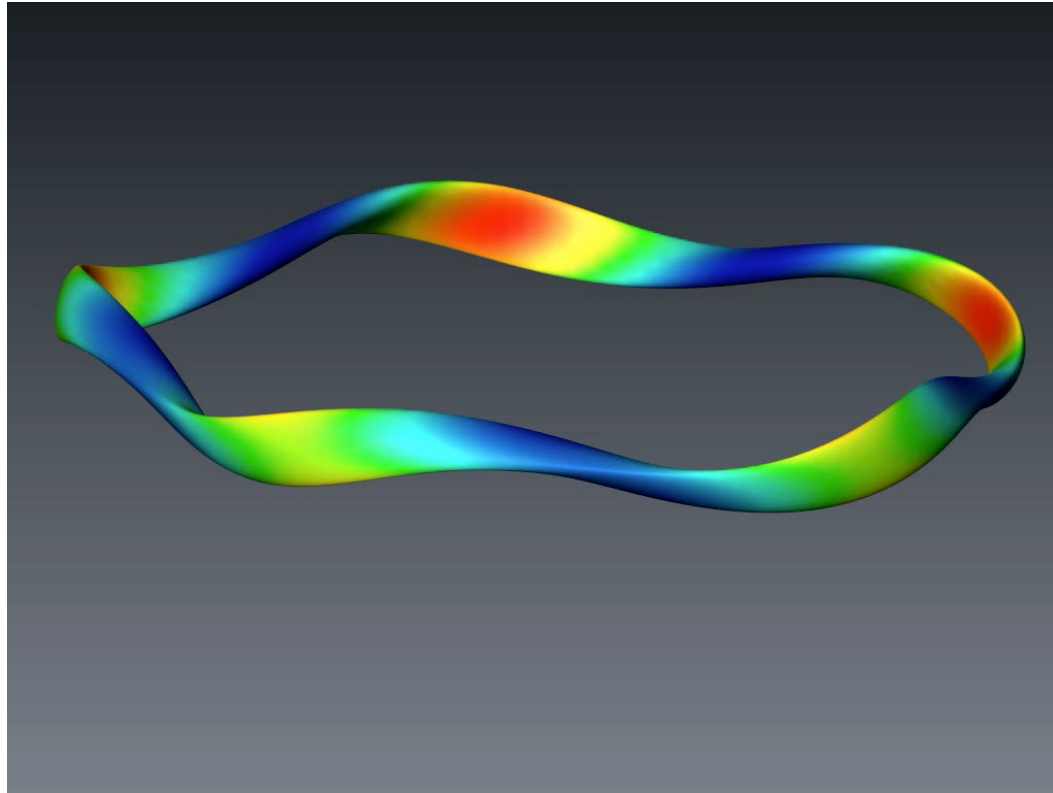
- Field lines on a given flux surface are indistinguishable
- Trapped particles remain confined
- *But:* axisymmetric plasmas...
  - Do not exhibit torsion of the magnetic axis or helical rotation
  - Require a large plasma current to generate the helical field line twist
  - Are not stellarators 🤔



Max Planck Institute for Plasma Physics

# With the right shaping, stellarators can be omnigenous, too!

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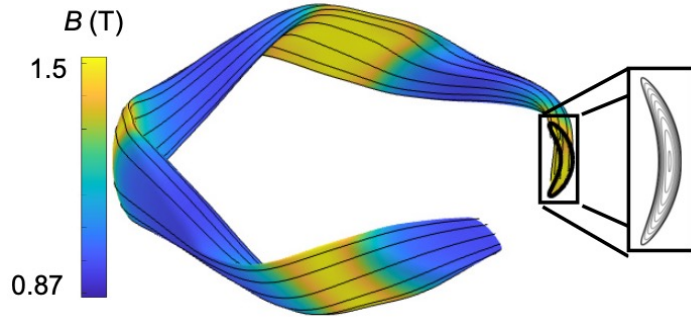


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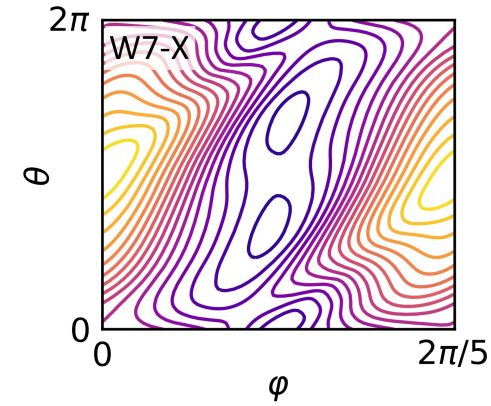
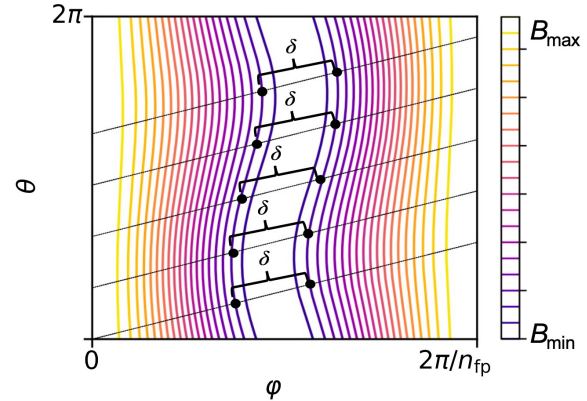
# Popular types of omnigenous stellarators

# Quasi-isodynamic (QI)

- Contours of  $|B|$  close poloidally around flux surfaces



- Example: Wendelstein 7-X
  - Many  $B$  contours not actually closed poloidally
  - Still exhibits substantially lower drift losses than in stellarators with no optimization



A. Goodman et al [9]

[9] A. Goodman et al., *PRX Energy* 3, 023010 (2024)

# Quasisymmetry (QS)

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- Note that integrand for  $\mathcal{J}$  can be expressed as a function of  $B(\ell)$  only
  - $B(\ell)$  = field strength at position  $\ell$  along field line

