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June 12, 2025 | Fusion Energy Division

Integrated Plasma Modeling:

*From the Power of the Sun to
Powering Society on Earth*

PRESENTED BY

Sebastian De Pascuale

PPPL SULI Introduction to Fusion Energy and
Plasma Physics Summer School Course



U.S. DEPARTMENT OF
ENERGY

ORNL IS MANAGED BY UT-BATTELLE LLC
FOR THE US DEPARTMENT OF ENERGY

About Me – From Astronomy to Space Physics



Grinnell College



University of Iowa

About Me – From Water Rocketry to Rowing on Water



Knox County, Ohio



Knox County, Tennessee

My Trajectory into the Field of Fusion Energy Science



2011 – Los Angeles
Research Experiences
for Undergraduates
(REU)



2015 - Los Alamos
Space Weather
Summer School
(SWSS)



2018 - Oak Ridge
Post-Doctoral
Research Associate
(ORNL)

“Just as the system of the sun, planets and comets is put in motion by the forces of gravity, and its parts persist in their motions, so the smaller systems of bodies also seem to be set in motion by other forces and their particles to be variously moved in relation to each other and, especially, by the electric force”

Sir Isaac Newton, *Principia* (1687)

Exploring Dusty Horizons at University of California, Los Angeles

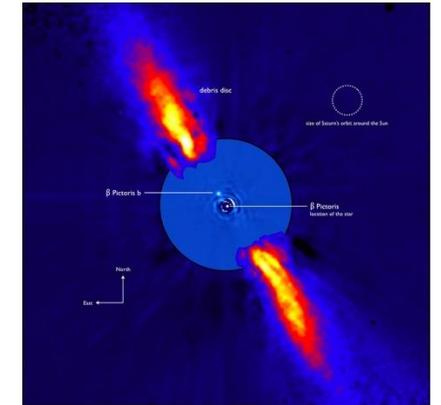
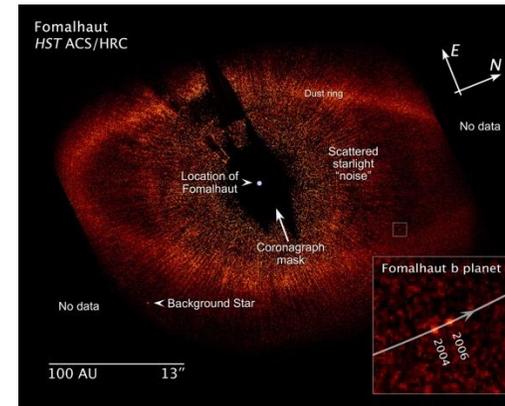


“Your reward will be the widening of the horizon as you climb”
- Cecilia Payne (figure from NASA’s Mars Perseverance Rover)

*Undergraduate Mentor Dr. Michael Fitzgerald

How do you find exoplanets orbiting other stars?

- Observe light curves from the occlusion of a planetary body in front of a star, or observe the gravitational disturbance of smaller bodies to the surrounding circumstellar debris disk



- Dust grain sized particles scatter light preferentially in the forward direction under Mie theory, causing the zodiacal light observed on Earth or the reddish haze observed on Mars



<https://usfusionenergy.org/opportunities>



Illustration by Thumý Phan
for U.S. Fusion Outreach

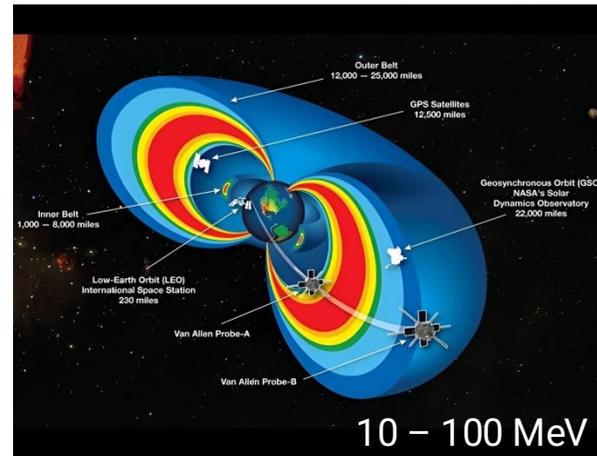
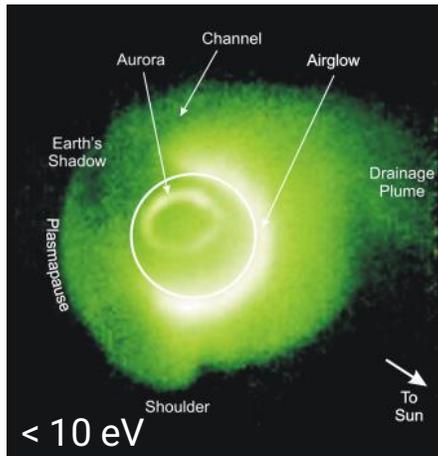
Cecilia Payne

What are stars made of? How do stars make their energy? We know the answers to these questions thanks to Dr. Cecilia Payne-Gaposchkin, who discovered that the stars, like our Sun, are mostly made of hydrogen and helium. This led to our understanding of fusion—the reaction that powers our sun and the stars. These elements are the “fuel” that powers fusion reactions inside the sun in a giant fusion engine. Together, we are working to build fusion engines here on Earth. Join the celebration and join the movement to bring the power of the stars to Earth.

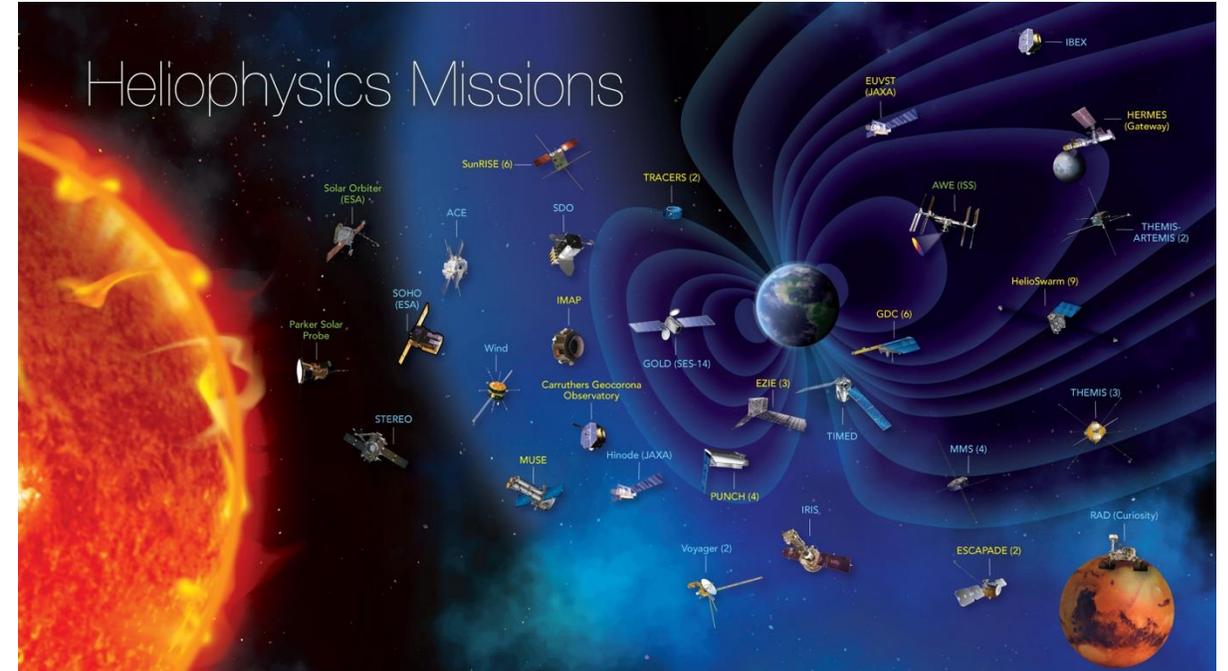
Pursuing Complex Science at Los Alamos National Laboratory

How does space weather impact life on Earth?

- Geomagnetic storms are caused by the plasma connection of Earth's magnetosphere with the enveloping solar heliosphere drawn out by coronal mass ejections and solar wind and flares



- Radiation belts trap energetic charged particles and can pose operational risks to commercial, national, and scientific satellites



"I think the next [21st] century will be the century of complexity"
-Stephen Hawking (figure from NASA GSFC Fleet ID:14718)

*Thesis Advisor Dr. Craig Kletzing and Vania Jordanova

□ De Pascuale et al., JGR Space Phys., 2018
<https://doi.org/10.1029/2018JA025776>

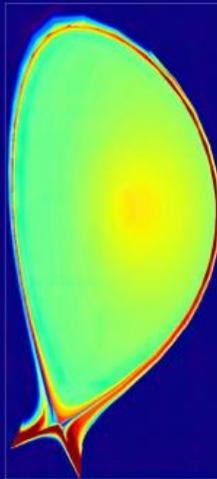
14TH ITER INTERNATIONAL SCHOOL



30 June (Monday) – 4 July (Friday) 2025
Aix-Marseille University, Aix-en-Provence, France

INTEGRATED MODELLING IN MAGNETIC FUSION PLASMAS

The subject of the 2025 school is "Integrated Modelling in Magnetic Fusion Plasmas." Reliable predictions of ITER plasmas—spanning the entire cross-section from the plasma core and scrape-off layer (SOL) to the material surface—are critical for achieving ITER's fusion power demonstration goals. Such predictions are indispensable for defining and preparing plasma operational scenarios, analyzing the plasma pulses planned for ITER, and evaluating required control schemes via simulations of measurements, actuators, and plasma responses.



Hosted by Aix-Marseille University, France
<http://www.univ-amu.fr/>
<https://iis2025.sciencesconf.org/>



Summer Schools

ITER International School is organized nearly every year to offer early career scientists and engineers multi-disciplinary networking and instruction on the exciting field of nuclear fusion.

Each school is organized around a specific scientific theme and hosted by different countries around the world at institutions with access to tours of experimental research facilities.

Previous topics include:

- Diagnostics and Data Science
- Impacts of Energetic Particles
- Operation Scenarios and Control
- Technology of Power Flux Handling
- Transport and Pedestal Physics
- Disruptions and Magneto hydrodynamics
- High Performance Computing
- Plasma-Surface Interactions



Solving Big Problems at Oak Ridge National Laboratory



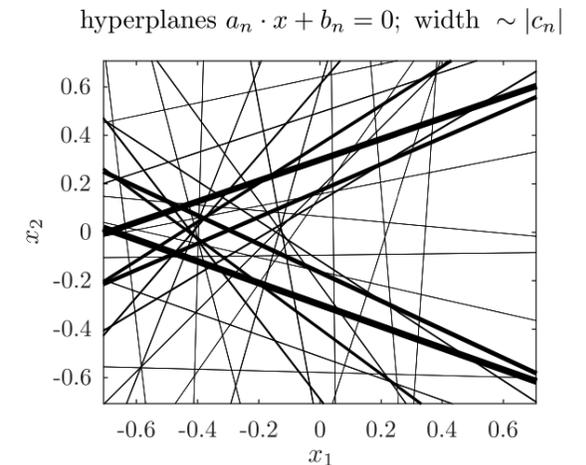
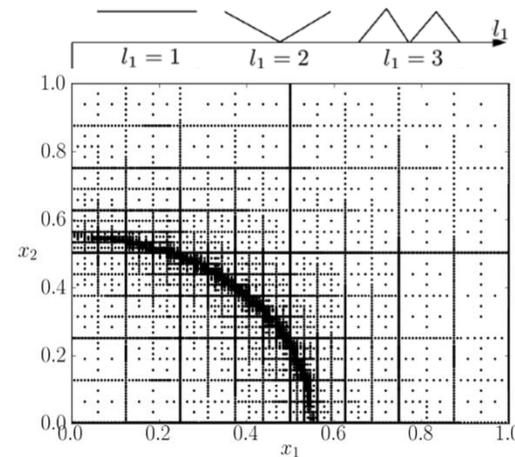
“We need people [at Oak Ridge] to solve big problems”
- Eugene Wigner (figure from ORNL Main Campus)

*Post-doc Advisor Dr. David Green

□ De Pascuale et al., J. Comp. Phys., 2023
<https://doi.org/10.1016/j.jcp.2023.112089>

How can we accelerate progress towards fusion?

- High fidelity fusion simulations are computationally prohibitive without accounting for numerical limitations on the treatment of multiple scales



- Sparse discretization methods in space and time offer tractable approaches to reduce computational overhead and develop reduced models of governing physical systems

<https://arxiv.org/pdf/1110.0010>

<https://arxiv.org/pdf/2410.02132>

Research Internships

ORNL educational programs provide experiences complementing academic programs and enabling opportunities for additional learning, development, and training in a laboratory workplace setting.

Commitment to enriching scientific culture and the future workforce through **learning objectives** based, **mission aligned** experiences with world-class researchers, staff, facilities.

Over **20 program offerings** are available at any point in your developing career including elementary education, high school, undergraduate, community college, and graduate student participants.

See you at the traveling science fair or during a fall, spring, or summer internship!



<https://education.ornl.gov>



Three Takeway Themes

QUANTUM
Empowering scientists to pursue quantum innovation >



FUSION
Fusion energy project among six ORNL Federal Laboratory Consortium wins >



AI
Leading research to ensure secure, trustworthy, and energy efficient AI >





First in Flight

“The period from 1885 to 1900 was one of unexampled activity in aeronautics, and for a time there was high hope that the age of flying was at hand...

But the public, discouraged by the failures and tragedies it had just witnessed, considered flight beyond the reach of humans, and classed its adherents with the inventors of perpetual motion...

Heavier-than-air flight had been **around the corner** for decades, with little to show for it...”