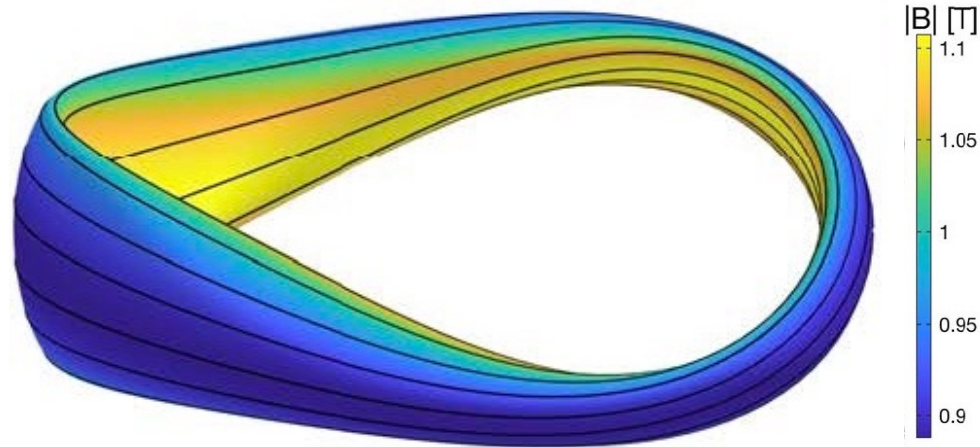
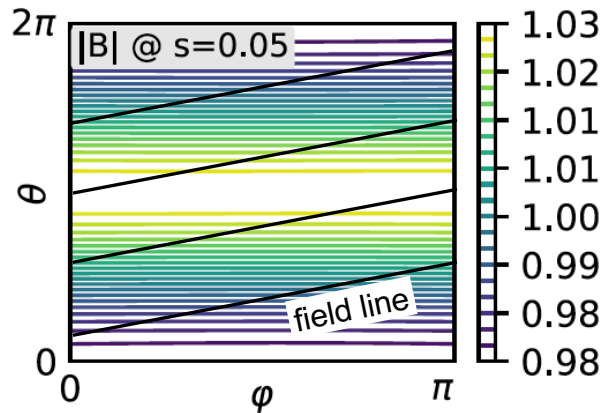
$$\mathcal{J} = \int_{\ell_1}^{\ell_2} m v_{\parallel} d\ell = \sqrt{2\mu m} \int_{\ell_1}^{\ell_2} \sqrt{B_* - B(\ell)} d\ell$$

$\uparrow$  magnetic moment       $\nwarrow$  field strength at bounce point

- Quasisymmetry:  $B(\ell)$  is the same for any field line on a flux surface
  - $\mathcal{J}$  is automatically uniform across the flux surface
    - Satisfies the condition for omigeneity
  - Particles on a given flux surface can't “tell” which field line they are on

# Quasi-axisymmetry

- Quasi-axisymmetry (QA): quasi-symmetry in which  $B$  contours are purely toroidal
- Non-axisymmetric analog to a tokamak

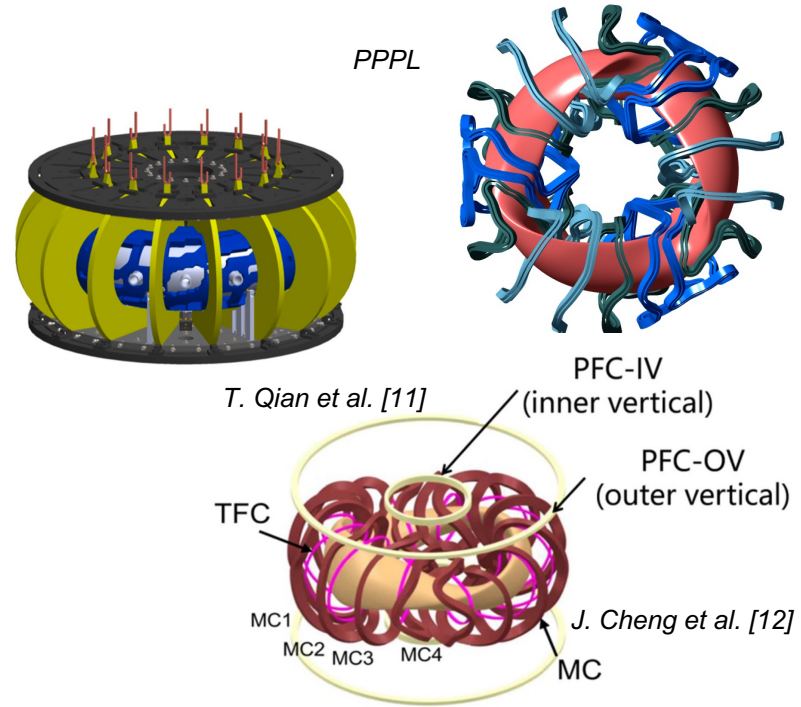


*Equilibrium with near-perfect QA found in 2022 [10]*

[10] M. Landreman and E. Paul, *Phys. Rev. Lett.* **128**, 035001, (2022)

# Quasi-axisymmetry: examples

- NCSX (National Compact Stellarator Experiment) [10]
  - Planned for PPPL, partly constructed
  - Cancelled due to cost and schedule overruns
- MUSE [11]
  - First and only permanent magnet stellarator
  - Tabletop scale with  $B \sim 0.15$  T
- CFQS (Chinese First Quasi-axisymmetric Stellarator) [12]
  - Modular coil device
  - Constructed and currently operating at low field (eventually will run at  $\sim 1$  T)



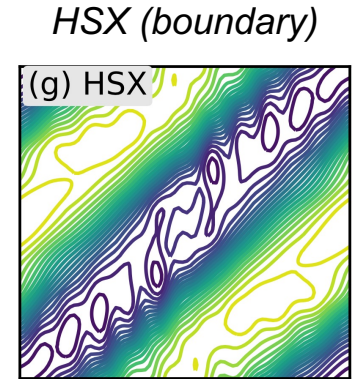
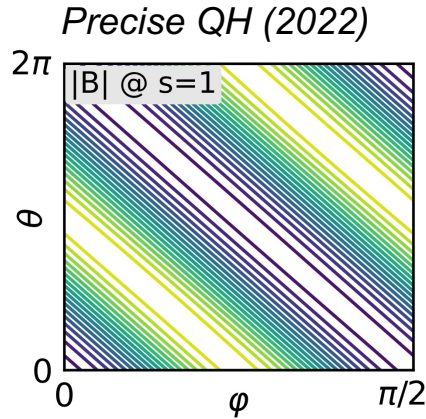
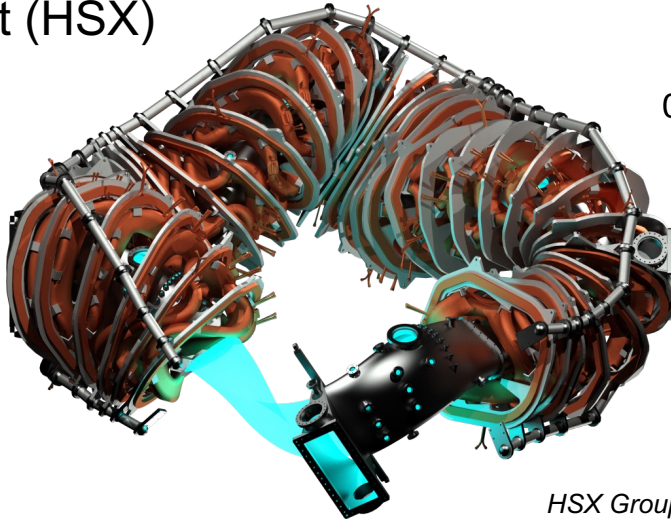
[10] M. Zarnstorff et al., *PPCF* **43**, A237 (2001)