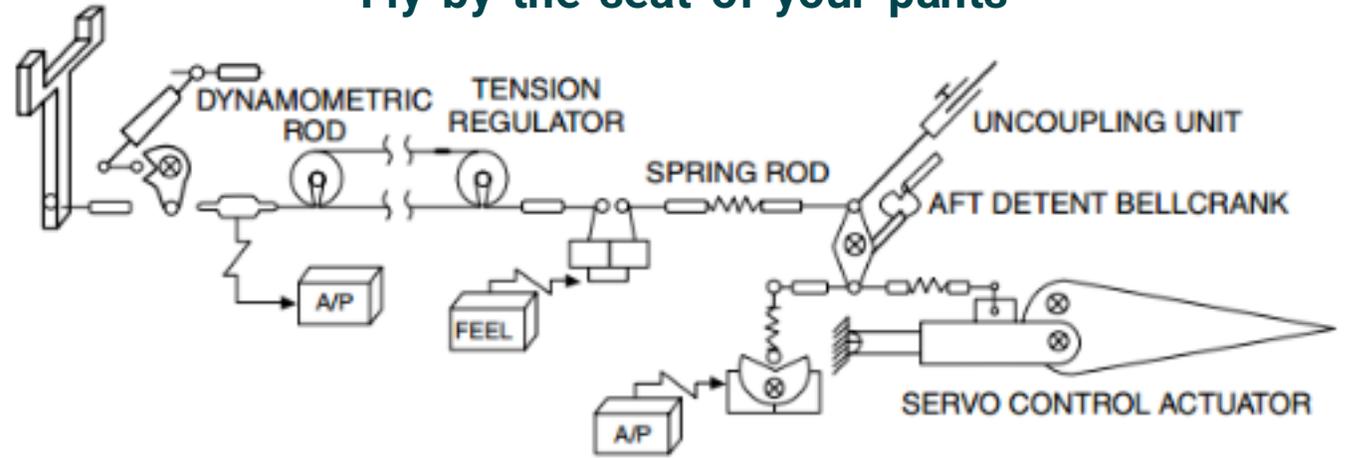


Controlled, powered flight

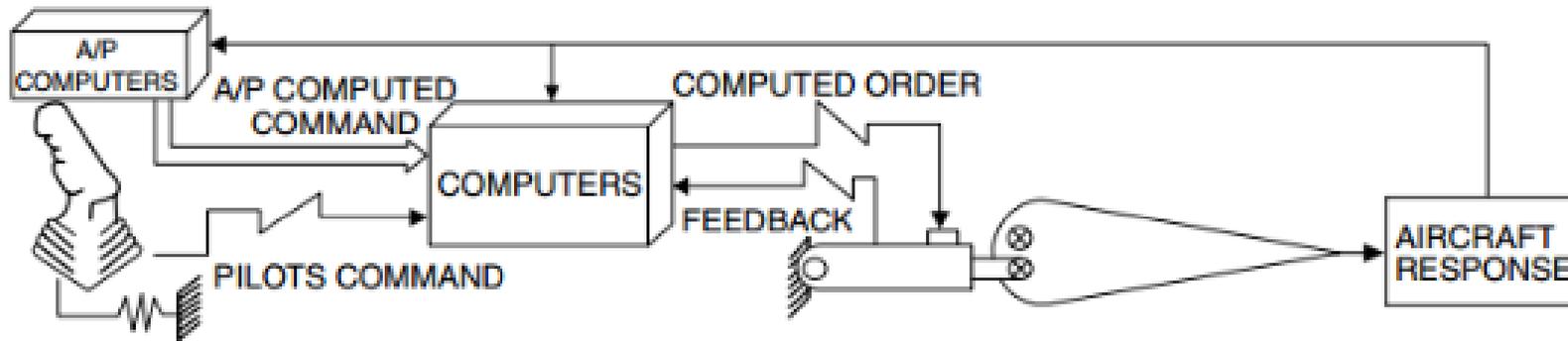
Flying, with Contemporary Style



Fly-by-the-seat-of-your-pants



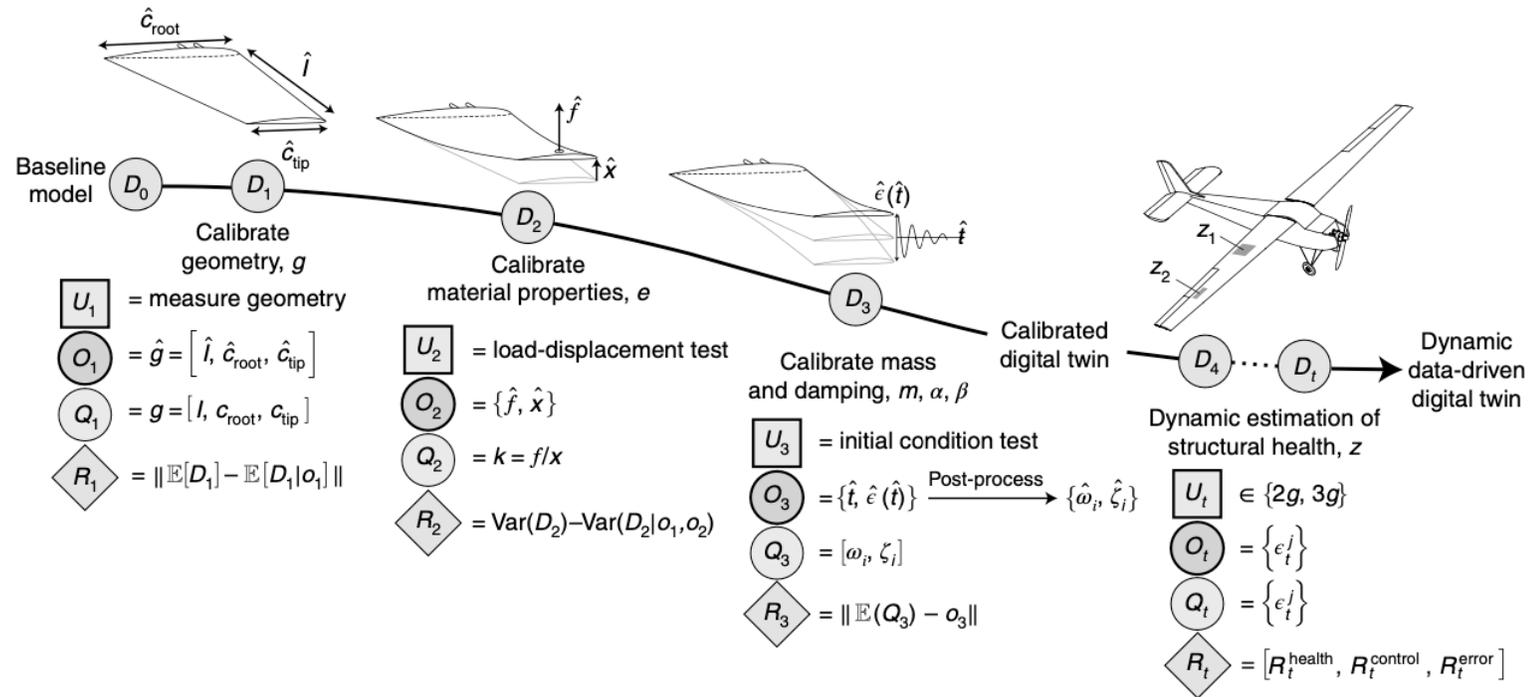
Fly-by-wire



Flying, with Common Sense



Fly-by-autonomous-self-organization



From Flight Simulator to Pulse Simulator

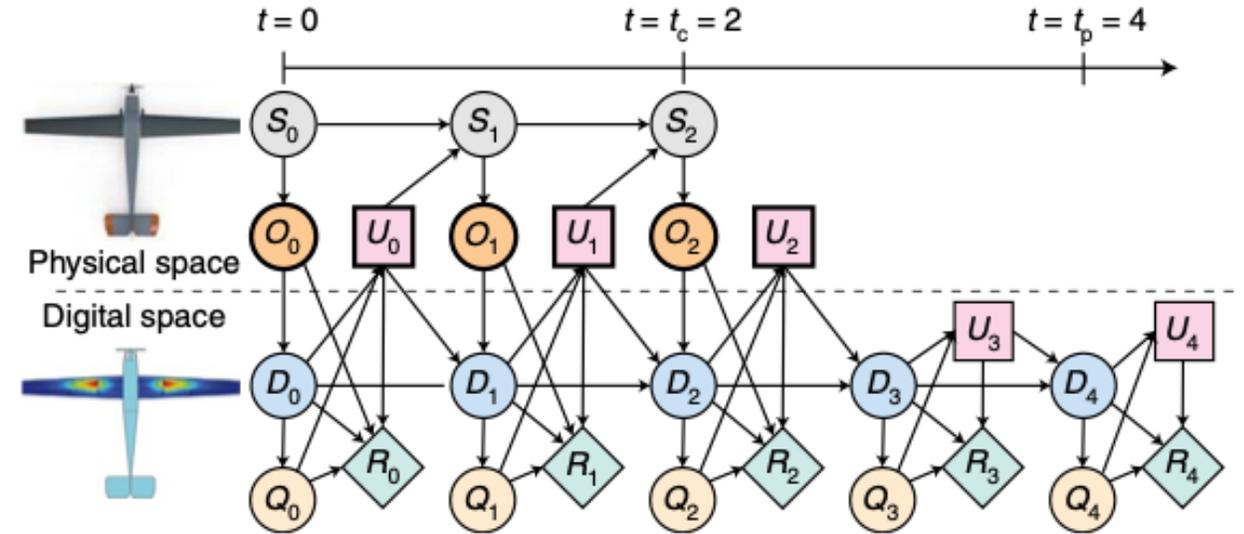
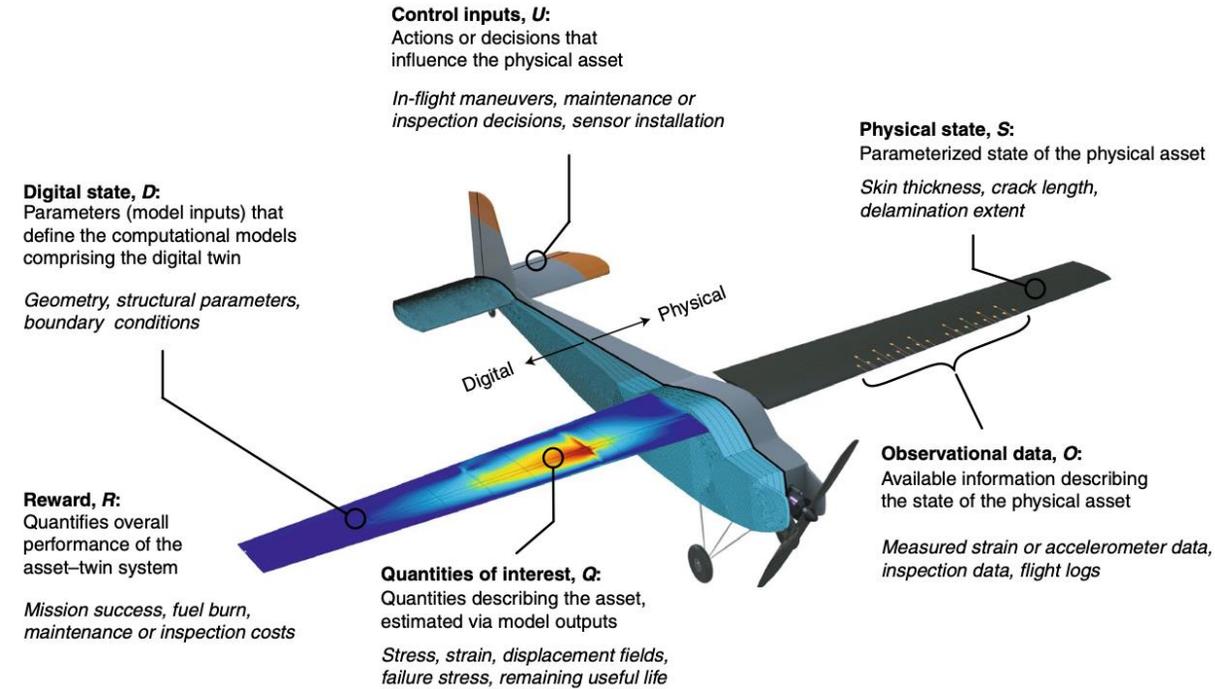


Modern Aircraft



Advanced Tokamak

Designing a Digital Twin



probabilistic graphical models with data assimilation

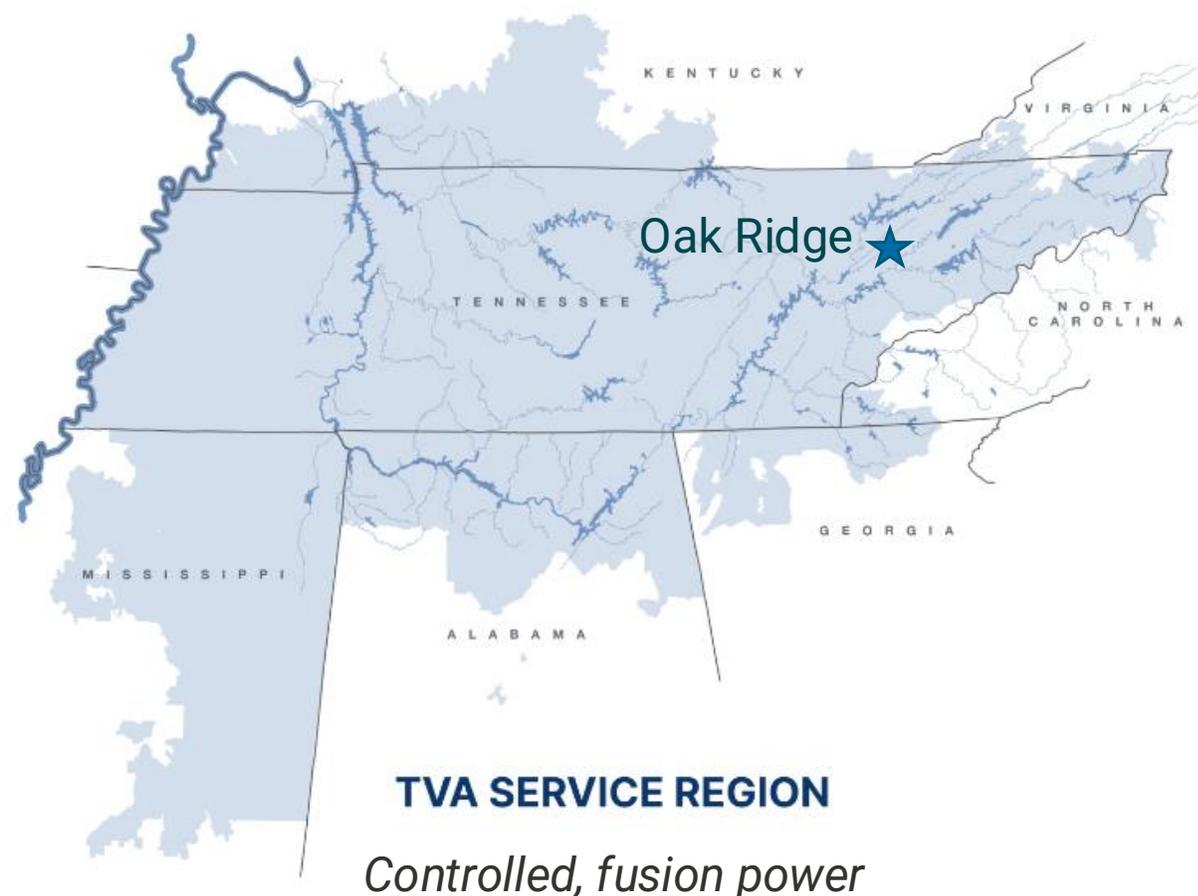
First in Fusion – Putting Another Star on the Map

“Fusion is unquestionably one of the key technologies that will shape the coming millennium” following energy crisis of the 70s

In the 1980’s Governor Alexander had a vision to create an Oak Ridge Corridor, and by 1995 an annual Science & Technology Summit was convened

During this time, fusion scaled back in enthusiasm, but reached new world records as well as substantial technological challenges

In 2010 “**Thirty years later**, tomorrow is here and –despite some delay– so is the bright promise” on ITER



Chesterfield

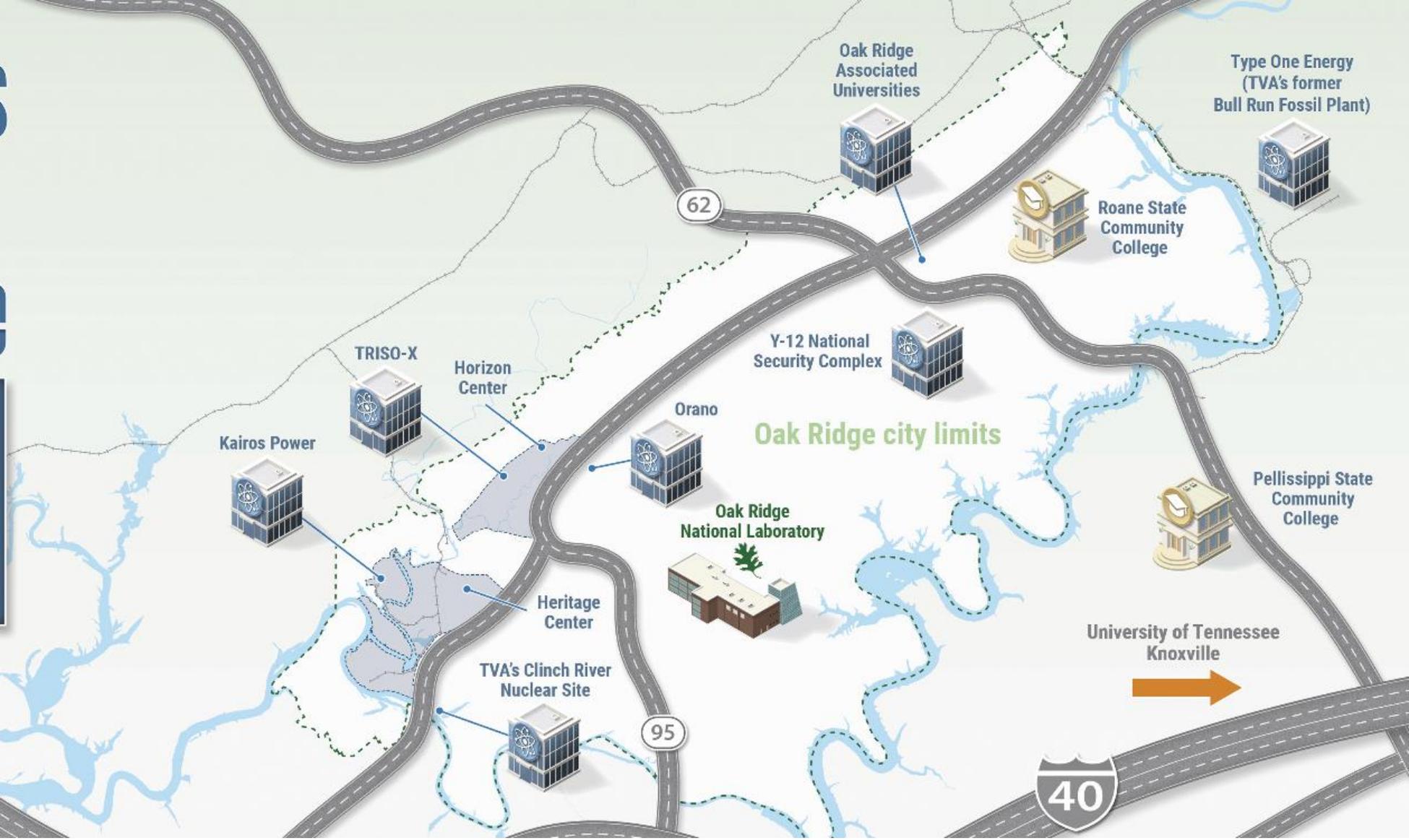


Kitty Hawk



Tennessee's nuclear renaissance

East Tennessee's nuclear industry is growing rapidly. Businesses are introducing new reactor and fuel designs. Institutions such as ORNL, the Tennessee Valley Authority and the University of Tennessee are providing knowledge and support. And local colleges are ensuring this growing industry has a skilled workforce. Here are just a few of the region's major players.



FUTURE FUSION ENERGY HUB

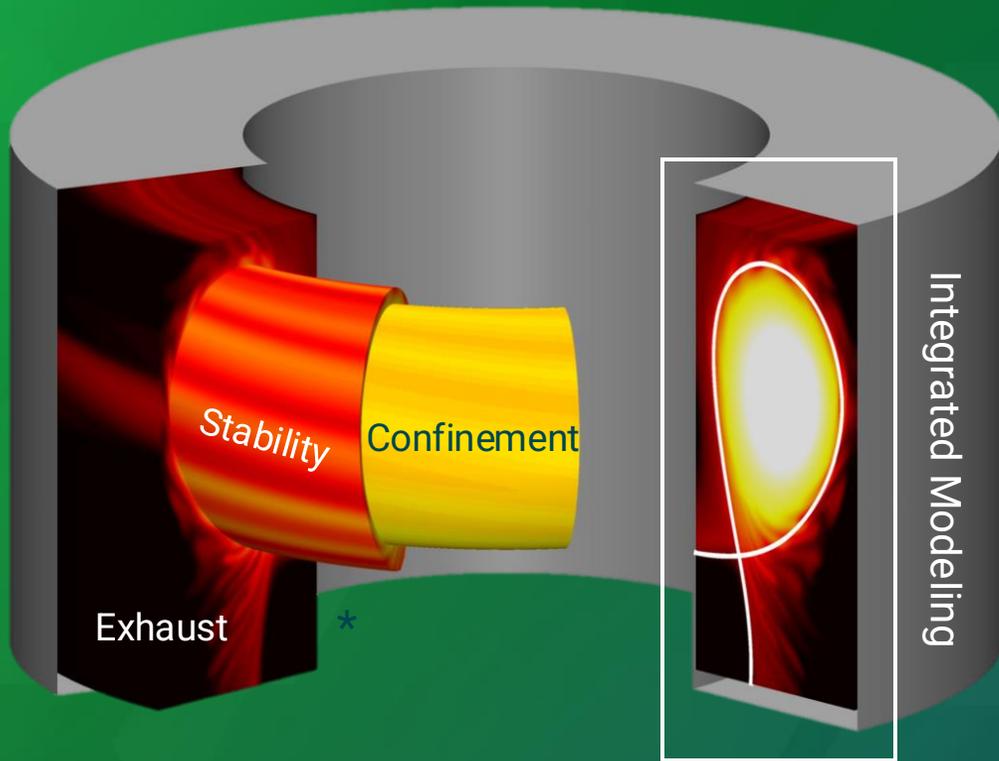
<https://www.ornl.gov/nuclearishere>

Building Bridges for Public-Private Partnerships to Enable a Thriving Fusion Workforce and Industry Ecosystem



<https://www.energy.gov/fusion-energy>

Transport Processes are Inherently Multi-Scale and Ubiquitous Across the Entire Fusion Plasma System



Core Confinement

- Maintain plasma performance under turbulence and other loss mechanisms

Edge Stability

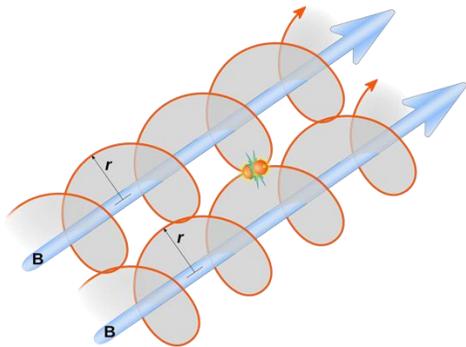
- Manage high density and temperature against magnetohydrodynamic effects

Power & Particle Exhaust

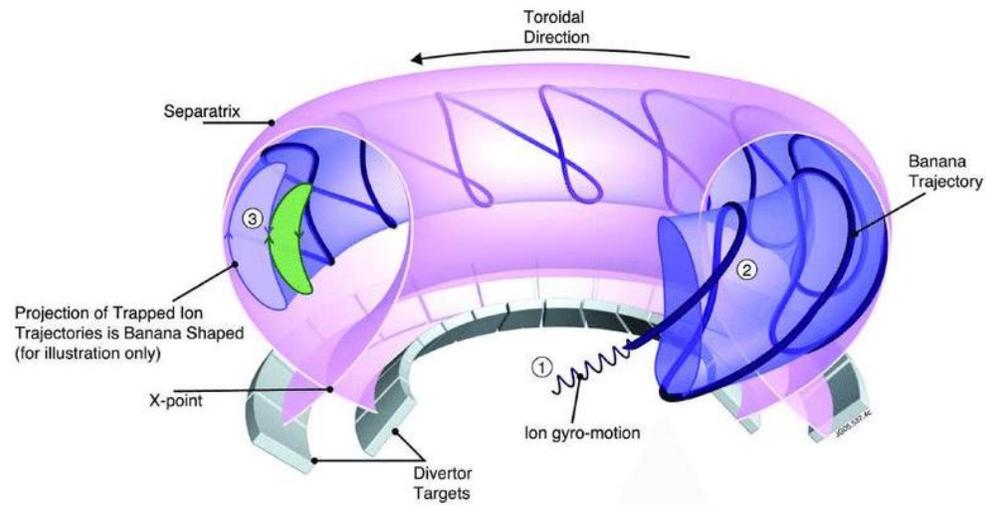
- Mitigate plasma facing component exposures to damaging conditions

A hierarchy of simulation fidelities are required to achieve self-consistent tokamak plasma solutions

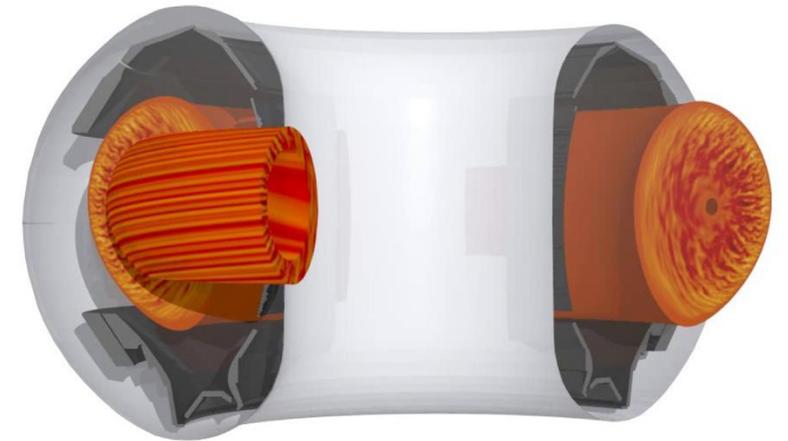
Physics of Gyrokinetic Turbulent Transport in Tokamaks