[11] T. Qian et al., *J. Plasma Phys.* **89**, 955890502 (2023)

[12] J. Cheng et al., *PPCF* **67**, 105011 (2025)

# Quasi-helical symmetry

- Quasi-helical symmetry (QH): quasi-symmetry in which  $B$  contours wrap both poloidally and toroidally
- Example: the Helically Symmetric Experiment (HSX)



*M. Landreman and E. Paul, [10]*

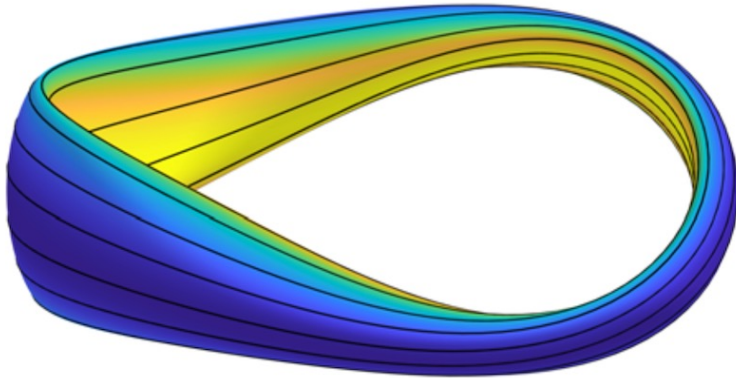
*HSX Group, University of Wisconsin*

# Designing a stellarator

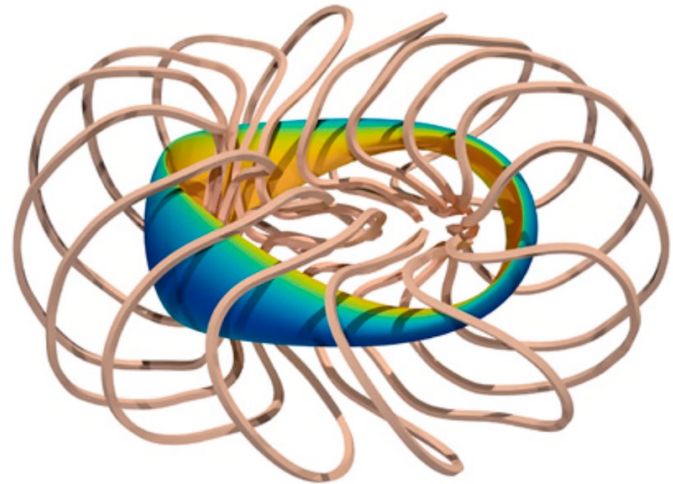
# Stellarator design is often done in two stages

---

1. Design the plasma for good physics properties
2. Design magnets that create the fields necessary to confine the plasma



*Landreman and Paul 2022*

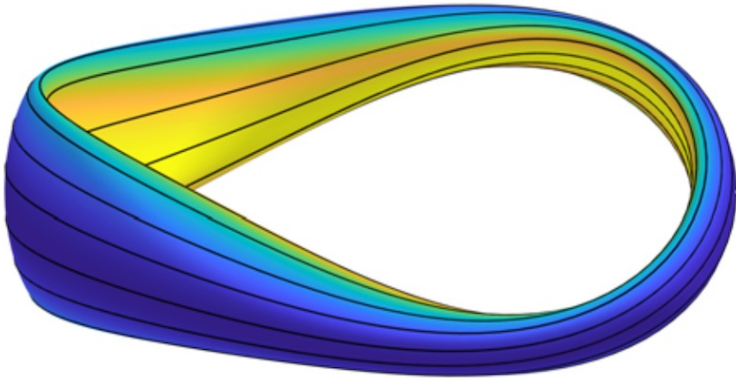


*Wechsung et al. 2022*

# Stellarator design is often done in two stages

---

1. Design the plasma for good physics properties
2. Design magnets that create the fields necessary to confine the plasma



*Landreman and Paul 2022*

# Plasmas can be designed with numerical optimization

---

- Decide on what you want to optimize for
  - Specific rotational transform profile
  - Quasisymmetry, quasi-isodynamicity, etc.
  - Fast particle confinement
  - MHD stability
  - Turbulence minimization
  - ...

# Plasmas can be designed with numerical optimization

---

- Formulate *objective functions* that penalize deviations from the desired attributes
  - $\mathbf{x}_{plasma}$  = parameters that specify the plasma
  - Individual objectives for each attribute:
    - $f_{quasisymmetry}(\mathbf{x}_{plasma})$  = deviation from quasisymmetry for plasma defined by  $\mathbf{x}_{plasma}$
    - $f_{stability}(\mathbf{x}_{plasma})$  = level of fast particle losses for plasma defined by  $\mathbf{x}_{plasma}$
    - $f_{turbulence}(\mathbf{x}_{plasma})$  = level of instability of plasma defined by  $\mathbf{x}_{plasma}$
    - ...
  - Total objective: weighted sum of individual objectives
    - $f(\mathbf{x}) = w_{quasisymmetry} f_{quasisymmetry}(\mathbf{x}_{plasma}) + w_{stability} f_{stability}(\mathbf{x}_{plasma}) + w_{turbulence} f_{turbulence}(\mathbf{x}_{plasma}) + \dots$
    - Weights  $w$  determine which objectives get penalized the most

# Plasmas can be designed with numerical optimization

---

- Use a numerical optimization algorithm to minimize the total objective
  - Supply an initial guess of the plasma ( $\mathbf{x}_{plasma,0}$ )
    - *Cold start*: simple, arbitrary plasma (e.g. an axisymmetric torus)
    - *Warm start*: pre-existing “ok” equilibrium that the optimizer will improve upon
  - Run an optimization code that adjusts  $\mathbf{x}_{plasma}$  to minimize  $f(\mathbf{x}_{plasma})$ 
    - SIMSOPT [6]
    - DESC [13]
    - STELLOPT [14]
    - ROSE [15]
    - ...
  - Output: improved  $\mathbf{x}_{plasma}$  specifying a plasma that better satisfies the objectives

[6] M. Landreman et al., *J. Open Source Soft.* **6**, 3525 (2021)

[13] D. W. Dudt et al., *J. Plasma Phys.* **89**, 955890201 (2023)