Classical Diffusion



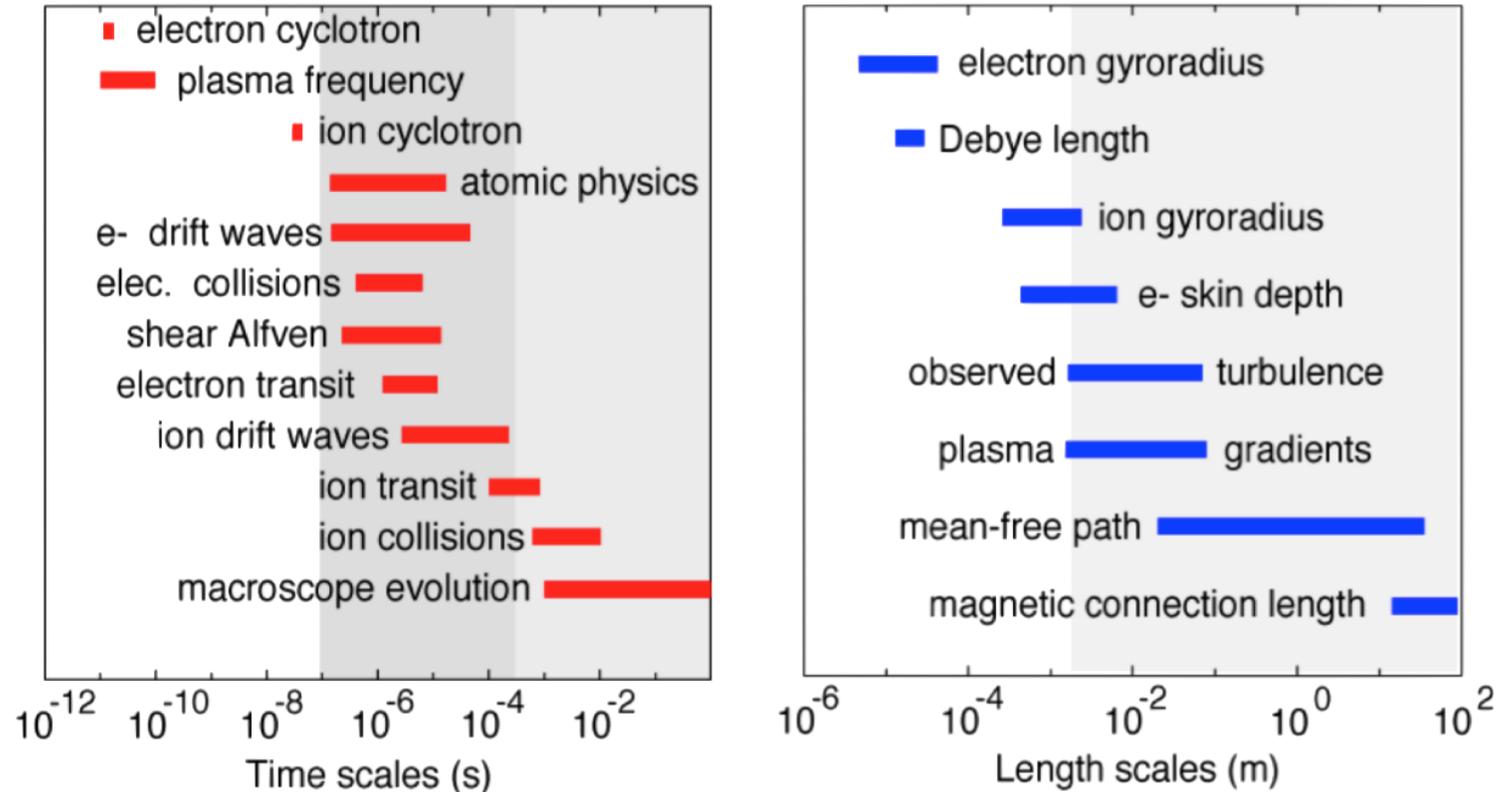
Neoclassical Transport



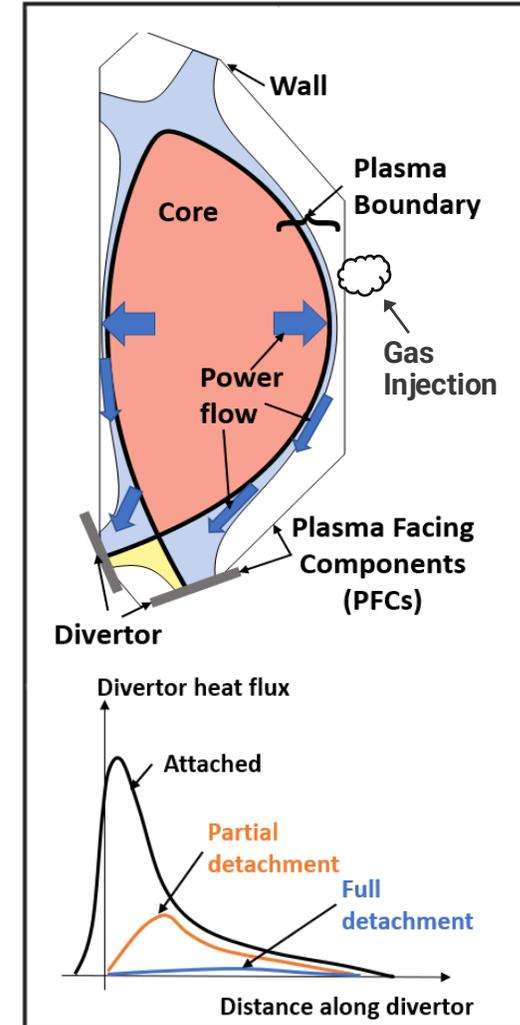
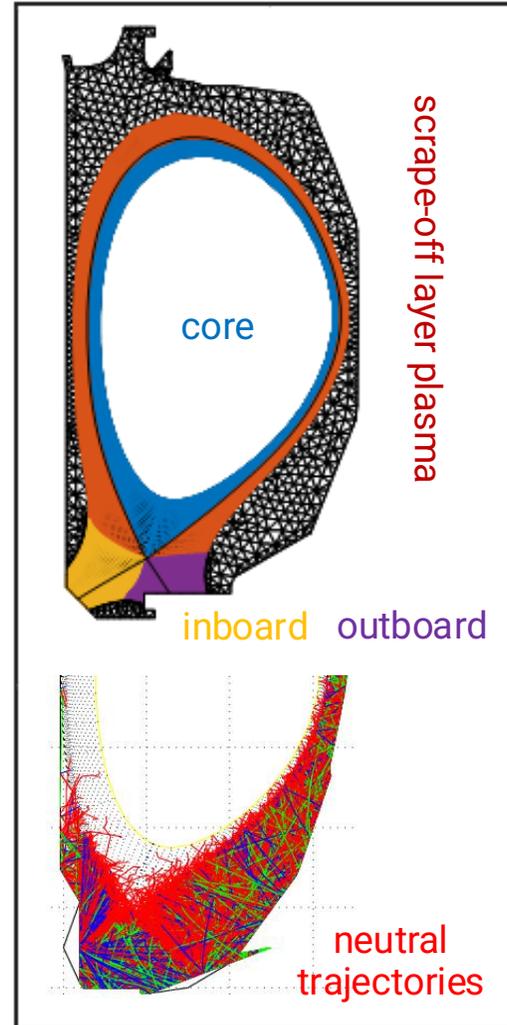
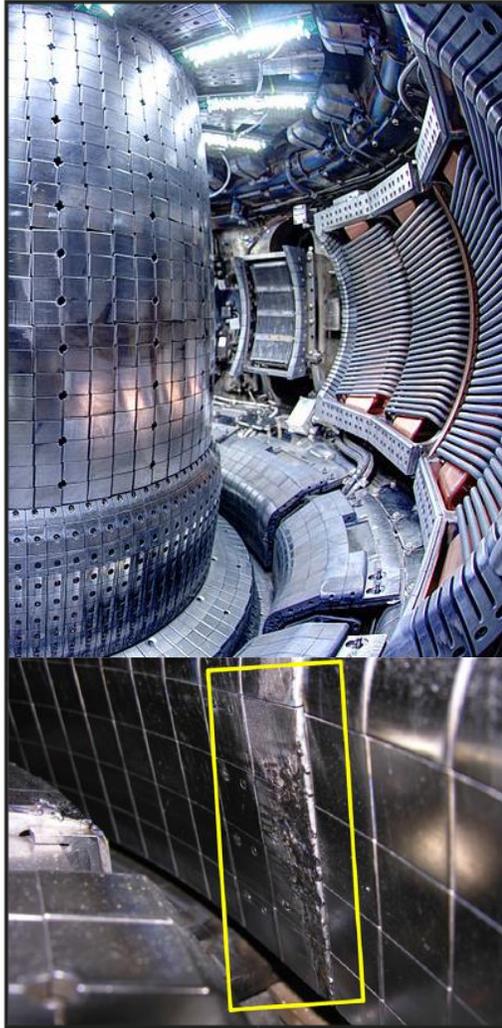
Turbulent Transport

Impact of Steep Gradients on Tokamak Edge Pedestal

Characteristic DIII-D Pedestal Scales

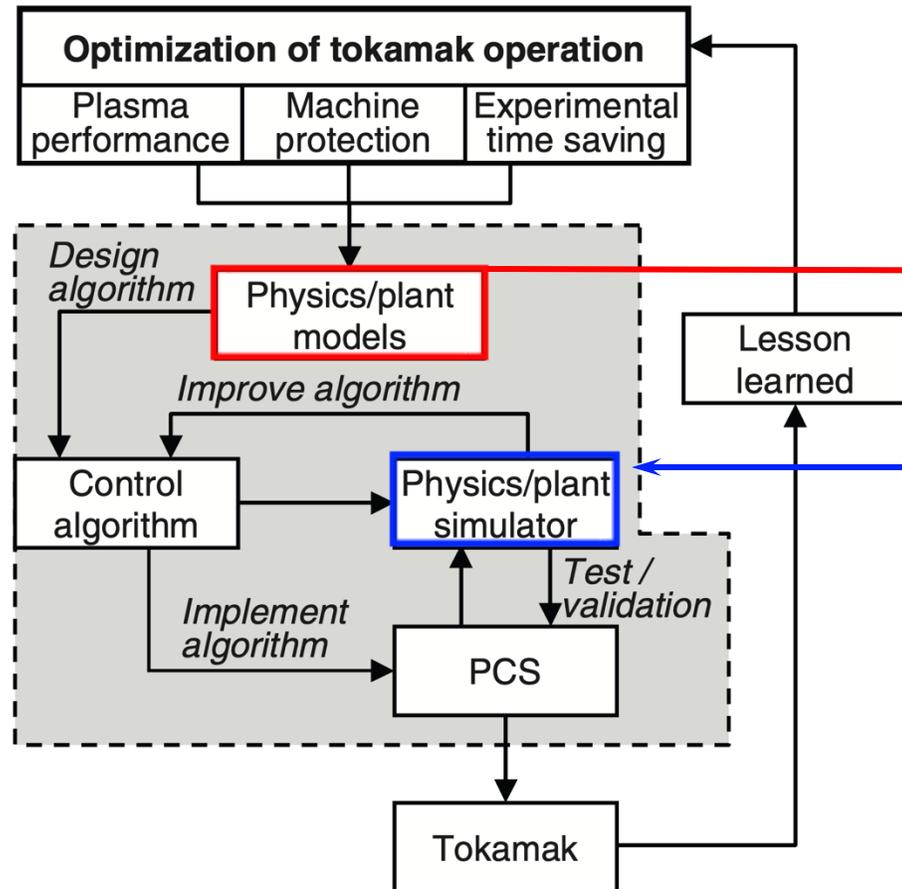


Consequence of Tokamak Power Exhaust and Particle Control

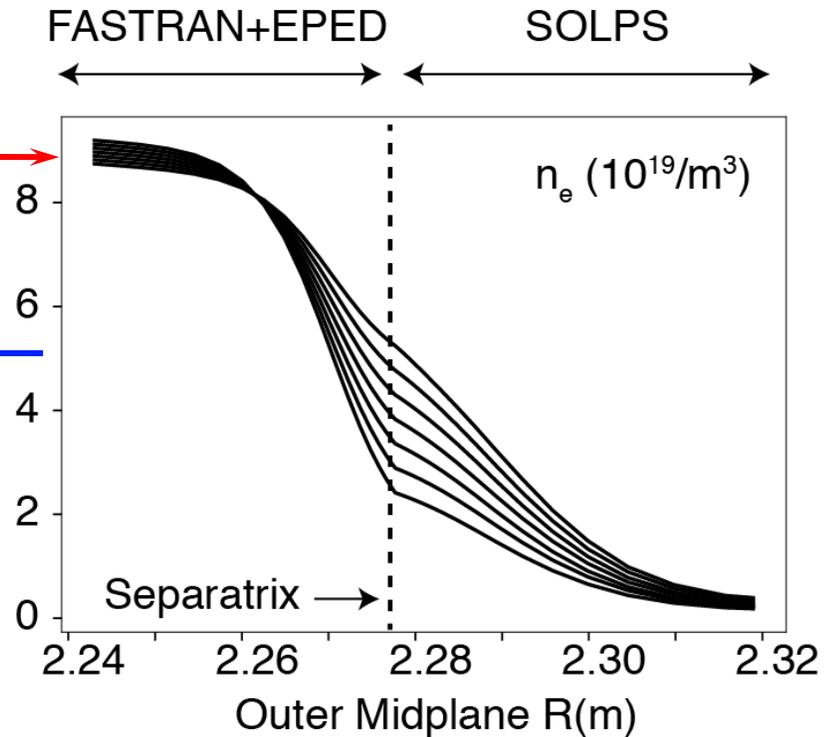


Components of Experimental Control and Reactor Design

Tokamak Operations

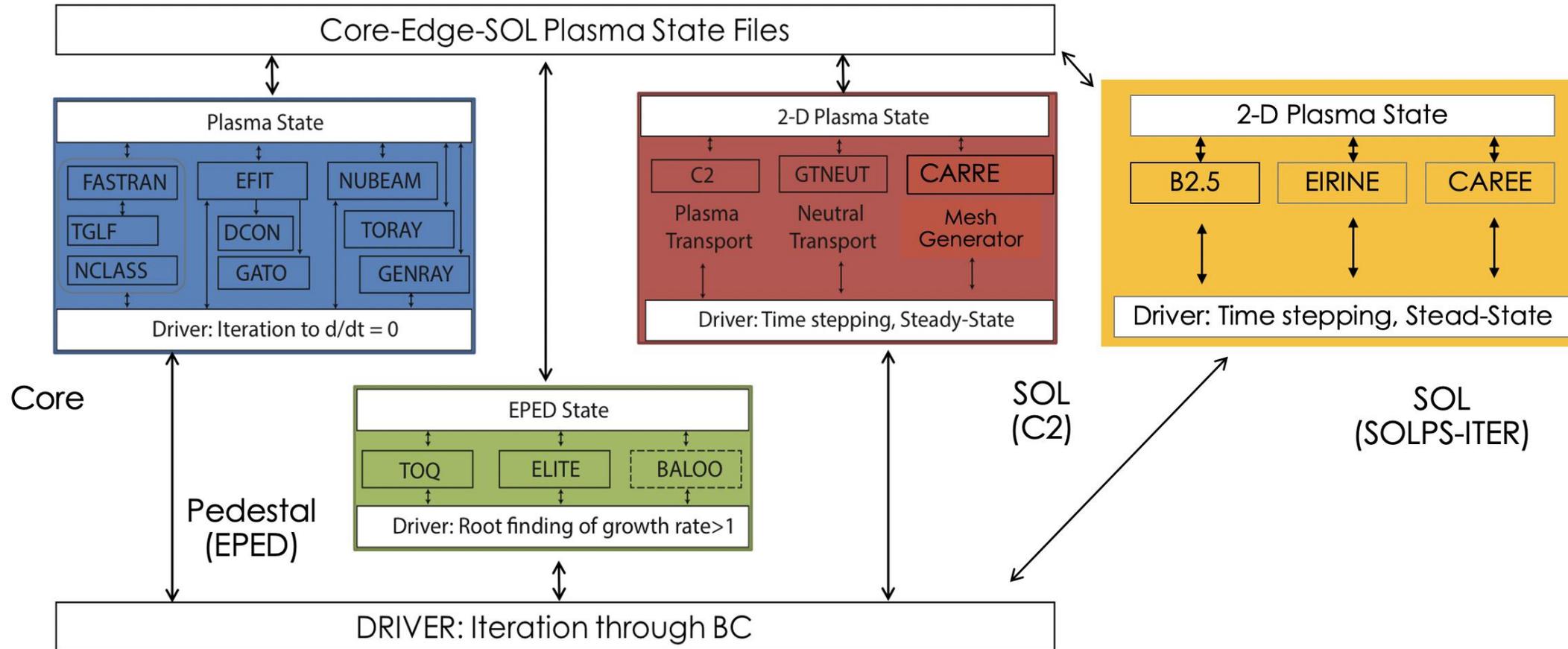


Tokamak Simulations



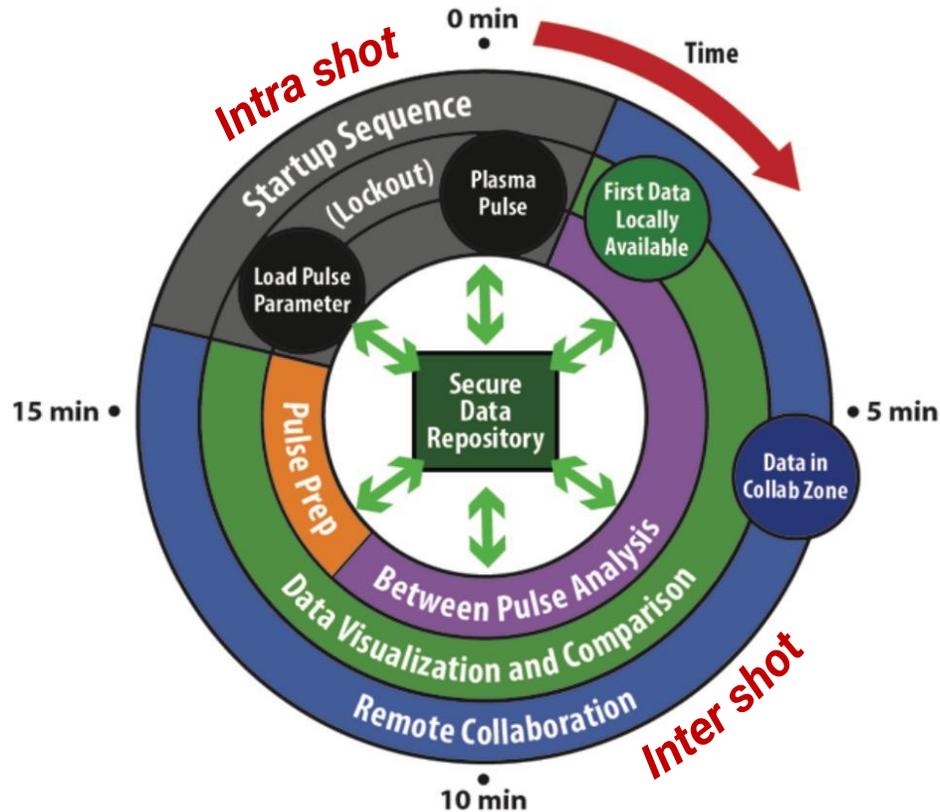
Components of Experimental Control and Reactor Design

Massively Parallel File-Based Model Coupling

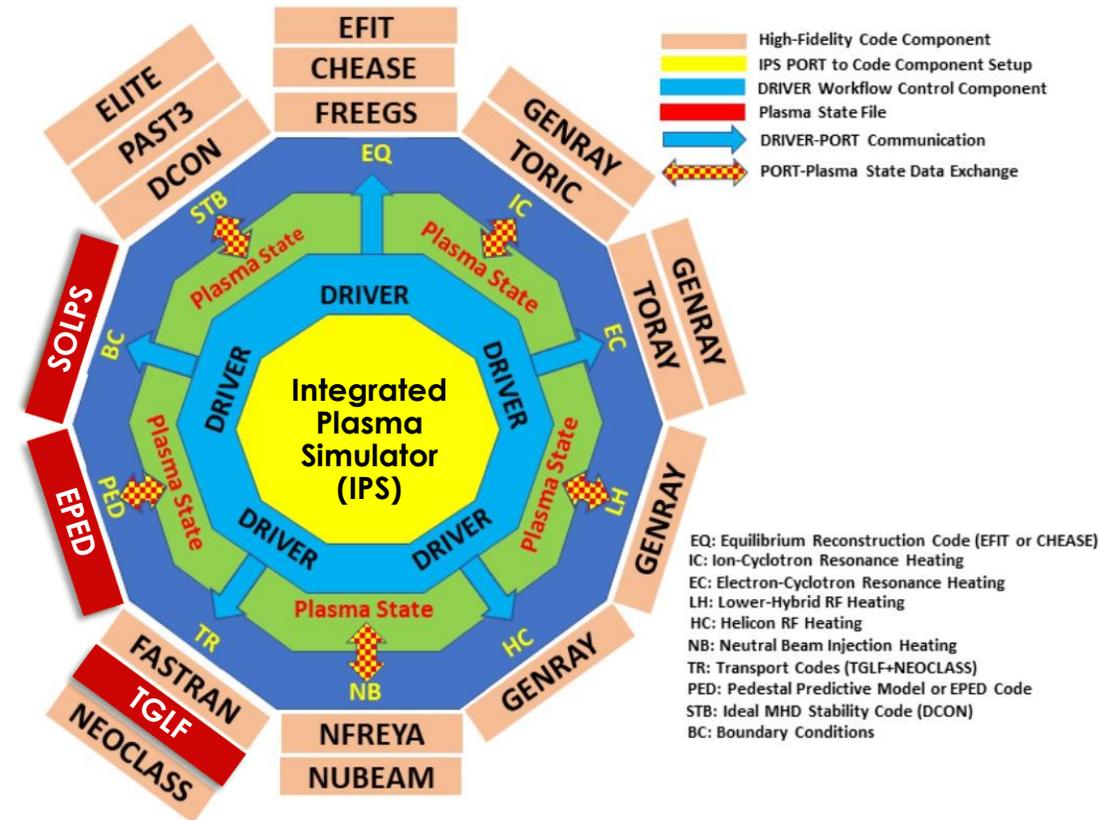


Components of Experimental Control and Reactor Design

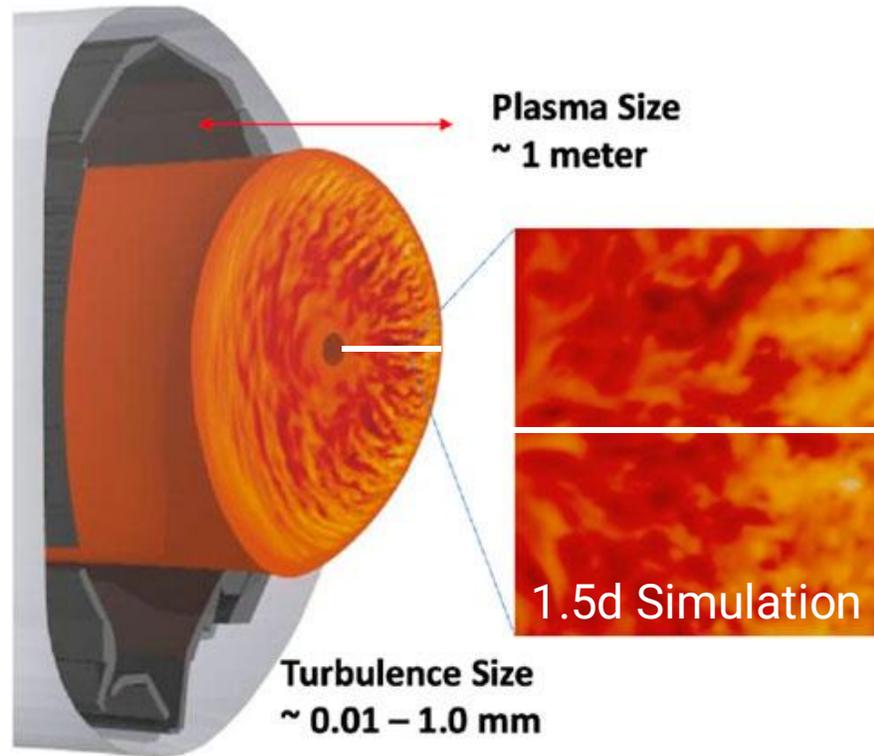
Tokamak Shot Cycle



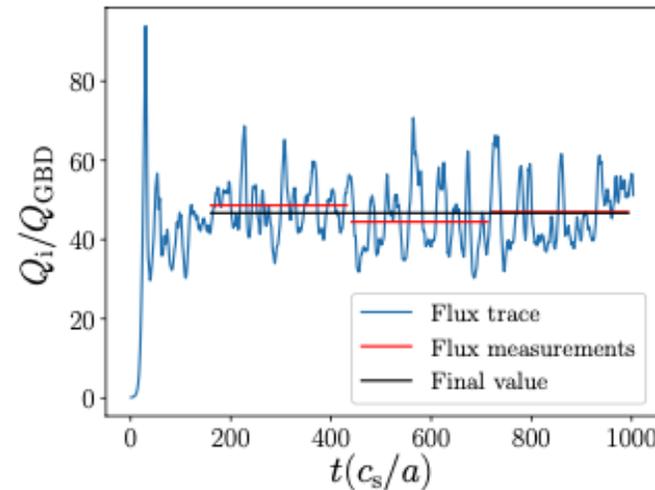
Integrated Plasma Simulator



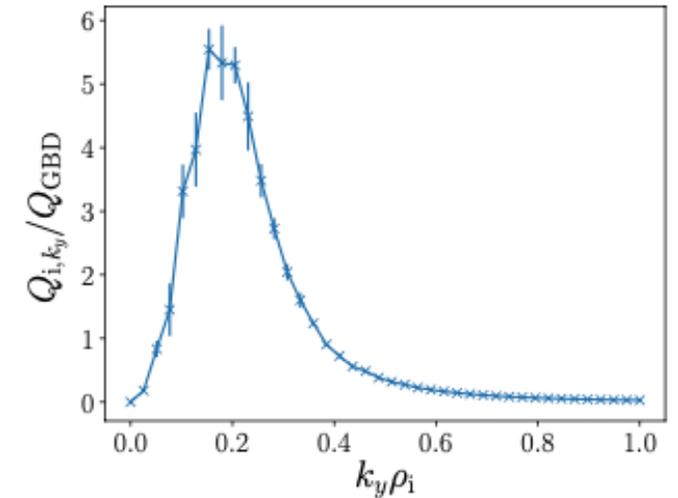
Reduced Models of Turbulent Transport Bridge Nonlinear Scales through Projected Linear Physics to Converged Saturation Level



Converged Fluxes



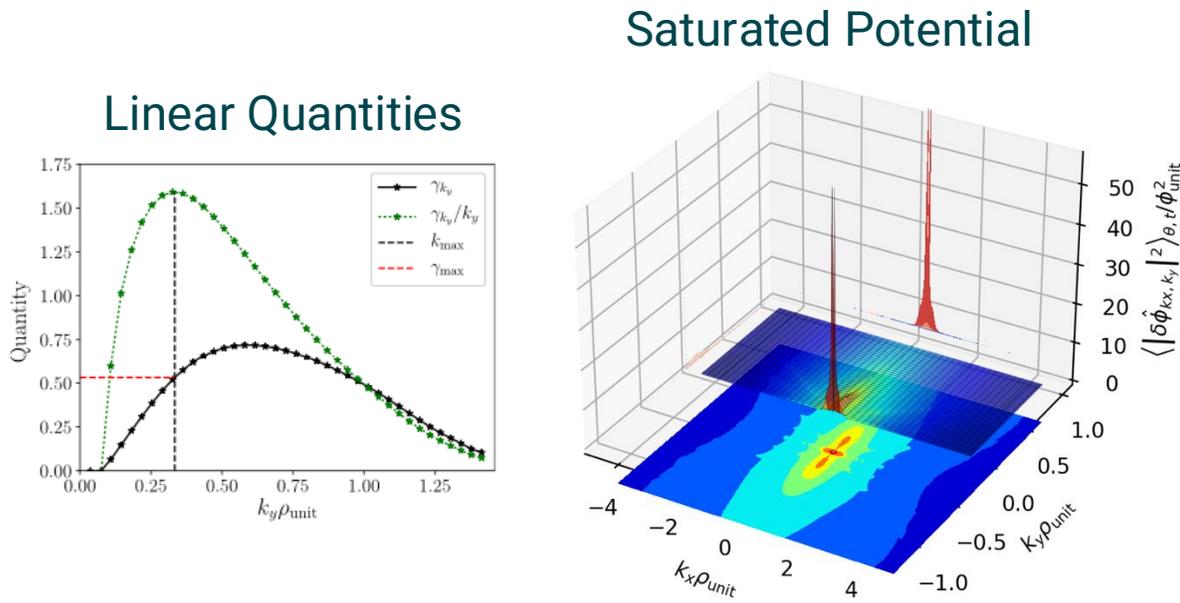
Saturation Rule



$$Q_s = 2 \sum_{k_y > 0} \Lambda_{s,k_y} W_{s,k_y}^L F_{s,k_y} \left[\frac{\langle |\delta \hat{\phi}_{k_y}|^2 \rangle_{x,\theta,t}}{\Delta k_y} \right] \Delta k_y$$

$$\frac{\langle |\delta \hat{\phi}_{k_y}|^2 \rangle_{x,\theta,t}}{\Delta k_y} = c_0 \sigma_{k_y}^{c_1}$$

Saturation Rules Employ Multi-Fidelity Modeling to Capture Complex Interaction Between Physical Plasma Parameters



$$k_y, \gamma, \omega \rightarrow \phi^2$$

Physical Parameter Space

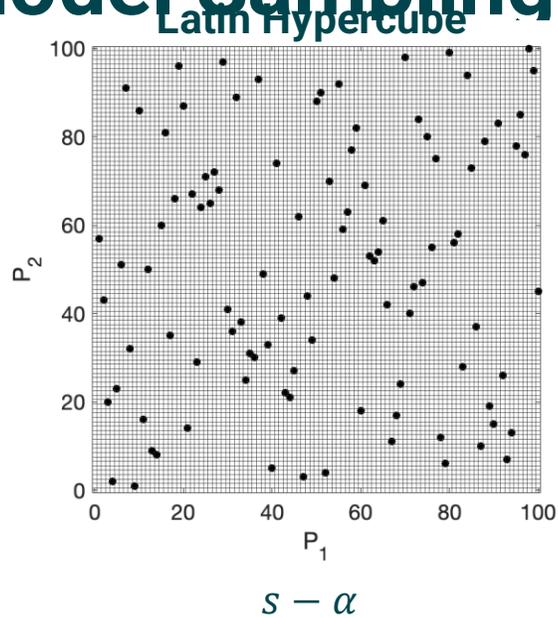
Nucl. Fusion **62** (2022) 096005

H.G. Dudding *et al*

Table 1. Details of the 43 nonlinear CGYRO simulations that form the database. Note that the labels in the second column correspond with those in figure 13(c), used for discerning the cases displayed in figures 13(a) and (b), 17 and 18.

Varied parameter	Values (label)	Fixed	Isotopes simulated
— (GA-std)	— (a)	—	H, D, T
$a/L_{Ti} = a/L_{Te}$	1.5 (b), 2.25 (c), 3.5 (d)	—	H, D, T
a/L_n	2.0 (e), 3.0 (f)	—	H, D, T
\tilde{s}	0.25 (g), 0.5 (h), 1.5 (i)	$\tilde{s}/w = 1/4$	D
$(a/c_s)\nu_{ee}$	0.01 (j)	—	H, D, T
$(a/c_s)\nu_{ee}$	1.0 (k)	—	D
T_i/T_e	0.5 (l), 1.5 (m)	—	D
q	1.5 (n), 2.5 (o)	—	D
κ	1.25 (p), 1.5 (q), 2.0 (r)	—	D
Δ	-0.125 (s), -0.25 (t), -0.5 (u)	—	D
r_0/R_0	1/4 (v), 1/12 (w)	—	D
$(a/c_s)\nu_{ee}$	0.01 (x)	$a/L_n = 3.0$	H, D, T
$(a/c_s)\nu_{ee}$	0.05 (y)	$a/L_n = 3.0$	H, T
$(a/c_s)\nu_{ee}$	1.0 (z)	$a/L_n = 3.0$	D

Global Quasilinear Transport is Governed by High-Dimensional Parameter Space Prohibiting Comprehensive Model Sampling



<https://zenodo.org/records/3274234>

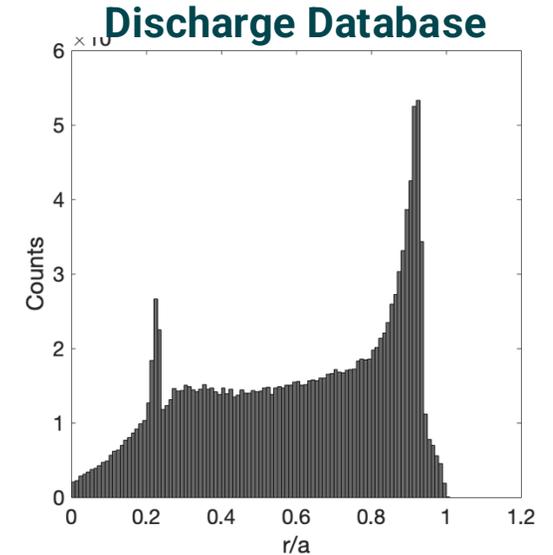
QuaLiKiz 4D, Citrin et al., (2015), *NF*, 55, 092001
 QuaLiKiz 6D, Citrin et al., (2023), *PoP*, 30, 062501
 QuaLiKiz 10D, van de Plassche (2020), *PoP*, 27, 022310

- Express curse of dimensionality when increasing degrees of freedom

QuaLiKiz 15D, Ho et al., (2020), *PoP*, 28, 032305

Dataset	Points	Time (CPU h)
10D	$\sim 3 \times 10^8$	$\sim 10^6$
Projected 15D	$\geq 7.2 \times 10^{10}$	$\geq 3 \times 10^8$
Actual 15D	3.38×10^7	3.5×10^5

- Supplement training data through experimental observations



$R(r, \theta | \kappa, \delta)$ $Z(r, \theta | \kappa, \delta)$

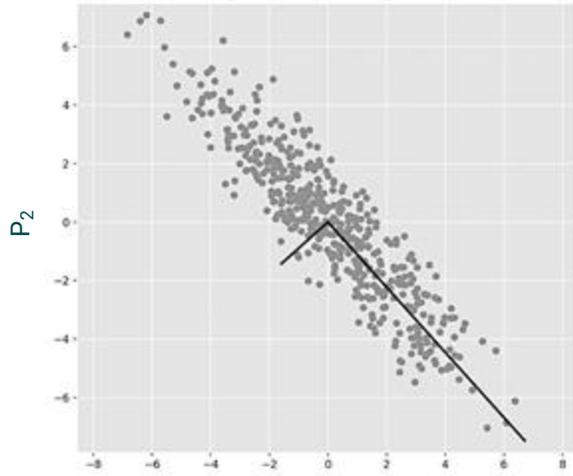
<https://zenodo.org/records/3274234>

TGLF 21D, Menenghini et al., (2014), *PoP*, 21, 060702
 TGLF 23D, Menenghini et al., (2017), *NF*, 57, 086034

Fully 5D gyrokinetic simulations take on the order of 10^4 hours of wall clock time per radial point

Local Dimensionless Plasma Parameters Have Associated Correlations and Accessibility in Present and Future Device

Principal Components



$$M = \sum_{i=1}^n \sigma_i u_i v_i^*$$

$$\|M - M_r\|_F = \sqrt{\sigma_{r+1}^2 + \dots + \sigma_N^2}$$

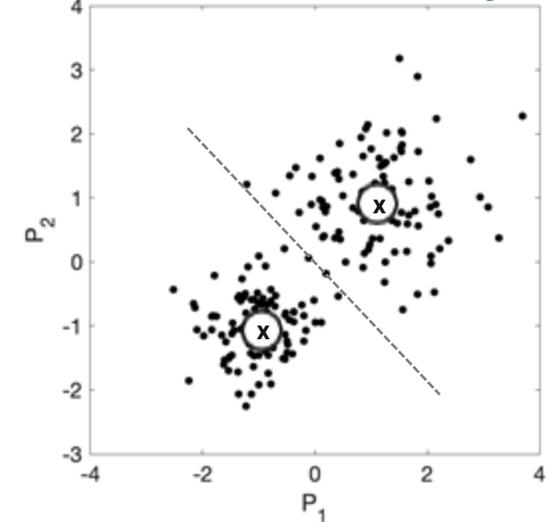
- Gyrokinetic validation data is several orders of magnitude more expensive

TGLF-GKNN, Neiser et al., (2024) TTF Conference

Saturation Rule	Validation	Machines
SAT0,1,2,3	CGYRO	DIII-D/MAST-U

- Supplement quasilinear model performance through residual

K-Means Clustering



$$\mu_i = \frac{1}{|S_i|} \sum_{y \in S_i} y$$

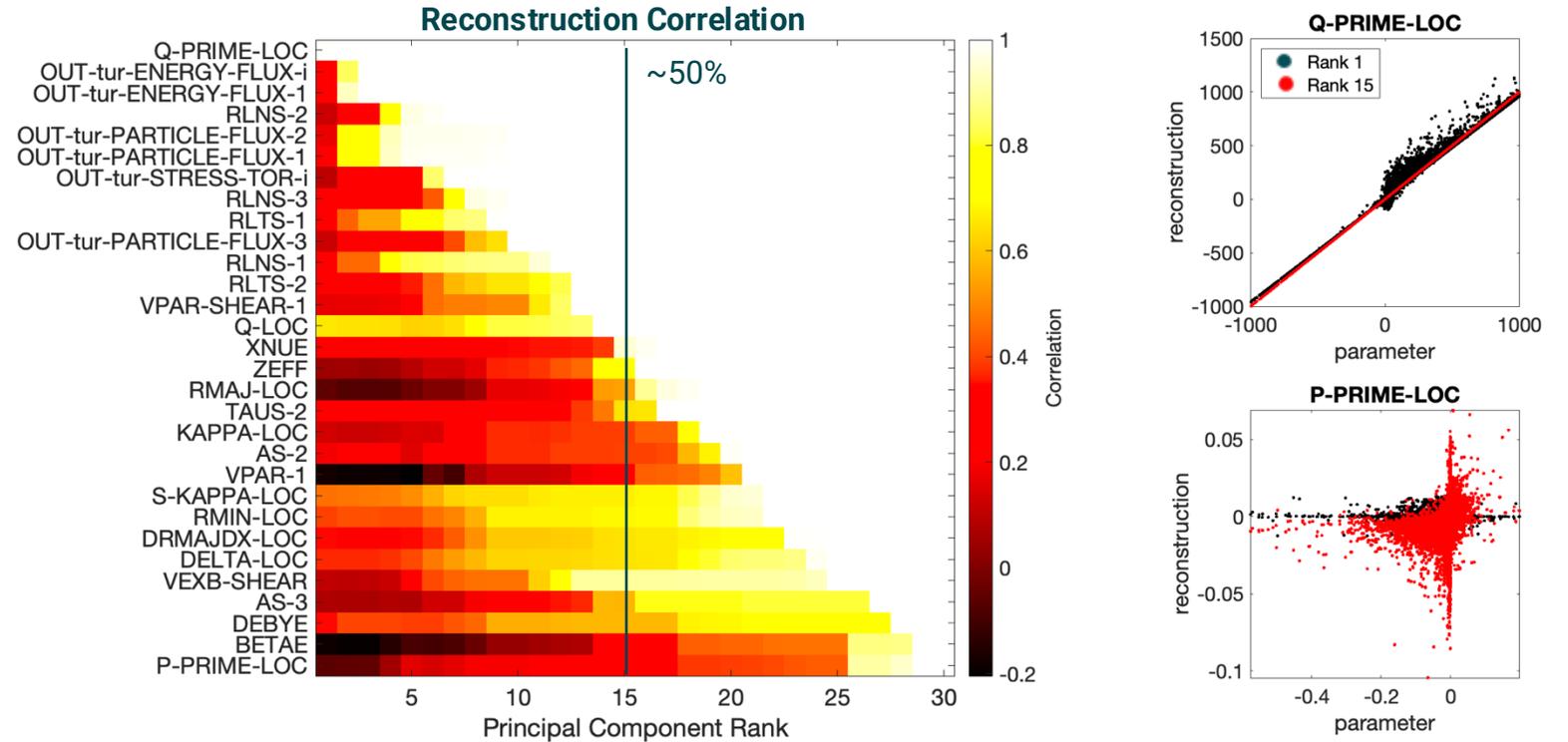
$$\sum_{i=1}^k \sum_{x \in S_i} \|x - \mu_i\|^2 = \sum_{i=1}^k |S_i| \text{Var } S_i$$

Quasilinear models, such as the **Trapped Gyro Landau Fluid (TGLF)**, reduce this cost down to **10⁰ seconds**

Matrix Factorization Cannot Fully Capture Nonlinear Relationships in Feature Space of Simulation Domain