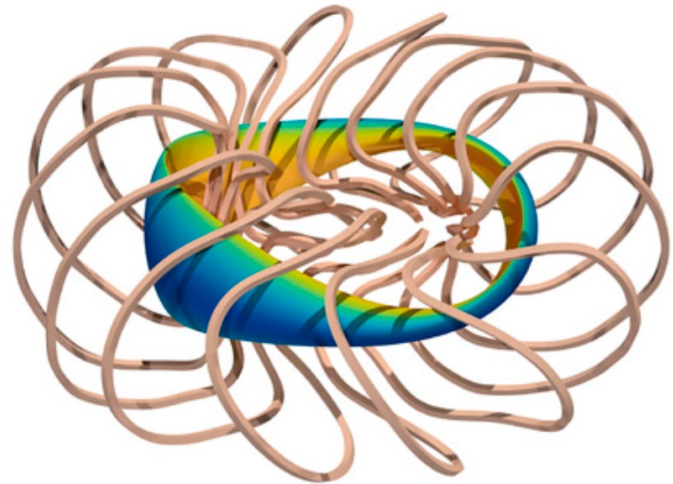
[14] S. Lazerson et al., *STELLOPT*,  
<https://doi.org/10.11578/dc.20180627.6>

[15] M. Drevlak et al., *Nucl. Fusion* **59**, 016010 (2018)

# Stellarator design is often done in two stages

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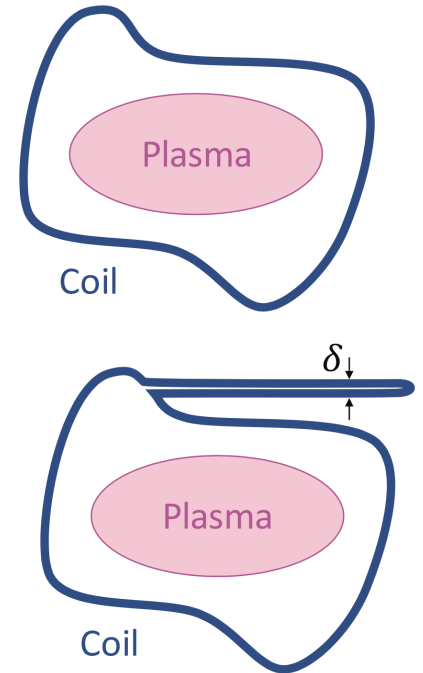
1. Design the plasma for good physics properties
2. Design magnets that create the fields necessary to confine the plasma



*Wechsung et al. 2022*

# Coil optimization allows for many approaches and design schemes

- Goal of coil optimization: find external magnets that produce the desired field within the plasma volume
- This is an *ill-posed problem* [16]
  - Example: adding an out-and-back current path to a coil will result in a different magnet shape but the same field in the plasma
- This can actually be a good thing
  - No one right way to design magnets
  - Lots of flexibility in how to define and parametrize the solution space
- On the other hand...
  - Difficult to know for sure if a given solution is optimal or if a better one exists

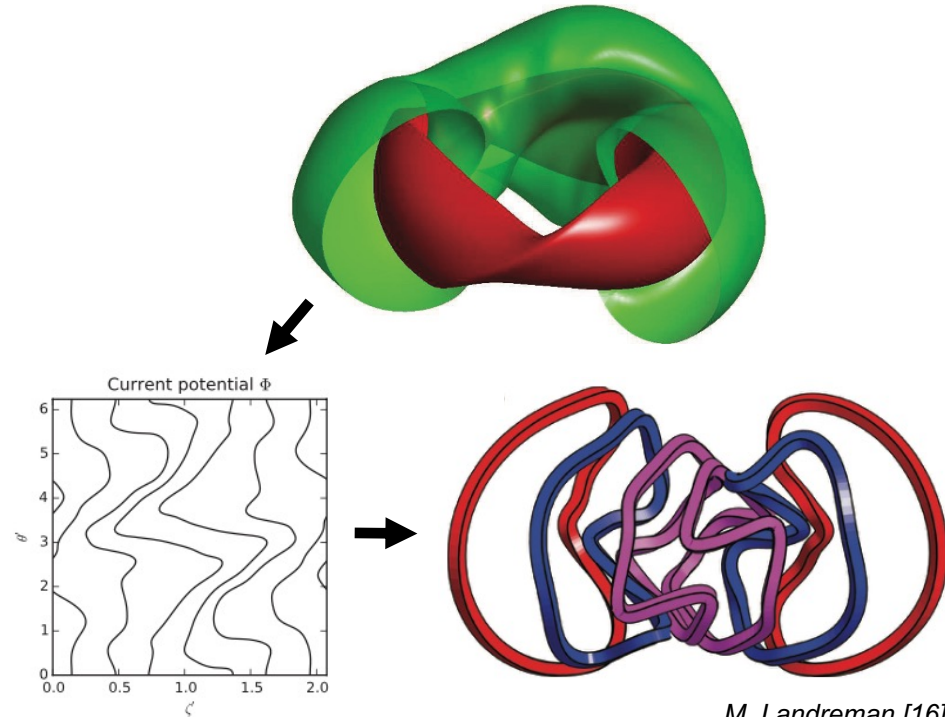


M. Landreman [5]

[16] M. Landreman, *Nucl. Fusion* **57**, 046003 (2017)

# Winding surface approach / Merkel's method [17]

- Define a toroidal *winding surface* enclosing the plasma
- Find an optimal current distribution on the winding surface that creates the necessary field on the plasma
- Discretize the current distribution to determine coil shapes
  - Calculate streamlines of surface current
  - Define elementary coils along those streamlines

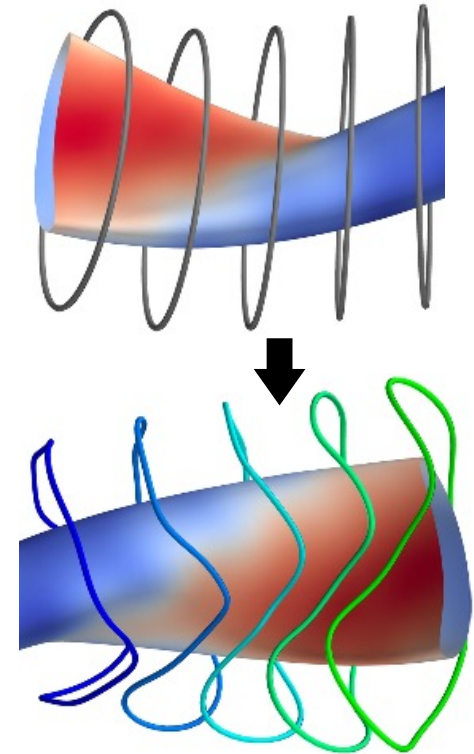


M. Landreman [16]

[17] P. Merkel, *J. Comp. Phys.* **66**, 83 (1986)