Variable	Definition
r/a	Normalized minor radius
R/a	Normalized major radius
κ	Elongation
$r \frac{\partial \kappa}{\partial r}$	Elongation shear
δ	Triangularity
$\frac{\partial R}{\partial r}$	Shafranov shift
q	Safety factor
$\frac{q^2 a^2}{r^2} \frac{\partial q}{\partial r}$	Safety factor shear
β_e	Kinetic to magnetic pressure ratio
ν_{ie}/ac_s	Collision frequency
T_i/T_e	Ion to electron temperature ratio
n_D/n_e	Deuterium to electron density ratio
n_C/n_e	Carbon to electron density ratio
Z_{eff}	Effective ion charge
a/L_{Te}	Electron temperature scale length
a/L_{Ti}	Ion temperature scale length
a/L_{ne}	Electron density scale length
a/L_{nD}	Deuterium density scale length
a/L_{nC}	Carbon density scale length
$\frac{qa^2}{r} \frac{\partial p}{\partial r}$	Total pressure gradient
$\text{sign}(I_p) R \omega_{\text{tor}} \frac{a}{c_s}$	Parallel velocity
$-\text{sign}(I_p) R \frac{\partial \omega_{\text{tor}}}{\partial r} \frac{a}{c_s}$	Parallel velocity gradient
$-\text{sign}(I_p) \frac{r}{q} \frac{\partial}{\partial r} \frac{V_{E \times B}}{R} \frac{a}{c_s}$	$E \times B$ velocity shear

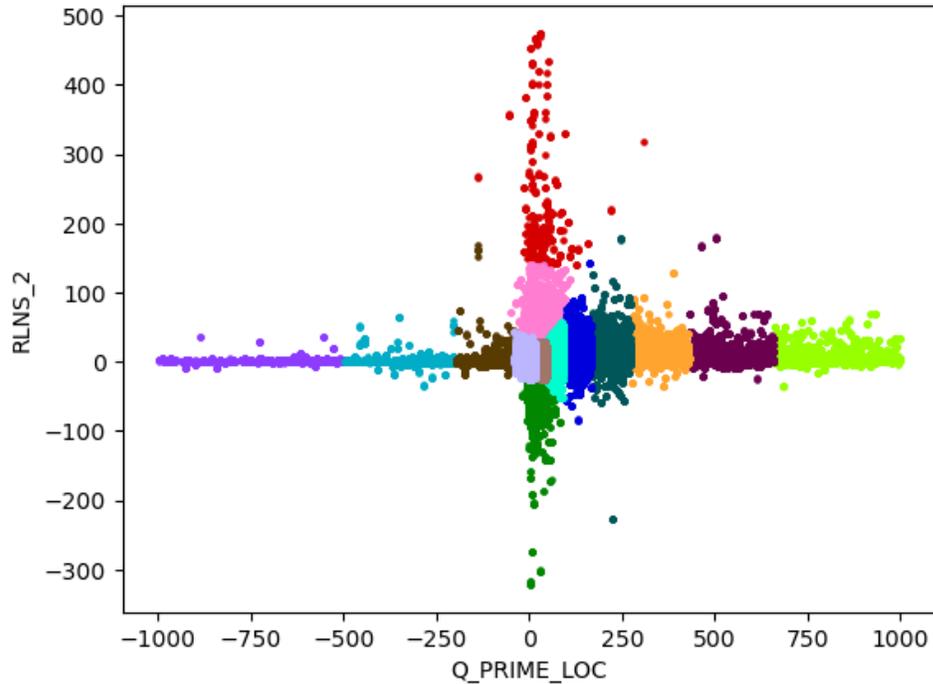
Menenghini et al., (2017) *NF*, 57, 086034



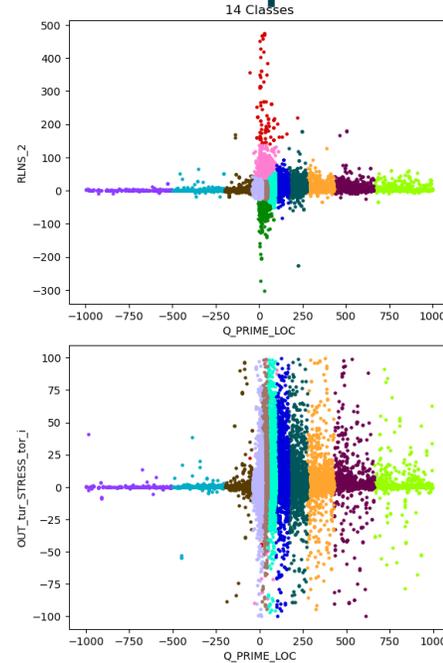
Surveyed database of DIII-D TGLF simulations spans **24D inputs** and **6D outputs** over range of experimental conditions and nearly 1.6 million points

Vector Quantization Groups Nearest Neighbor Simulations Separating Out Over-Sampled Parameter Regions in Domain

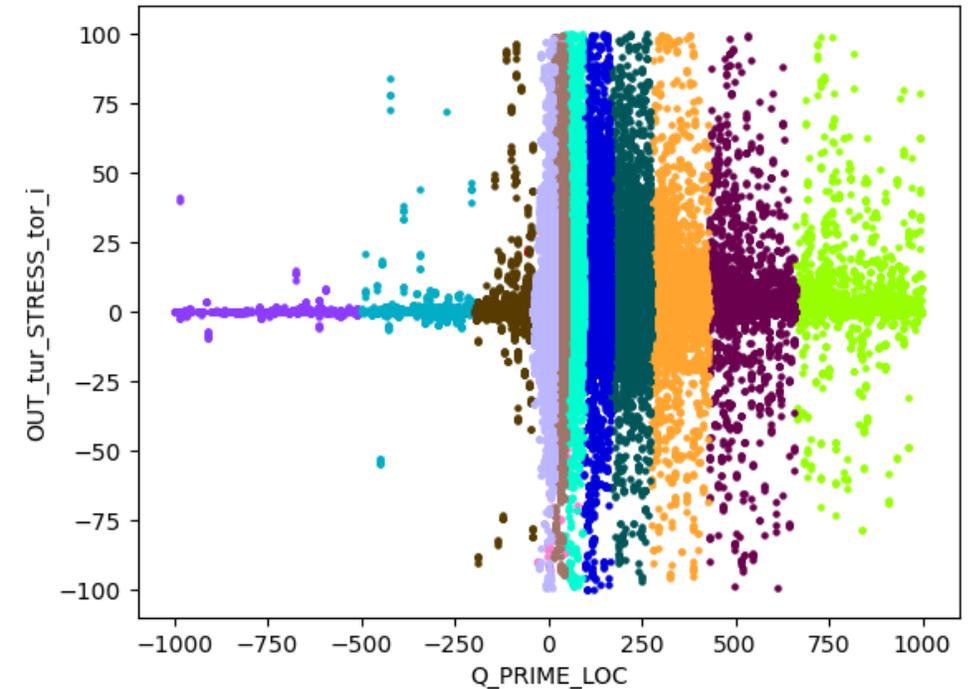
Input by Input



20% Compression



Input by Output

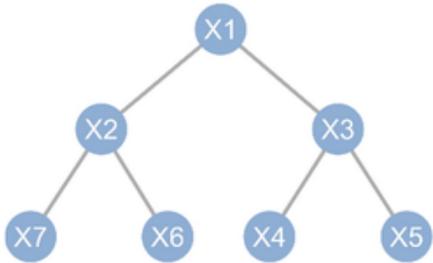


Dataset can be compressed by nearly an **order of magnitude** when clustered in the input space alone, reducing the number of simulations required for machine learning

Accelerated Predictions of Quasilinear Transport Simulations in Many Degrees of Freedom Are Enabled by Machine Learning

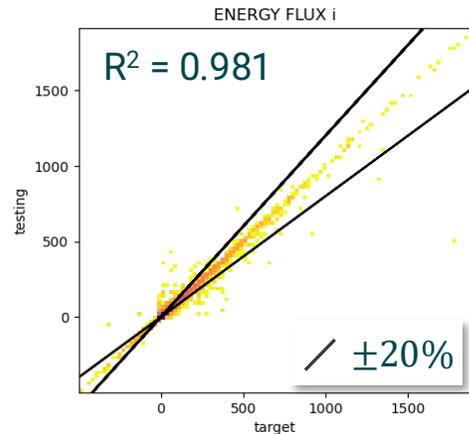
XGBoost: gradient boosted split level branching tree

- memory efficient
- fast to train
- extensible structure

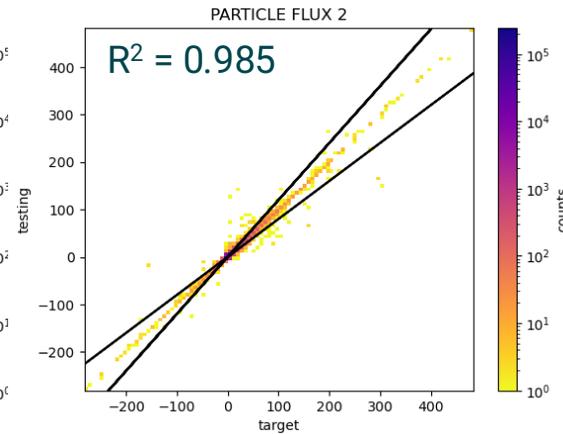


$$\hat{y} = \sum_{k=1}^n f_k(x)$$

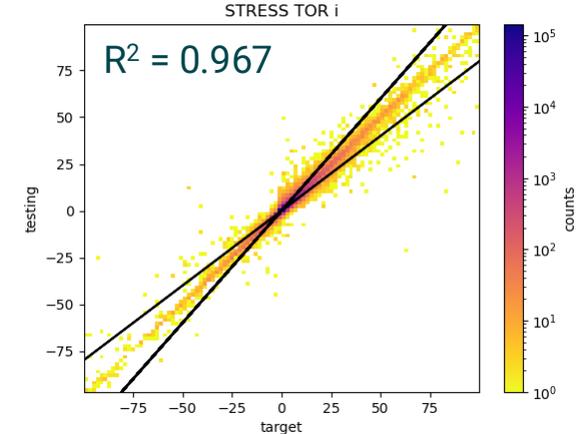
Energy Flux



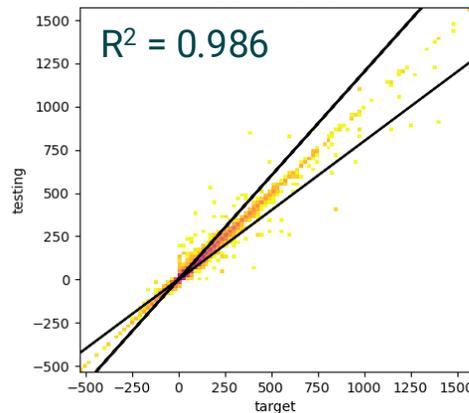
Particle Flux



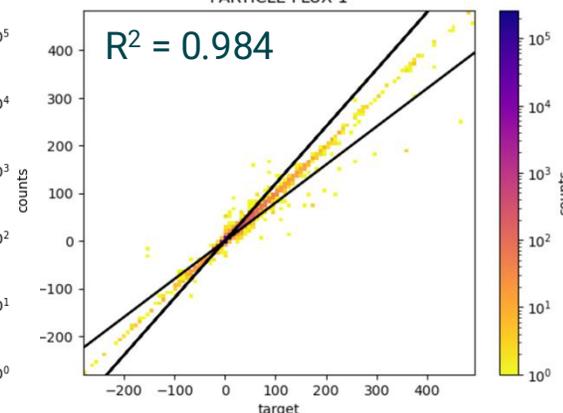
Momentum Flux



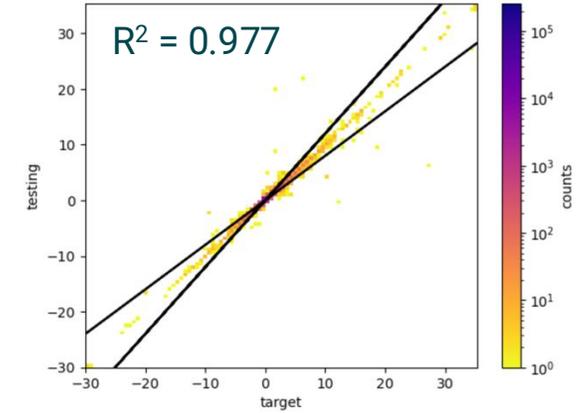
ENERGY FLUX 1



PARTICLE FLUX 1



PARTICLE FLUX 3

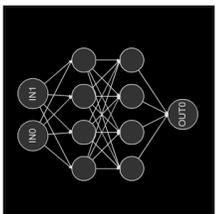


Impurity Flux

Sensitivity of Surrogate Modeling Framework is Captured by Computationally Tractable Calculation Over Large Datasets

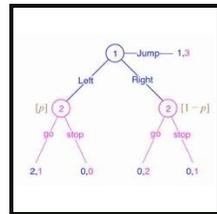
Shapley values: game theoretic description of importance distribution over possible features

1. additive explanation of each parameter for each prediction
2. interaction of each parameter to overall model behavior
3. dominant feature selection for each independent output



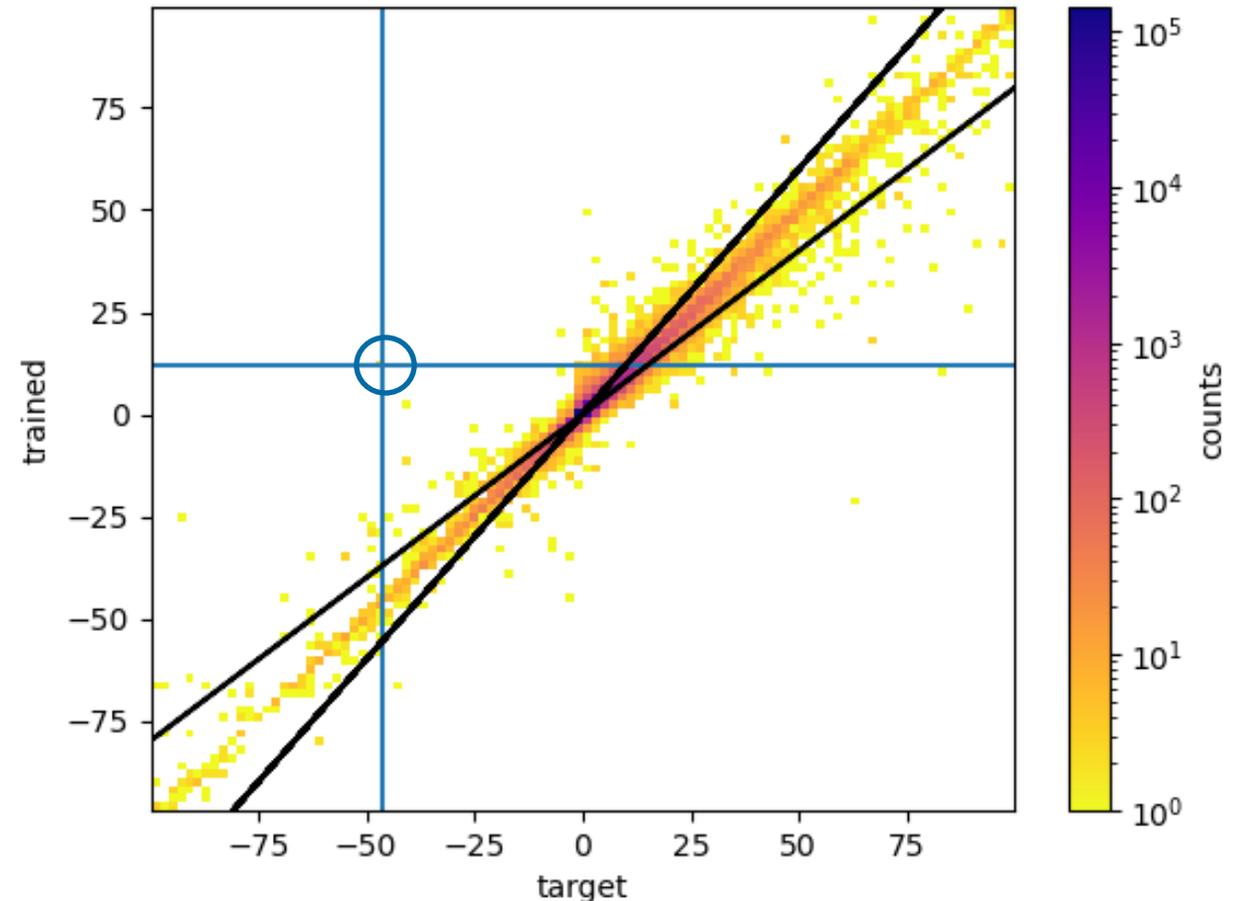
Neural Network

interpretable



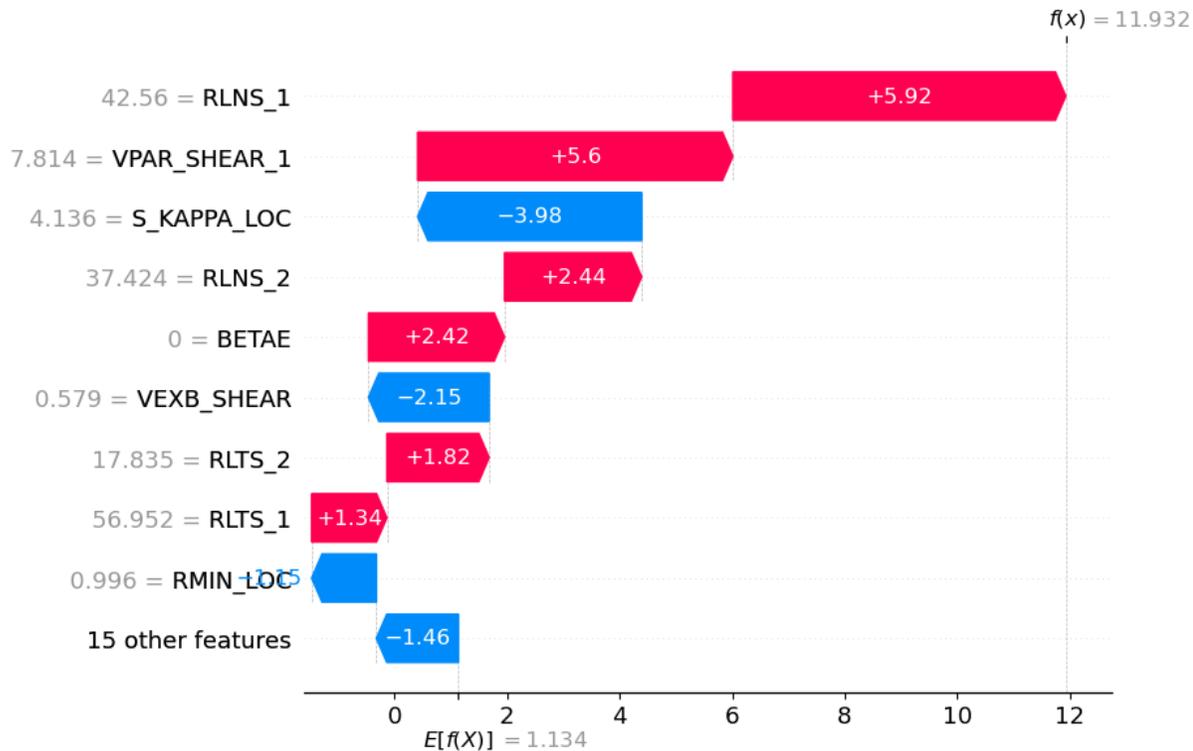
Decision Tree

Ion Momentum Flux

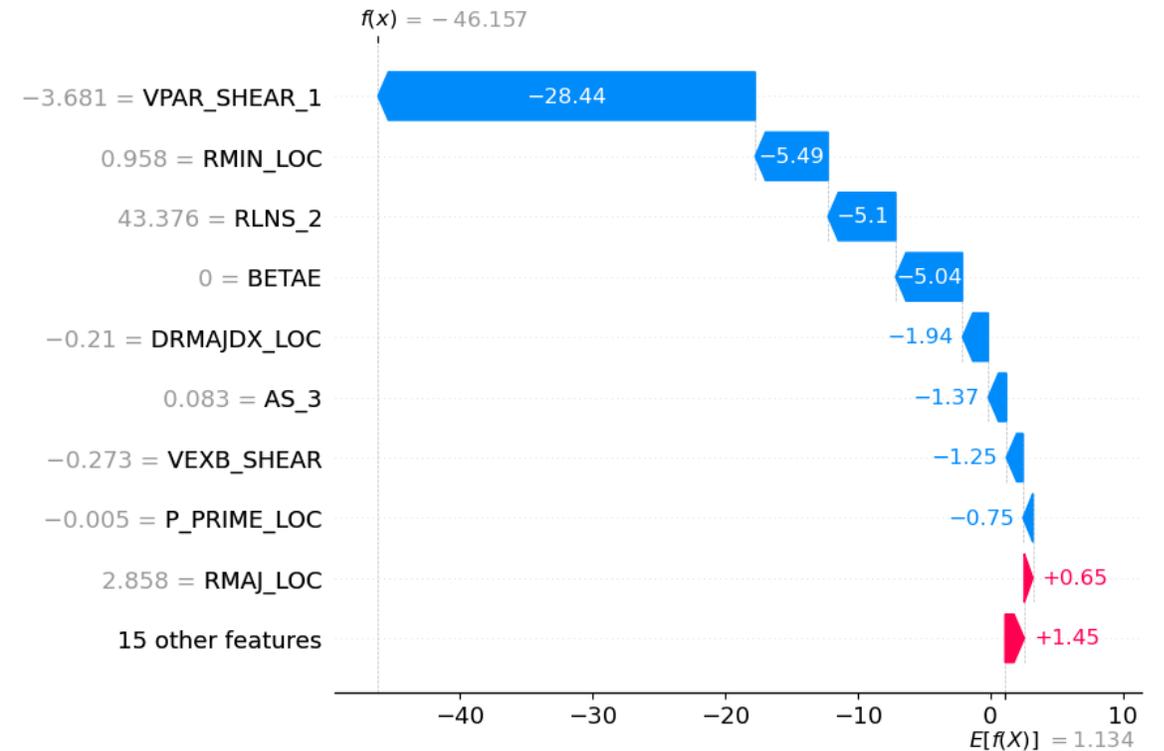


Target Predictions are Evaluated in Terms of Parameter Marginal Contributions Across All Possible Coalitions

Outlier Model Prediction



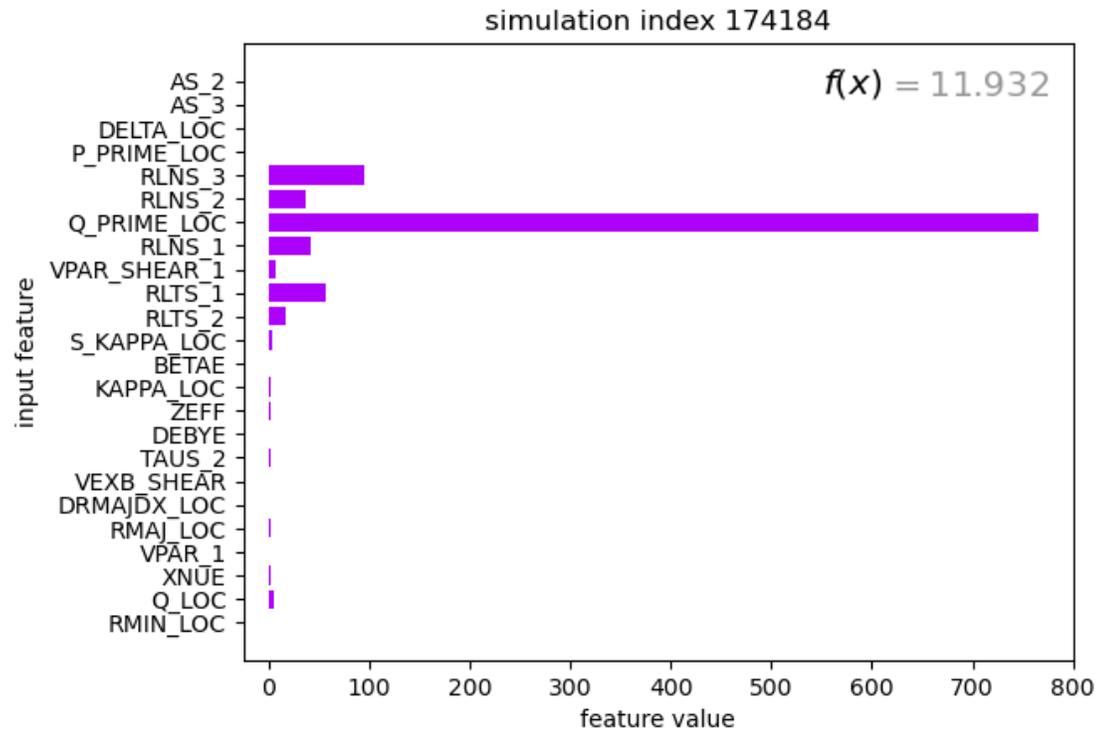
Matched Model Prediction



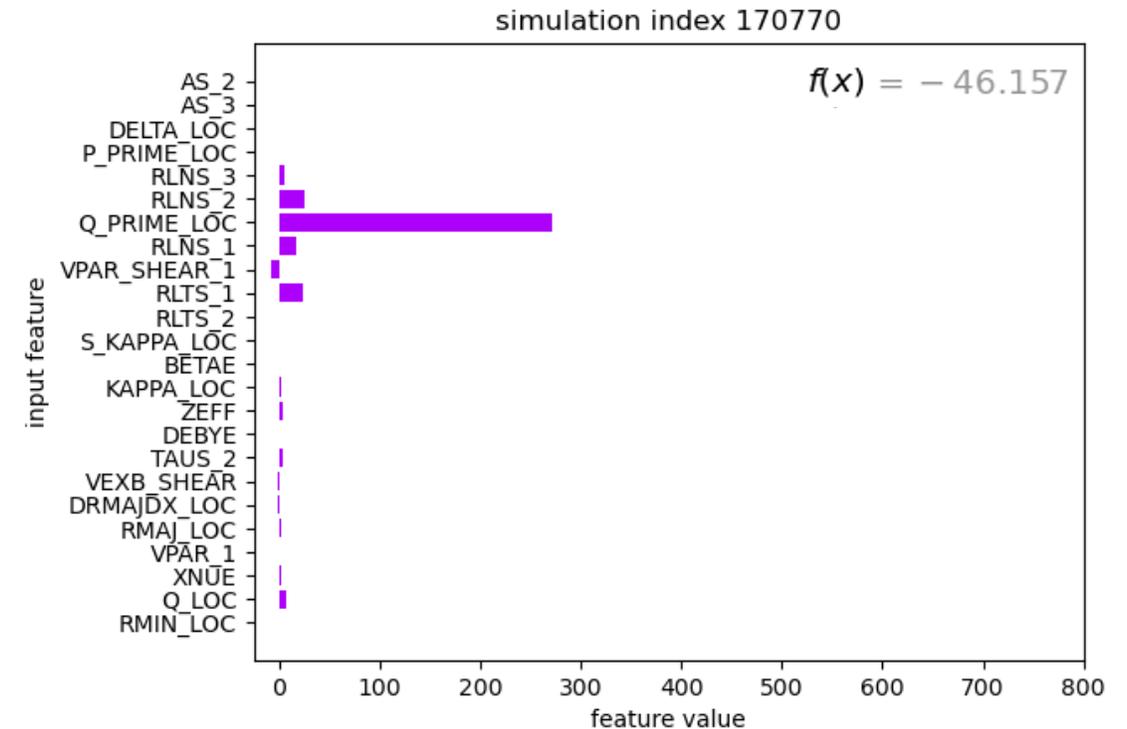
Predictions exhibit contrasting **characteristic response** to model input parameters

Transport Simulations Can be Augmented with Expanded Search Around Particular Coordinate Set Combinations

3 Neighbors Identified as Outliers

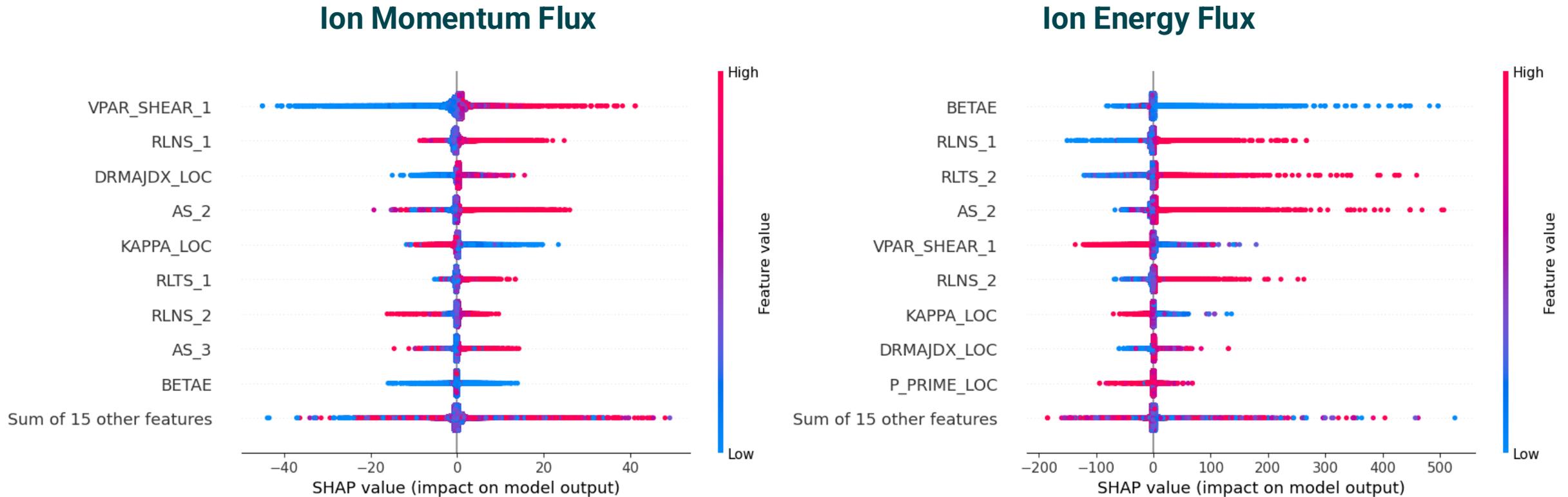


52 Neighbors within 10% of Target



Simulations **sparingly cover** proximal experimental conditions near extrema

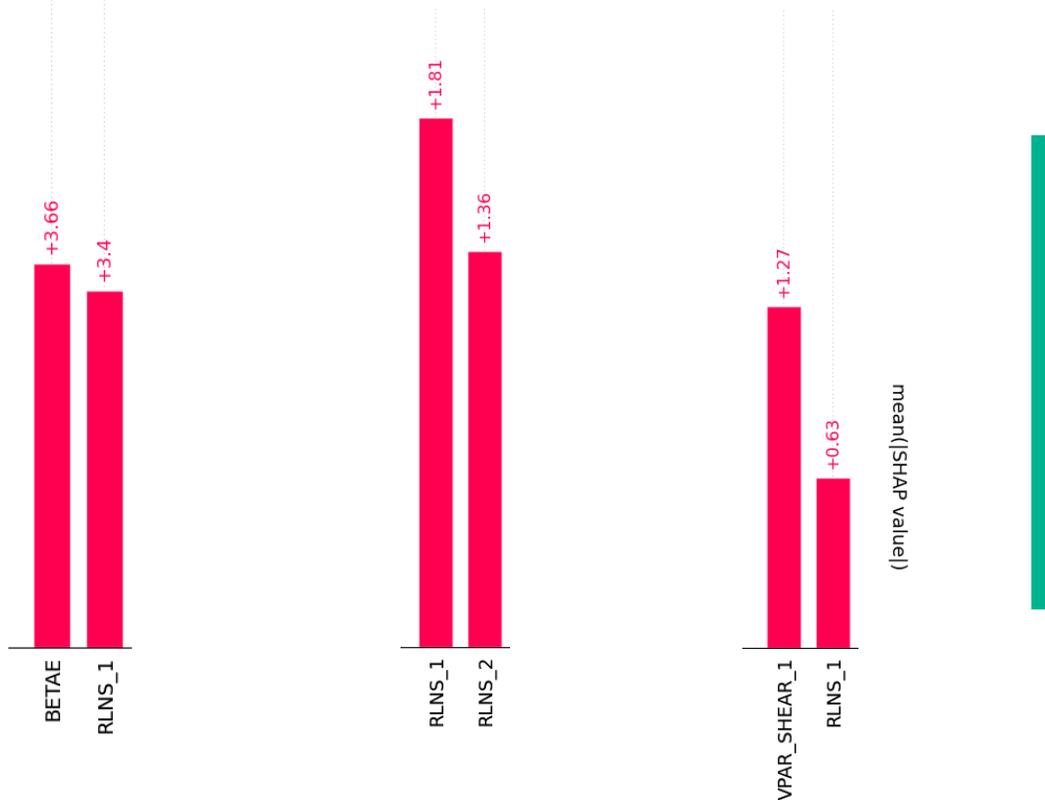
Survey of Database Indicates Parameter Interaction Strengths In Terms of Degree and Direction of Model Importance



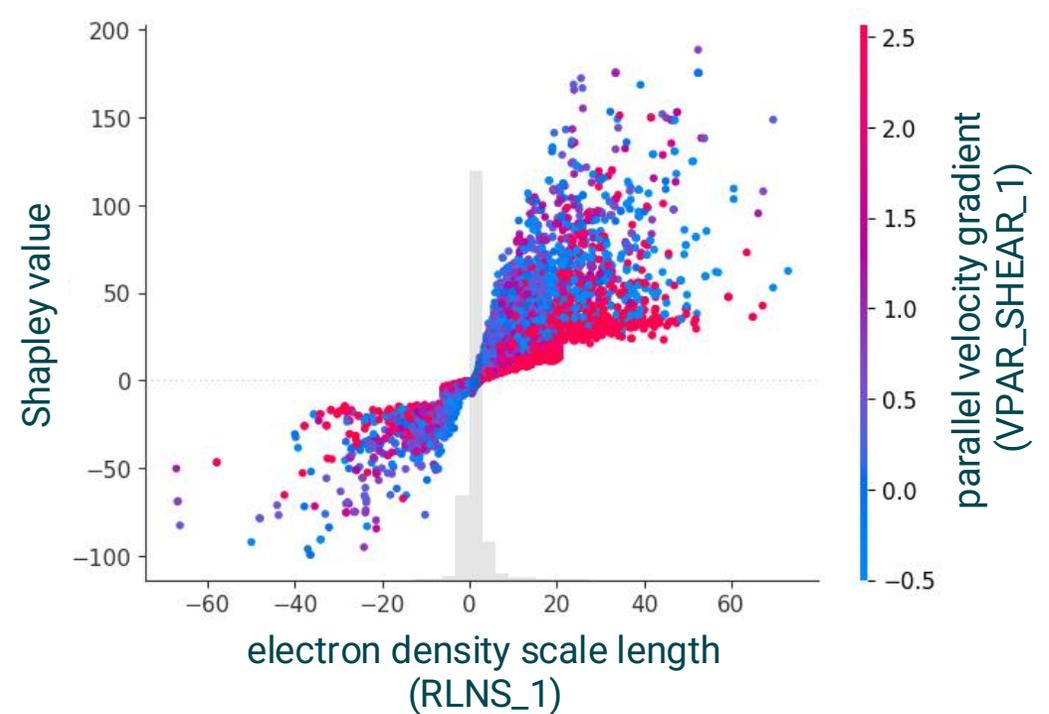
Model sensitivity reveals **significant partial populations** providing minimal impact

Explainable Machine Learning Unlocks Physical Insight to Determine Parameter Interdependence for Efficient Sampling

Decoupled Dominant Features



Paired Model Contributions

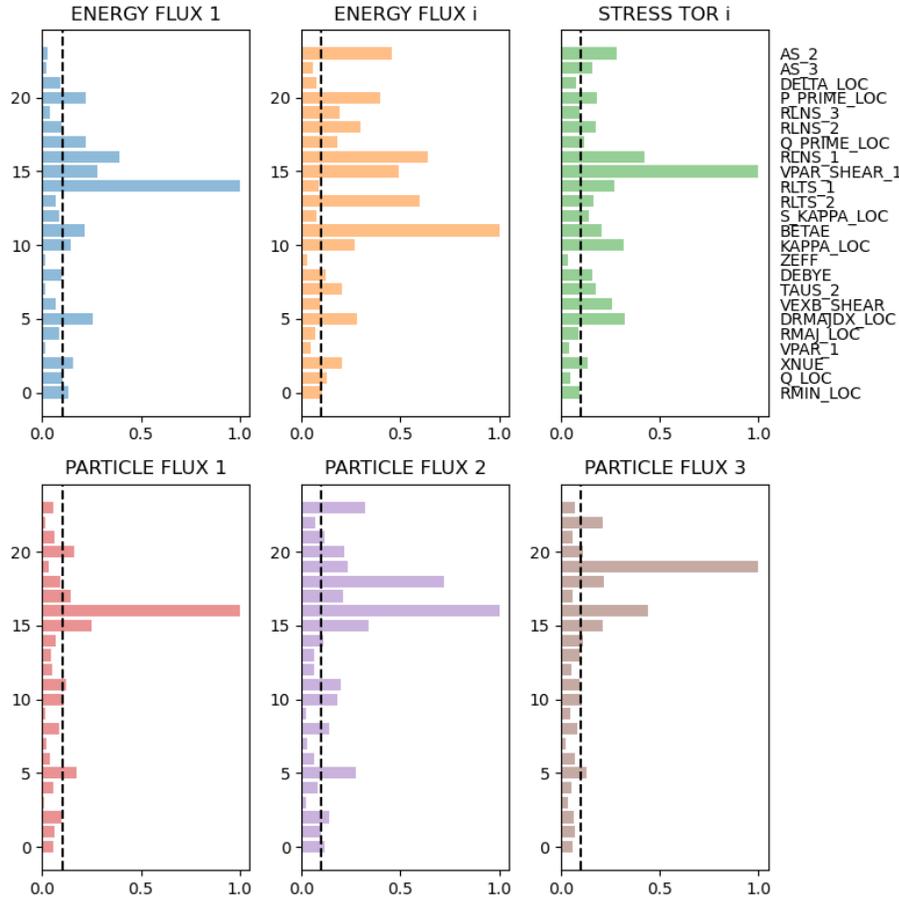


Ion Energy Flux Ion Particle Flux Momentum Flux

Electron Particle Flux

Explainable Machine Learning Unlocks Physical Insight to Determine Parameter Interdependence for Efficient Sampling

Total Parameter Contributions

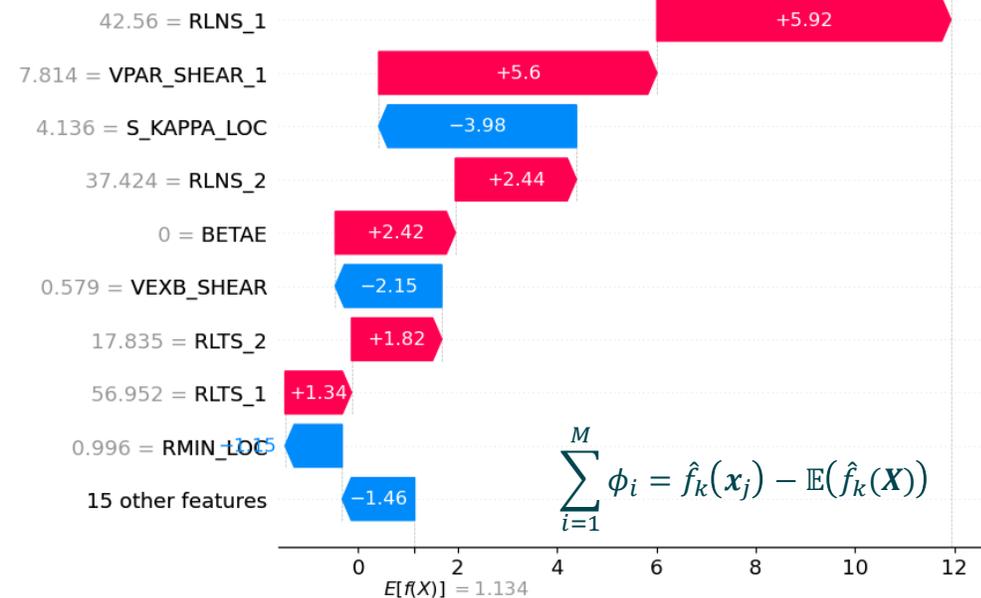


Shapley Value Assessment

$$\phi_i(v) = \frac{1}{n!} \sum_{S \subseteq N \setminus \{i\}} |S|! \cdot (n - |S| - 1)! \cdot (v(S \cup \{i\}) - v(S))$$