# Space curve optimization

- Initialize a specific number of coils around the plasma
  - Define geometric parameters  $\mathbf{x}_{coils}$  specifying the shape of each coil
  - Simple initial guesses (i.e. planar circles) are typically fine
- Minimize an objective function  $f(\mathbf{x}_{coils})$  that penalizes field error and engineering properties
  - $f(\mathbf{x}_{coils}) = f_{accuracy}(\mathbf{x}_{coils}) + f_{curvature}(\mathbf{x}_{coils}) + f_{separation}(\mathbf{x}_{coils}) + \dots$
- Optimized shapes ( $\mathbf{x}_{coils}$ ) form the basis for the design of the coils

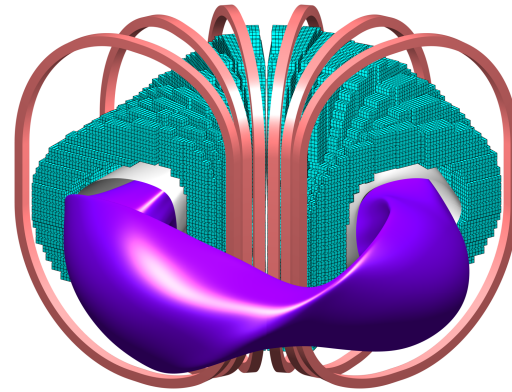


C. Zhu et al. [18]

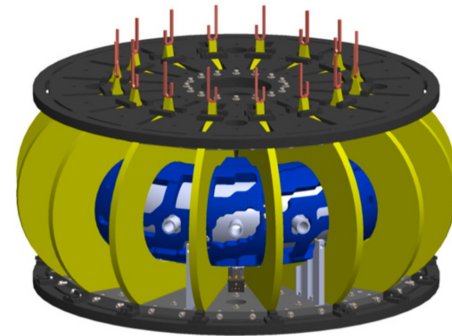
[18] C. Zhu et al., *Nucl. Fusion* **58**, 016008 (2018)

# Permanent magnet optimization

- Permanent magnets can be used to simplify coils
  - Cannot confine a plasma alone – some coils are needed to generate a toroidal field
  - Permanent magnets can generate 3D shaping so that coils can be less complicated
- Optimize dipole moments of magnets at selected positions around plasma



*K. C. Hammond et al [19]*



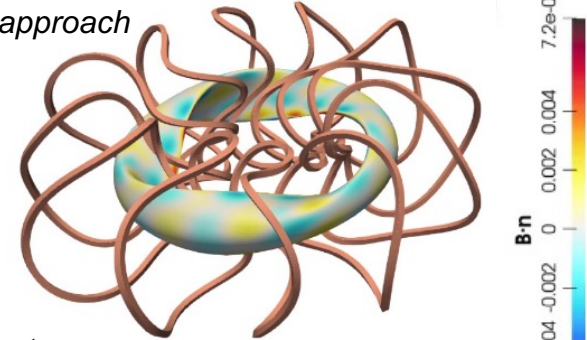
*T. Qian et al. [11]*

- [11] T. Qian et al., *J. Plasma Phys.* **89**, 955890502 (2023)  
[19] K. C. Hammond et al., *Nucl. Fusion* **62**, 126065 (2022)

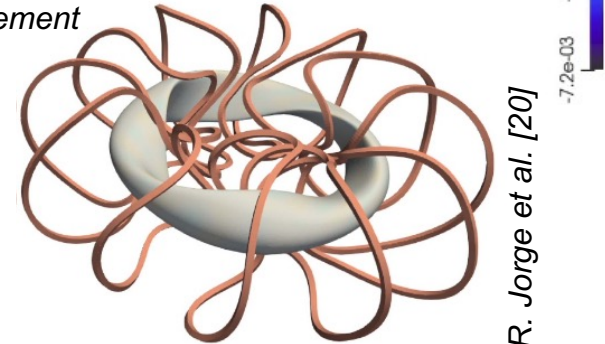
# Single stage optimization

- Optimize plasma and coils at the same time
  - Minimize a combined objective:  $f(\mathbf{x}_{plasma}, \mathbf{x}_{coils})$
- Advantage: steers solution toward plasmas that are feasible to build coils for
- Disadvantage: more computationally expensive than separate stage-1 and stage-2 optimizations
- Tends to be more successful when done from a warm start with plasmas and coils that are already “pretty good”

Initialization from two-stage approach



Single-stage refinement



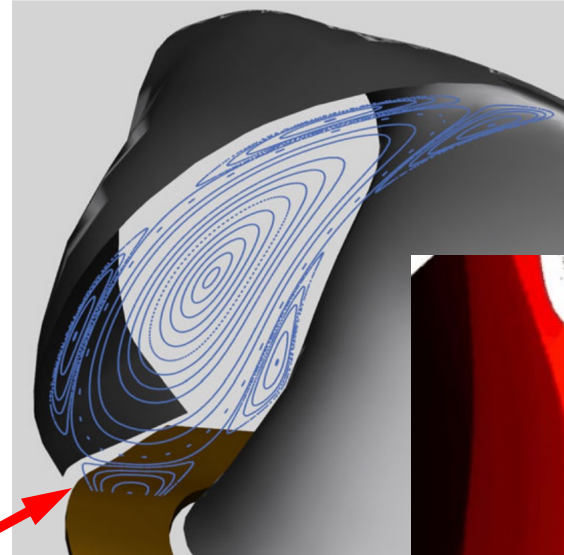
R. Jorge et al. [20]

[20] R. Jorge et al., *Plasma Phys. Control. Fusion* 65, 074003 (2023)

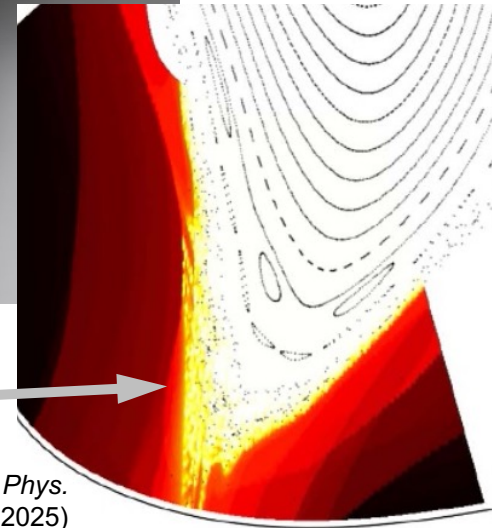
# Some current research areas

# Handling heat and particle exhaust

- Fusion devices need to exhaust He ash efficiently without overheating
- Divertor:
  - Receives majority of plasma flux
  - Directs exhaust gas to pumps
- Stellarator divertors deal with (or take advantage of) complexity of 3D magnetic fields
  - *Island divertors*: “magnetic islands” isolate plasma from targets (used in W7-X)
  - *Non-resonant divertors*: chaotic field lines isolate plasma from targets (not extensively tested)