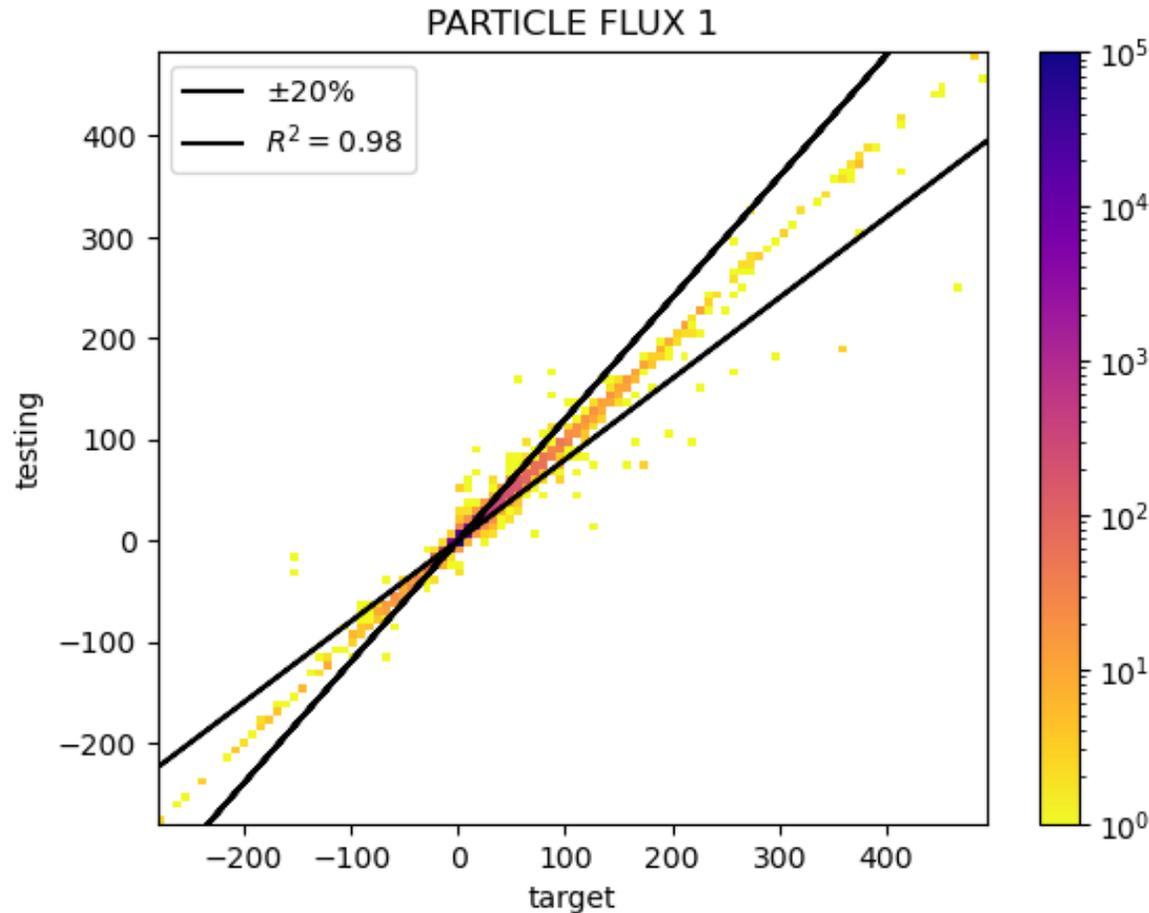
Annotations:
 - $\frac{1}{n!}$: average
 - $|S|! \cdot (n - |S| - 1)!$: multiplicity
 - $(v(S \cup \{i\}) - v(S))$: contribution value
 - $v(S \cup \{i\})$: coalition value

932

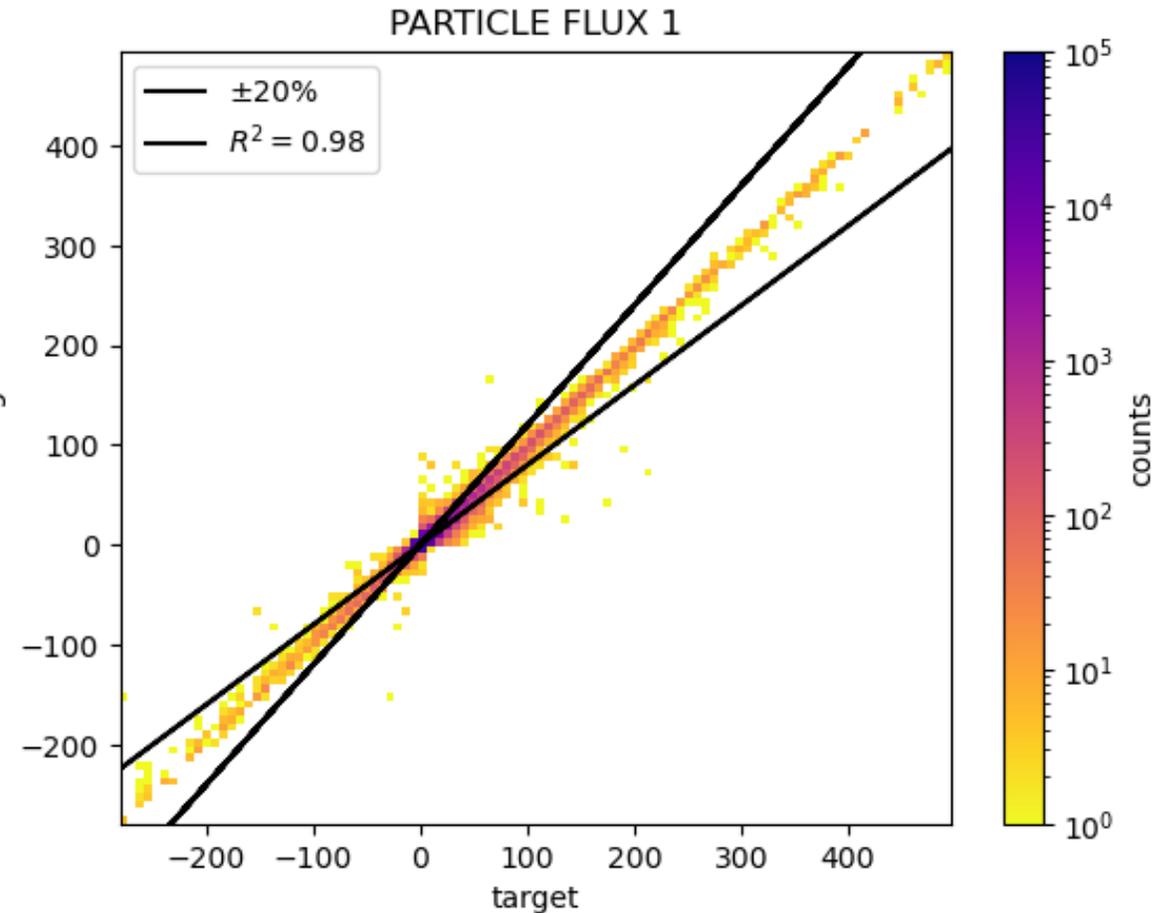


Drastic Reduction in Degrees of Freedom by Feature Importance Sampling of Surrogate Model Shapley Values

24 Input Components

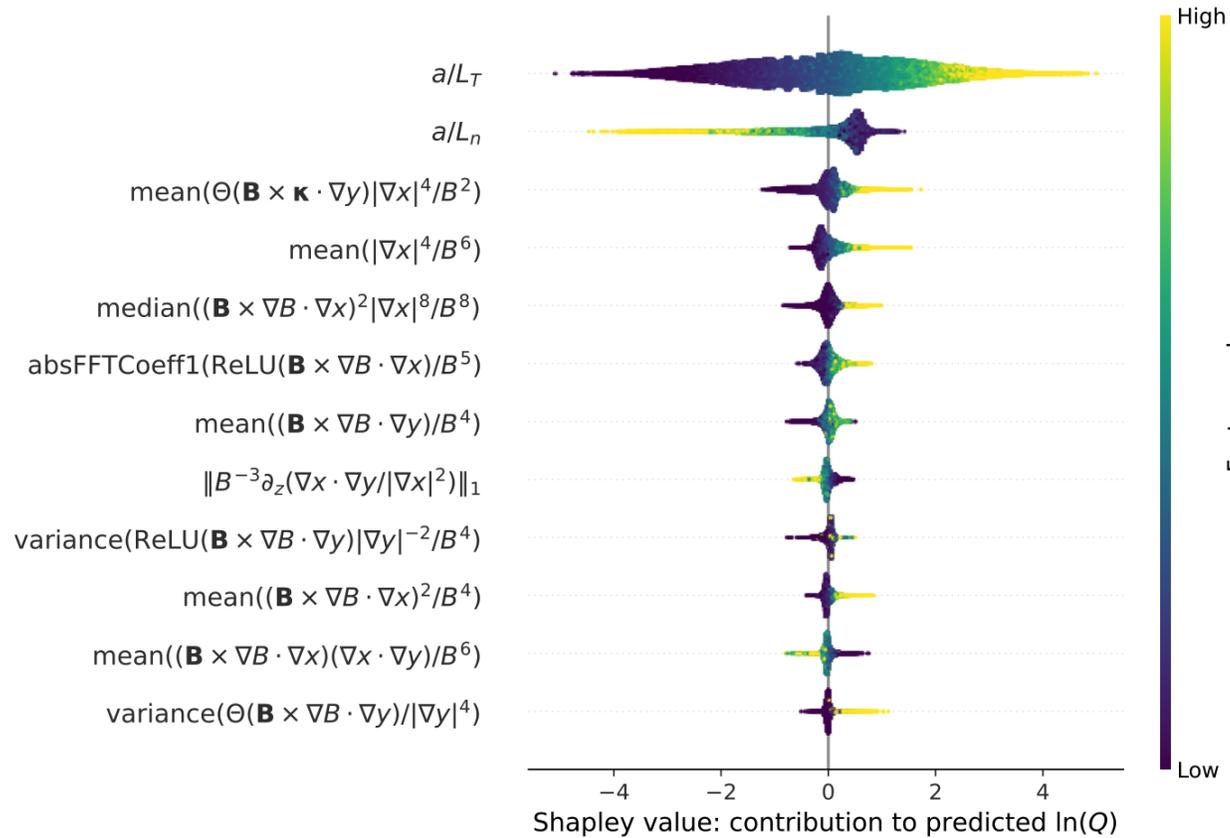


4 Input Components Selected by Shapley

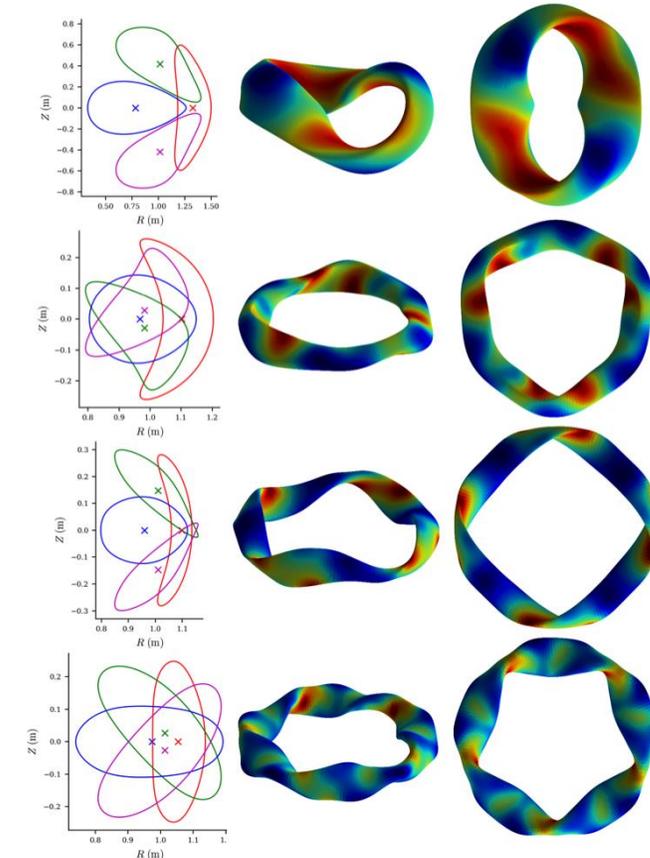


Similar Studies Provide Data-Driven Insight from Machine Learning on Stellarator Turbulent Transport Configurations

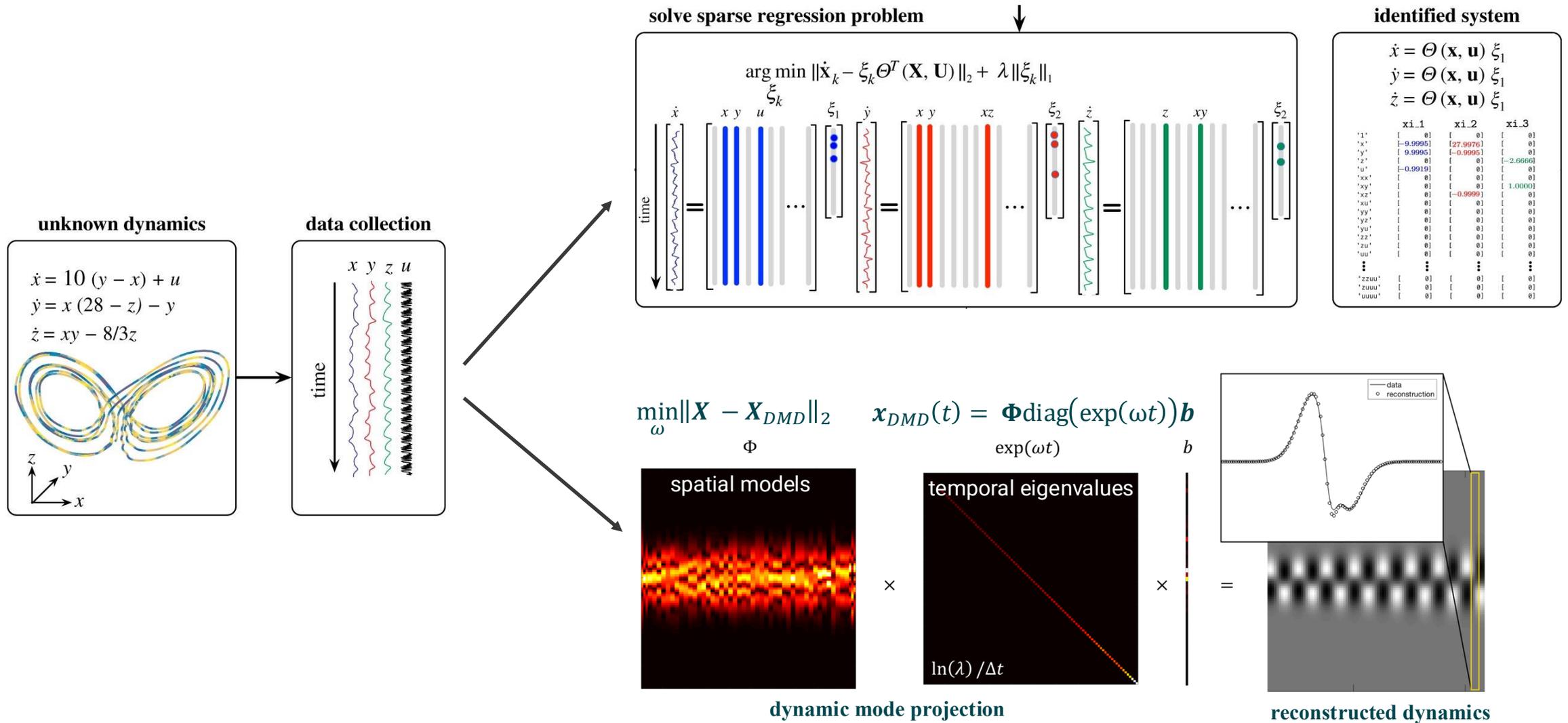
Shapley Value Parameter Contribution



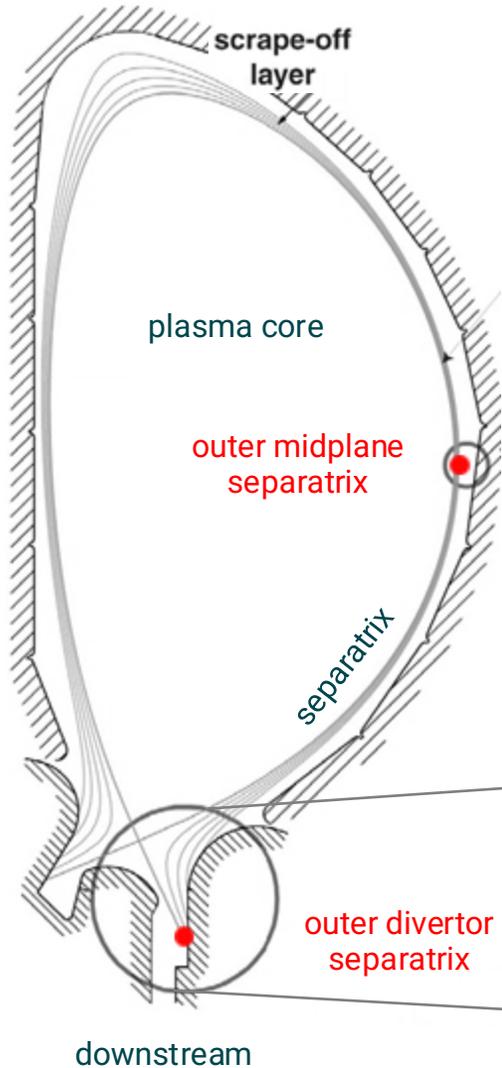
Quasi-Axisymmetric Configurations



Taking Interpretable Machine Learning One Step Further



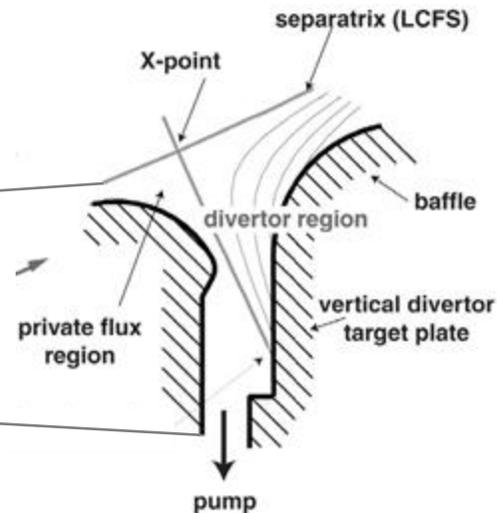
Operational Performance Can be Governed by Key Observable Quantities Linking Actuators to Divertor Targets