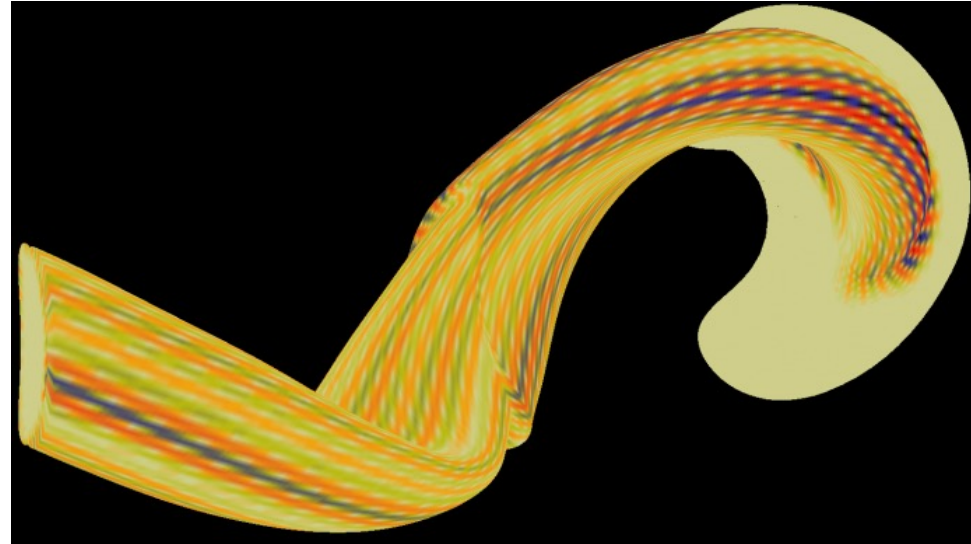
[21] A. Bader et al.,  
*J. Plasma Phys.* **91**,  
E67 (2025)



[22] K. Garcia et al., *Plasma Phys.*  
*Control. Fusion* **67**, 035011 (2025)

# Reducing turbulent transport

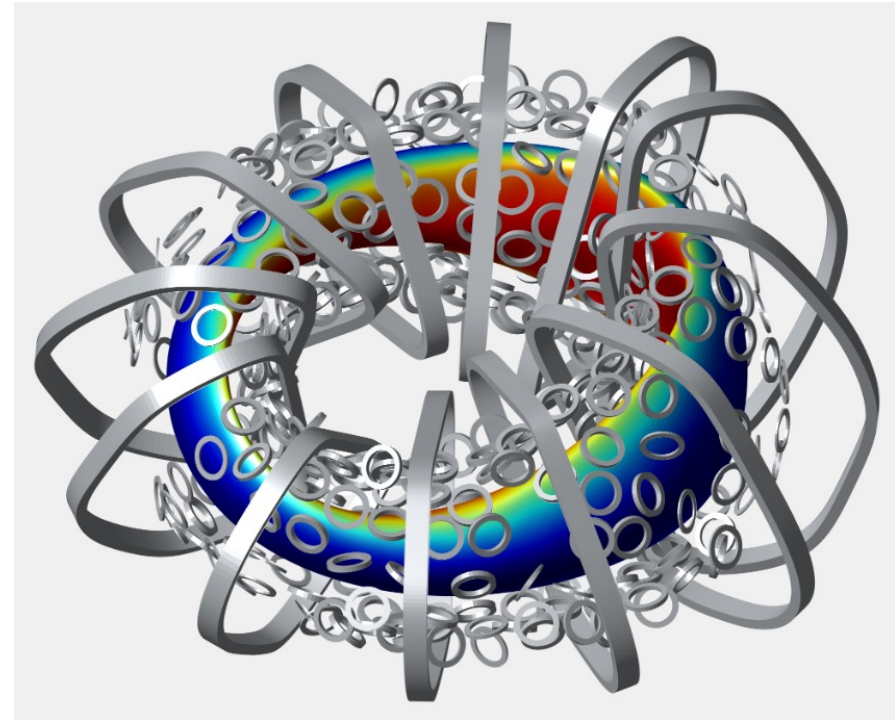
- Turbulence: small-scale fluctuations that can push heat and particles out of the plasma
- Strategies for mitigating turbulent transport
  - Optimize the magnetic geometry to reduce susceptibility to micro-instabilities that cause turbulence
  - Optimize plasma heating and fueling to achieve density and temperature gradients that stabilize micro-instabilities



I. Calvo et al., <https://sites.fusion.ciemat.es/multitransell/>

# Simplifying the coils

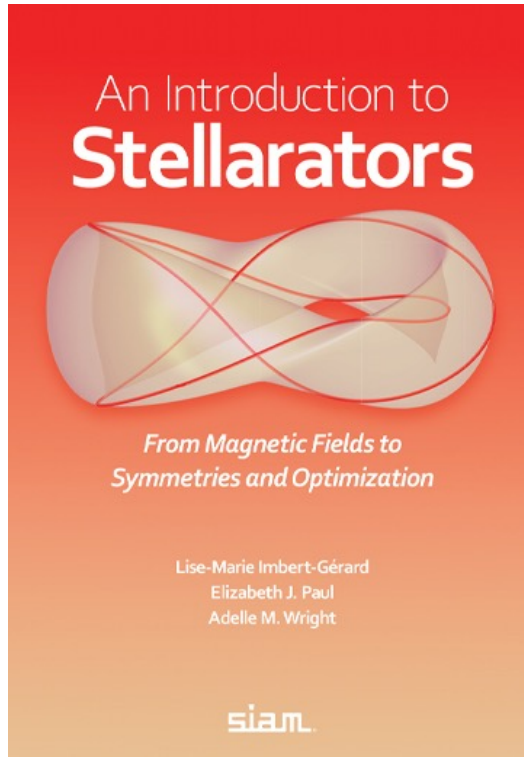
- Coils are a major cost driver of modern stellarators
  - Complex shapes
  - Tight engineering tolerances ( $\sim 10^{-3}$ - $10^{-4}$ )
- Ideas for simplifying coils in the future
  - Stochastic optimization to reduce sensitivity to misalignments
  - Designs with large arrays of simple coils (e.g. Thea Energy)



[27] C. P. S. Swanson et al., <https://arxiv.org/abs/2512.08027>

# Recommended reading

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## An Introduction to Stellarators

From Magnetic Fields to  
Symmetries and Optimization

Lise-Marie Imbert Gérard, Elizabeth J. Paul,  
and Adelle M. Wright

<https://doi.org/10.1137/1.9781611978223>

ArXiv preprint: <https://doi.org/10.48550/arXiv.1908.05360>

# Stellarators

**Kenneth C. Hammond**

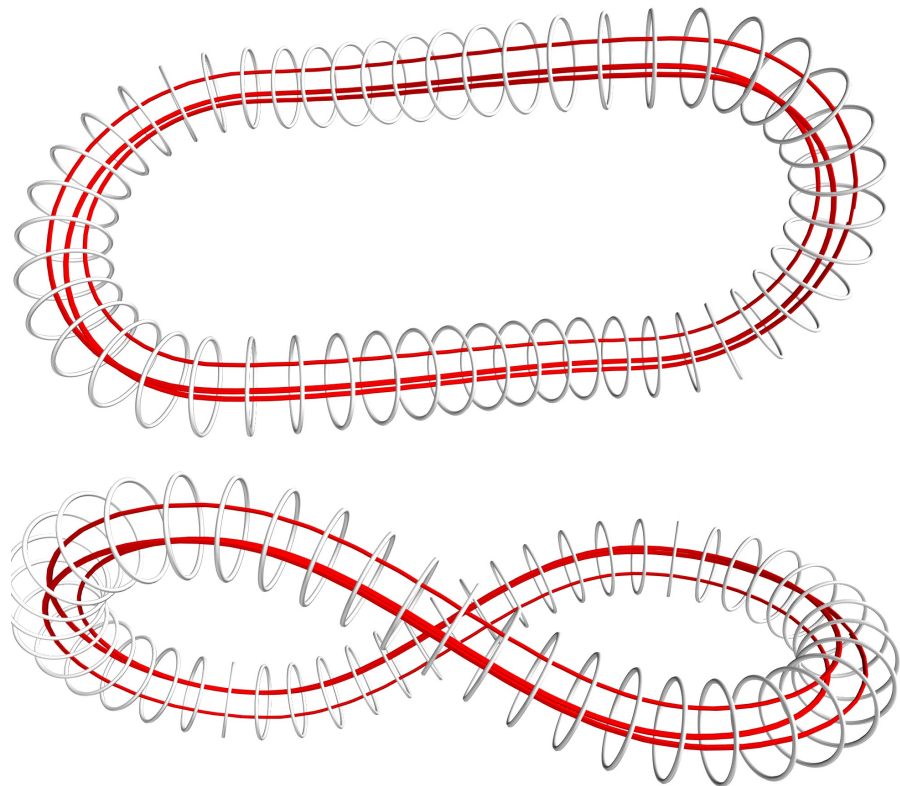
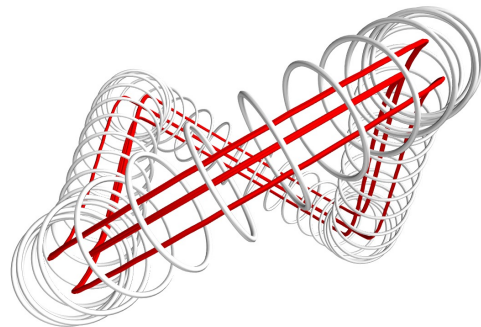
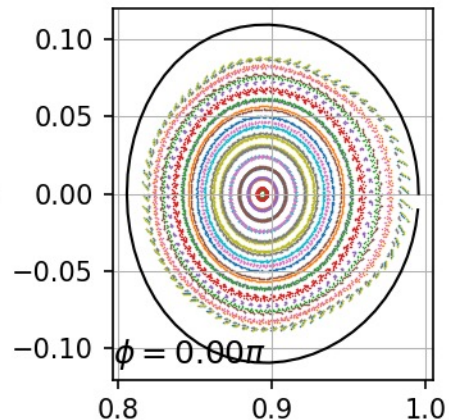
*Research Physicist*

*Princeton Plasma Physics Laboratory*

Back-up slides

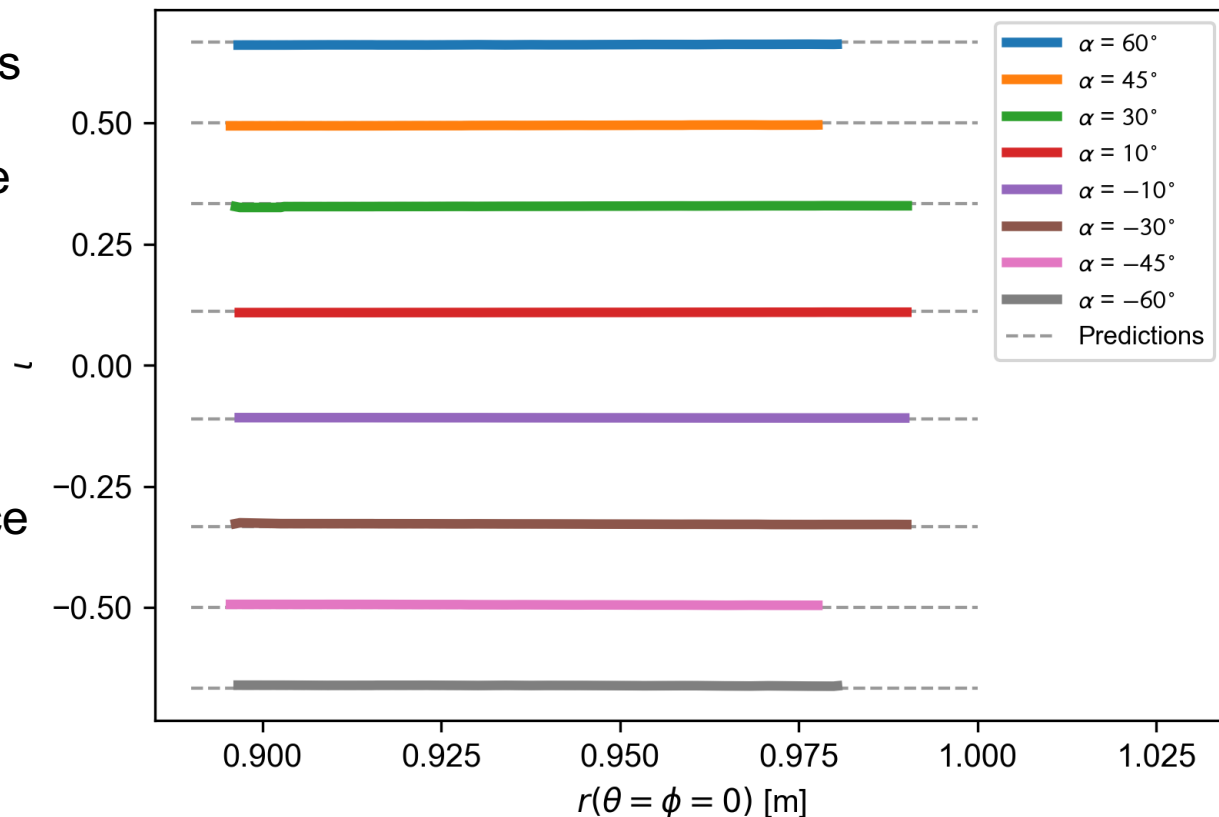
# Numerical model of figure-8 stellarator agrees very well with simple geometric predictions

- $\alpha = \pi/6 \rightarrow$  prediction:  $\iota = 1/3$ 
  - Field lines “bite their tails” after three transits

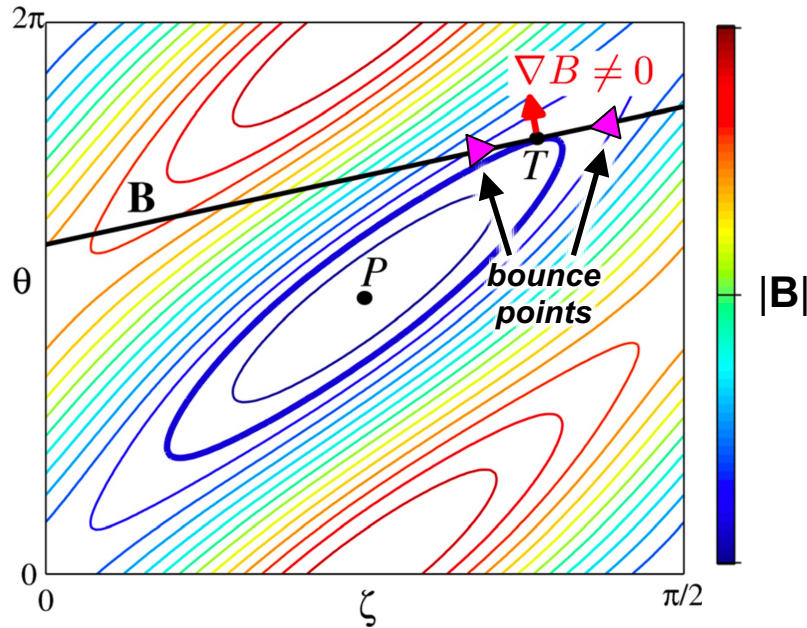


# Numerical model of figure-8 stellarator agrees very well with simple geometric predictions

- Rotational transform profiles agree to within 3% of predictions for a wide range of tilt angles
  - *Profile*: dependence of quantity on minor radius (distance from core/axis)
- Flatness of profiles (zero *shear*) affirms independence of  $\iota$  on minor radius



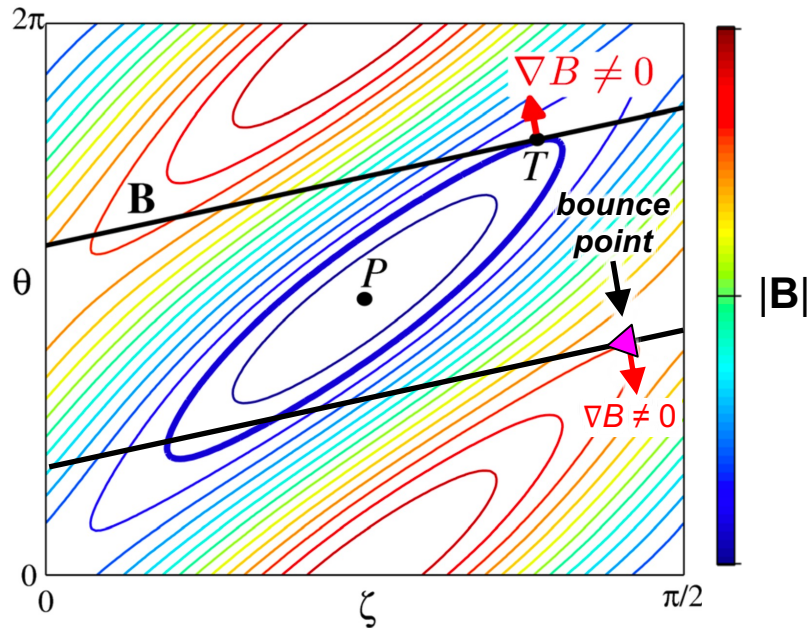
# Key property of omnigenous stellarators: no local point minima of $|\mathbf{B}|$ on flux surfaces [28]



- Some contours of  $B = |\mathbf{B}|$  will be tangent to field lines
- If a particle is deeply trapped in the vicinity of the tangency point:
  - $\mathbf{B} \times \nabla B$  has the same sign on both sides of the minimum
  - Radial drift velocity is in the same direction for entire bounce trajectory
  - Particle exhibits nonzero mean radial drift: violation of omnigenity

[28] M. Landreman and P. Catto, *Phys. Plasmas* **19**, 056103 (2012)

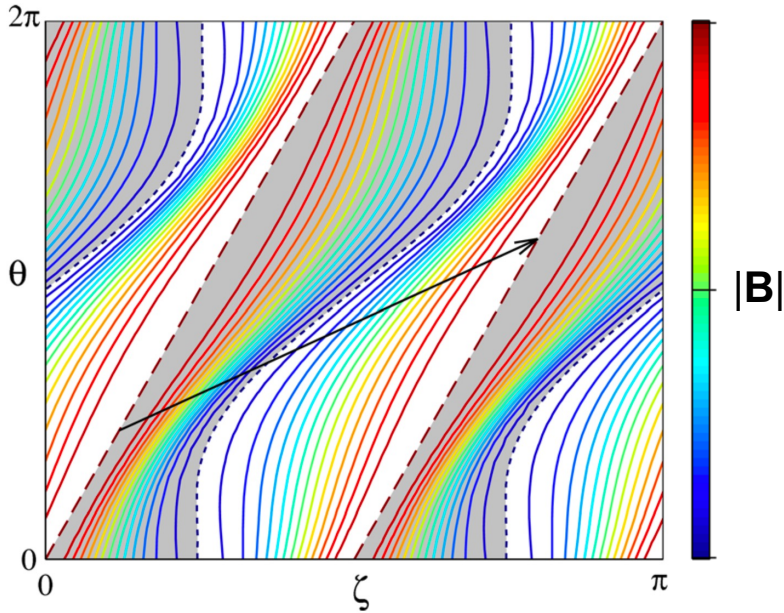
# Key property of omnigenous stellarators: no local point maxima of $|B|$ on flux surfaces



- Drift velocity is nonzero at the tangency point
- Barely trapped particles stay near the bounce point for a long time
  - May be long enough to drift out of the plasma

[28] M. Landreman and P. Catto, *Phys. Plasmas* **19**, 056103 (2012)

# Therefore: for omnigenity, all $|B|$ contours must wrap toroidally and/or poloidally



- Drift velocity is in the opposite direction on each side of the minimum  $B$
- Drift velocity is zero at local maxima in  $B$  along field lines  $\rightarrow$  weakly trapped particles do not escape
- Additional properties
  - Distances along field lines between points of equal  $B$  are the same across the flux surface
  - Contours of the maximum  $B$  are straight in Boozer coordinates

[28] M. Landreman and P. Catto, *Phys. Plasmas* **19**, 056103 (2012)