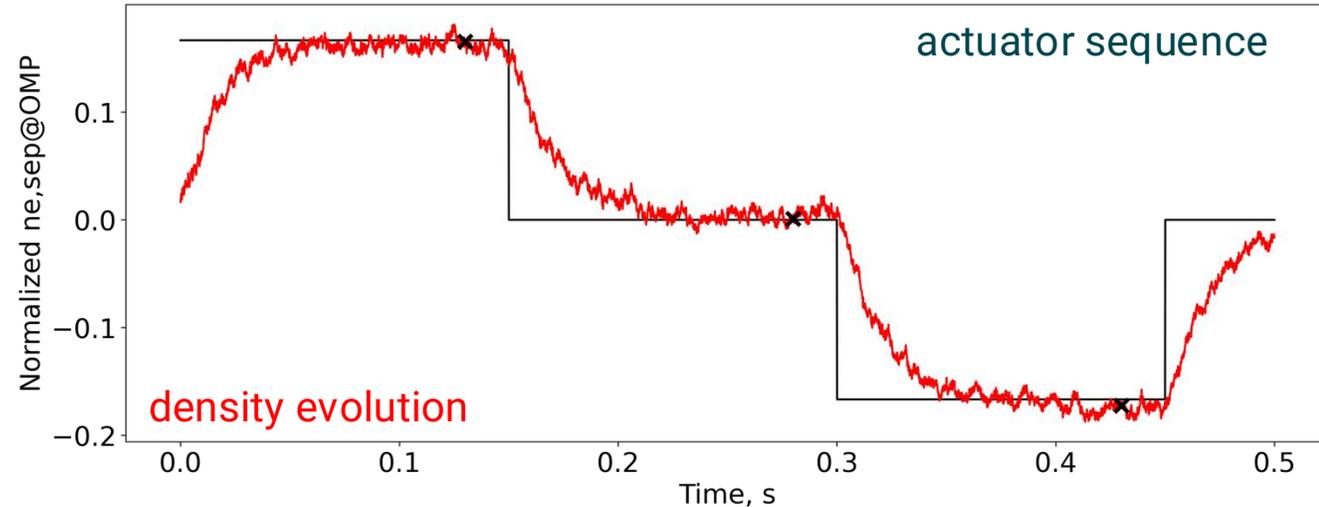
Two-Point Model

$$T_{\text{et}} = \left[\frac{2m_f}{e^3\gamma^2} \left(\frac{2\kappa_{0e}}{7} \right)^{4/7} \right] \left[\frac{q_{||u}^{10/7}}{n_u^2 L^{4/7}} \right] \times \left[\frac{(1 - f_{\text{cooling}})^2}{(1 - f_{\text{mom-loss}})^2} \right] \text{ (eV)},$$

upstream



Gas Puff Response

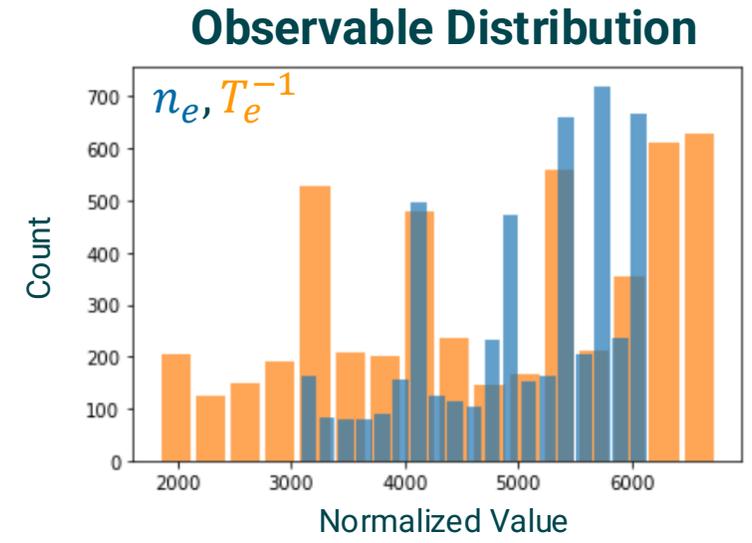
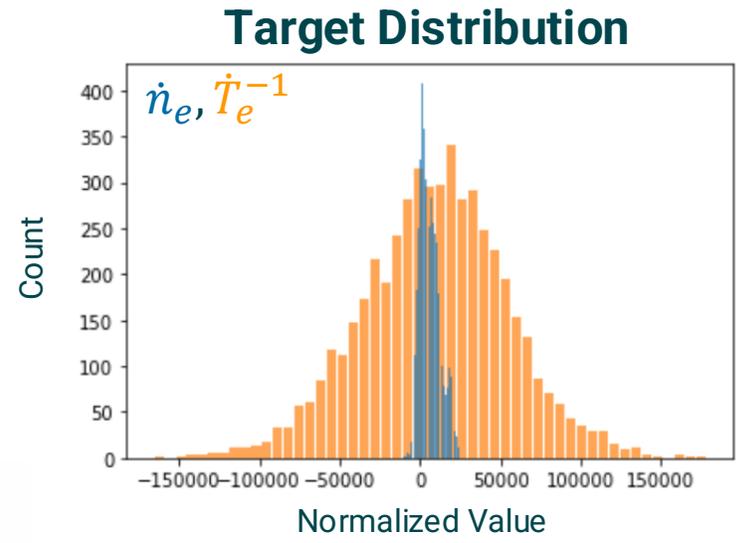
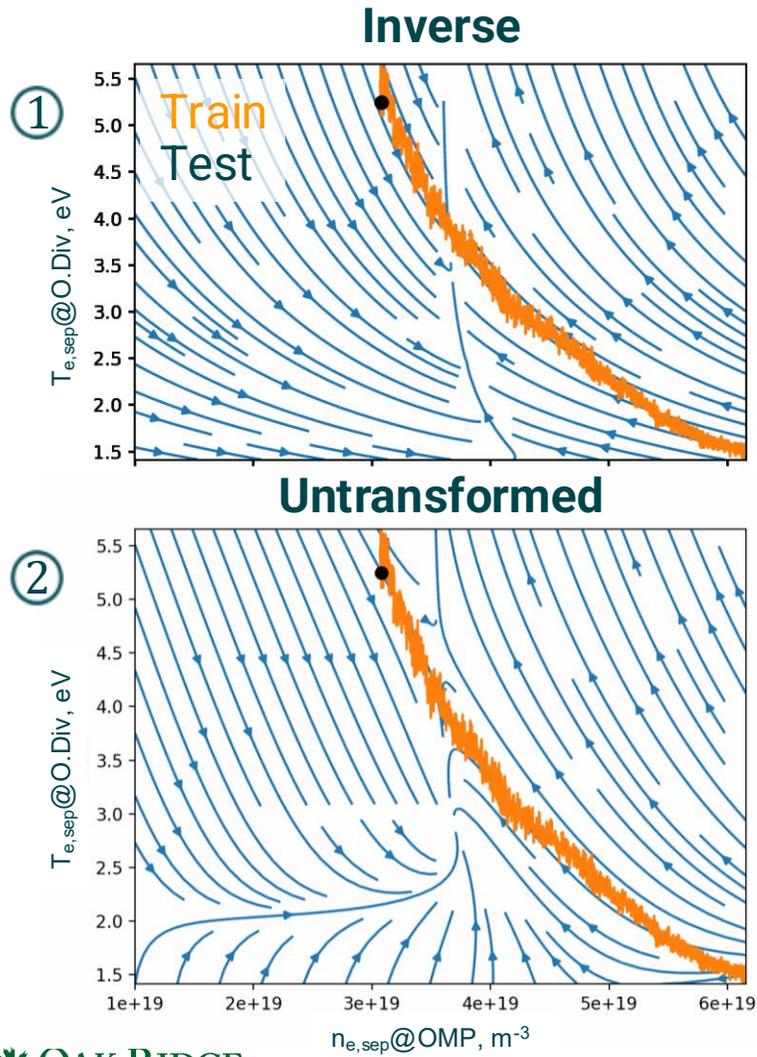


Separatrix values of n_e and T_e exhibit rise and fall times on the order of $\sim 60 - 70$ ms with step gas puff actuation

- Preliminary investigation of DIII-D shots shows timescales on the order of $\sim 50 - 100$ ms
- First-order Dead Time (FODT) linear fit shows minimal delay in simulated plasma response to direct actuation

□ De Pascuale et al., Phys. Plasmas, 2022
<https://doi.org/10.1063/5.0110393>

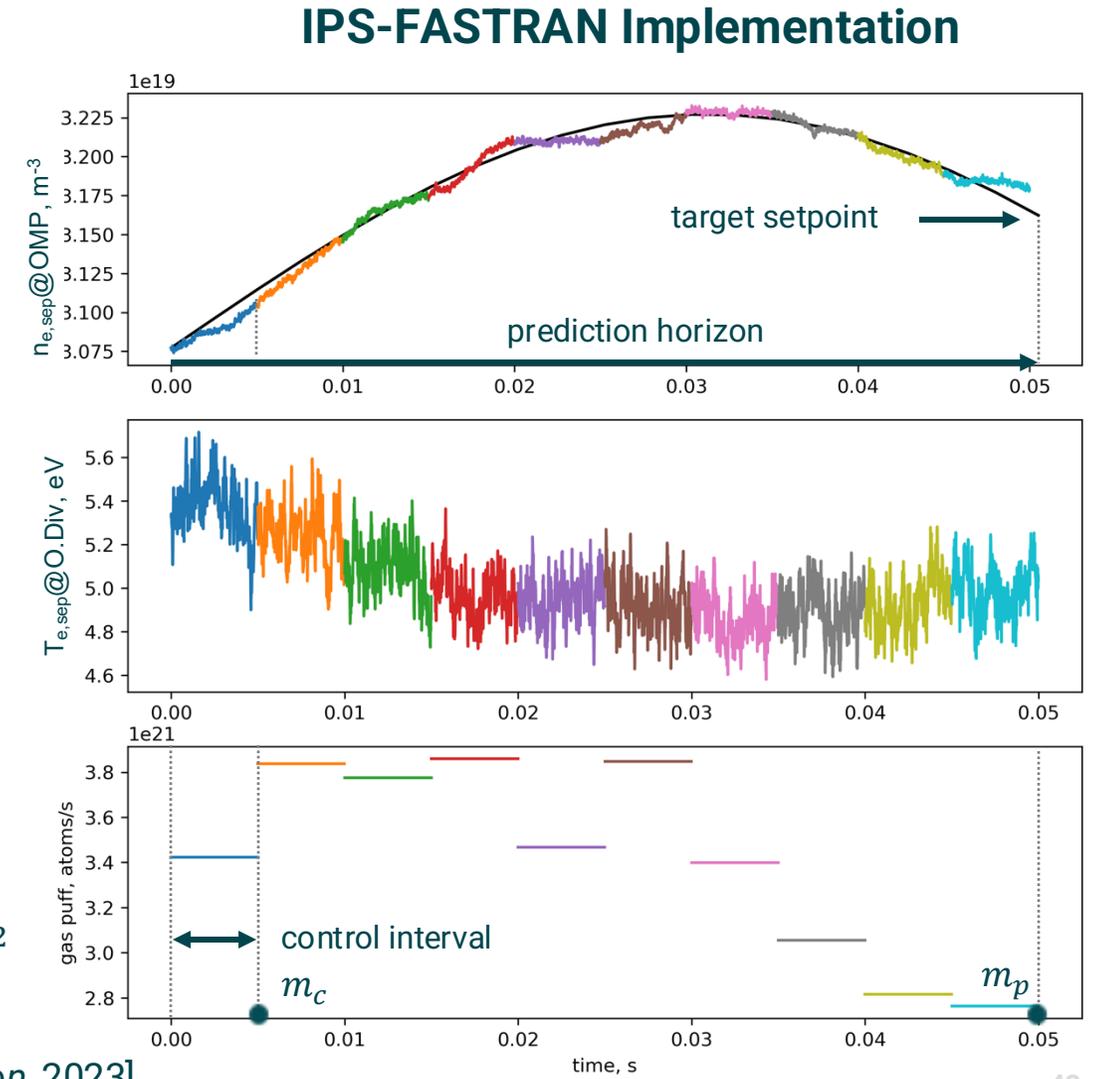
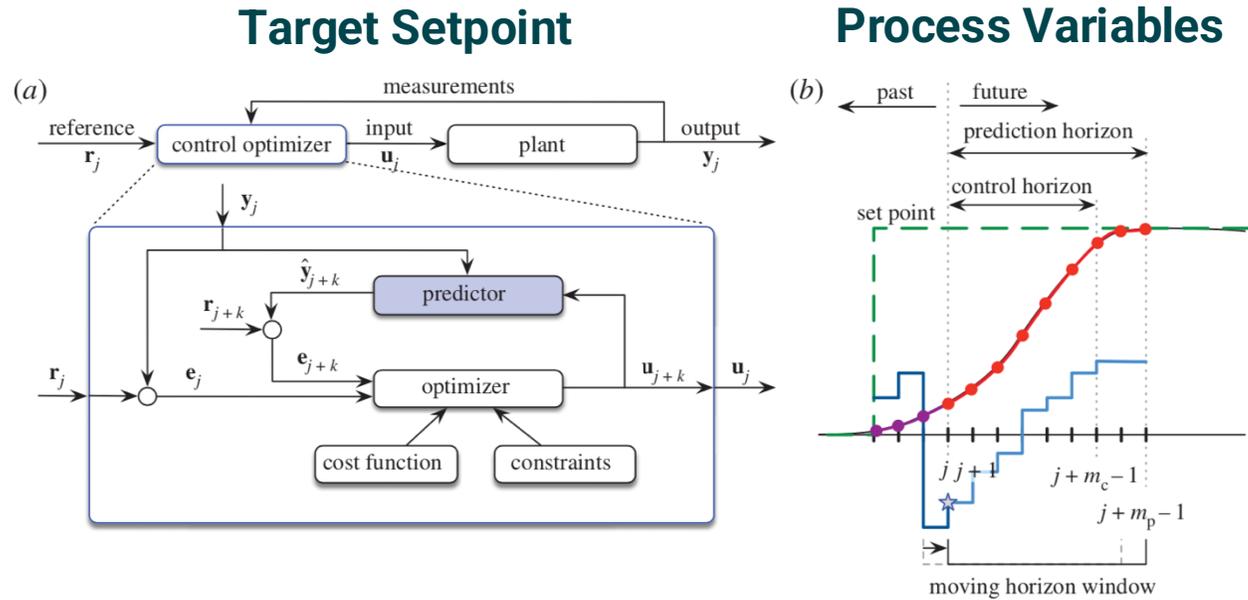
Dynamics are Best Characterized by Inverse Dependence of Key Plasma State Variables on Outer Divertor Target



- Data rescaling allows for physically consistent exploration of input parameter space for determination of controlled trajectories
- Identified coupled nonlinear equations are not sparse, but can be analyzed for stability characteristics to prevent over-fitting

□ Lore et al., Nucl. Fusion, 2023
<https://doi.org/10.1088/1741-4326/acbe0e>

Integration of Workflow Demonstrates Unique Capability for Evaluation of Strategies for Performance Optimization



- Model predictive control minimize cost function subject to constraints on input and output quantities
- IPS-FASTRAN provides framework for SOLPS-ITER runs with feedback from extracted reduced models

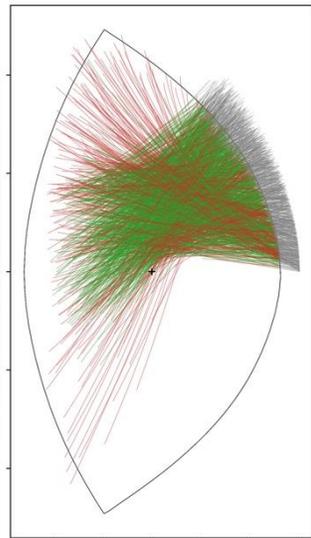
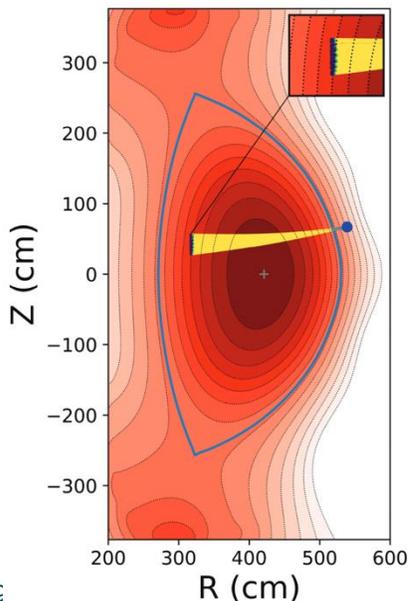
$$J = \left\| x_{j+m_p} - \tilde{x}_{m_p} \right\|^2 + \sum_{k=0}^{m_p-1} \left\| x_{j+k} - \tilde{x}_k \right\|^2 + \sum_{k=1}^{m_c-1} \left\| u_{j+k} \right\|^2 + \left\| \Delta u_{j+k} \right\|^2$$

Scientific Achievement

- Surrogate model accelerates FPP design and optimization of plasma profile control system utilizing electron cyclotron heating and current drive
- 25x reduction in compute time to generate electron cyclotron heating and current drive profile predictions based on 9 free parameters

Advance Tokamak

Ensemble Simulations

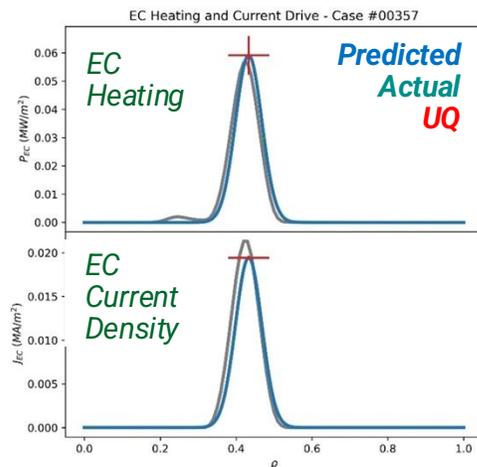


Significance and Impact

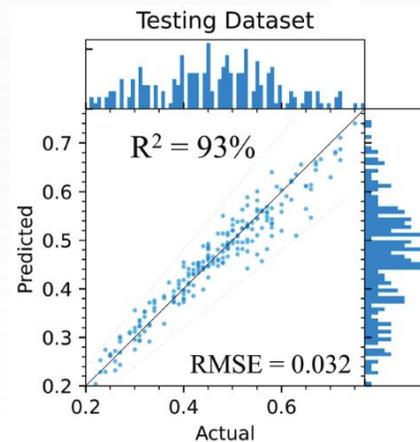
- Surrogate model of TORAY ray tracing code provides rapid predictions with uncertainty quantification
- Inverse-solvable model can predict inputs needed to generate a desired output without the need for large parameter scans for rapid optimization

□ Irvin et al., Fus. Sci. and Tech., 2025
<https://doi.org/10.1080/15361055.2025.2476829>

Predicted Profiles



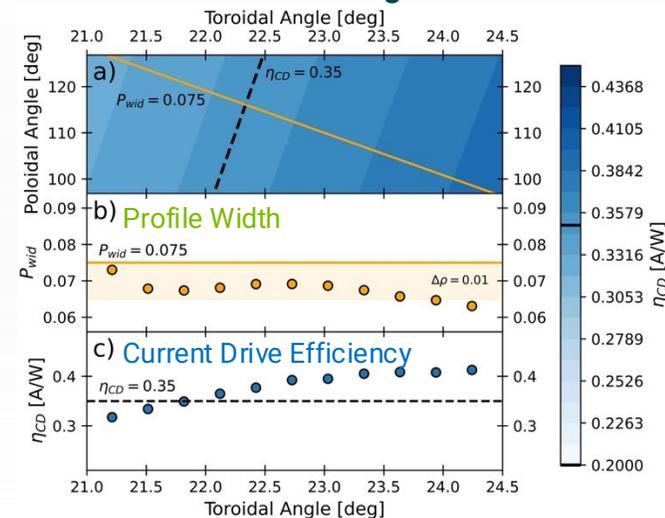
Parameter Regression



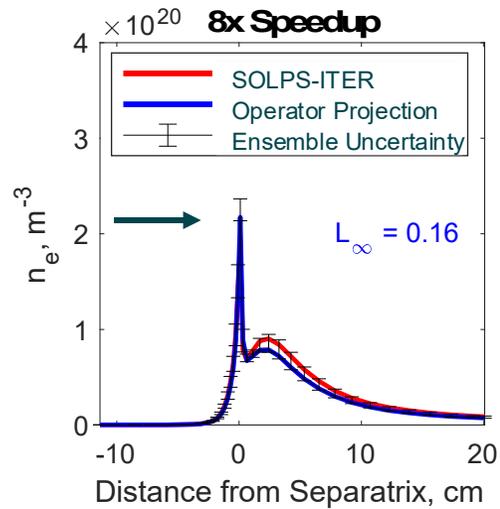
Research Details

- Former SULI/ECO internship (E. Hassan and S. De Pasquale) and current UT-Knoxville graduate student assistantship (L. Casali)
- Regression using database of 1600 points for a compact advanced tokamak (CAT) using IPS-TokDesigner

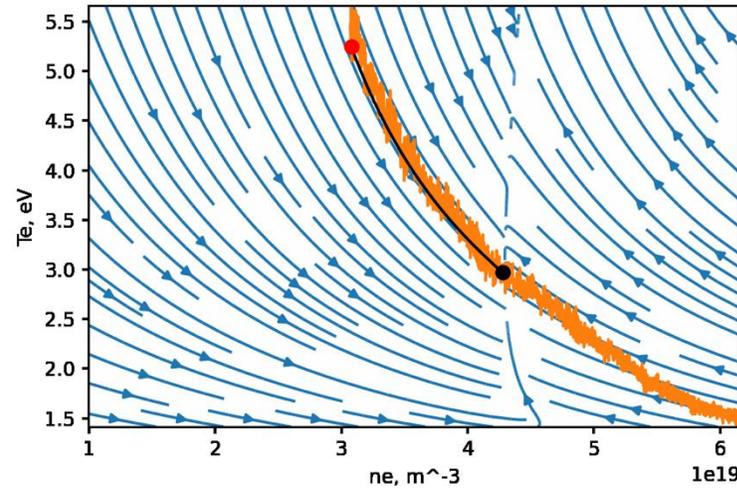
Inverse Modeling



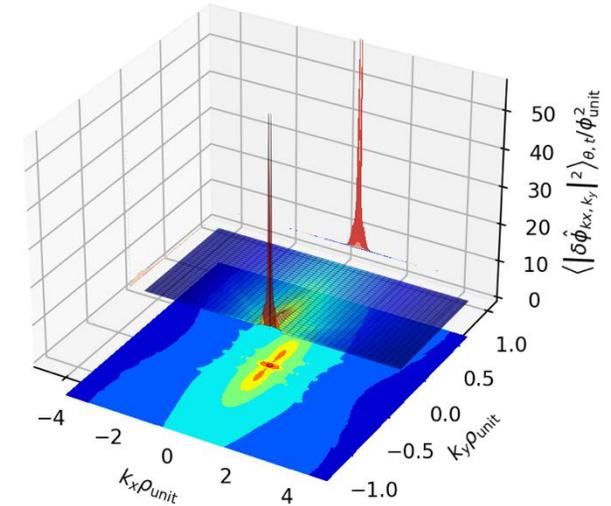
Three Takeway Themes



Reduced Modeling



Sparse Optimization



Machine Learning

$$x_{k+1} = Ax_k$$

$$x_{k+1} = Ax_k + Bu_k$$

$$\frac{dx}{dt} = \sum_{i=0}^3 \sum_{j=0}^3 a_{ij} x^i u^j$$

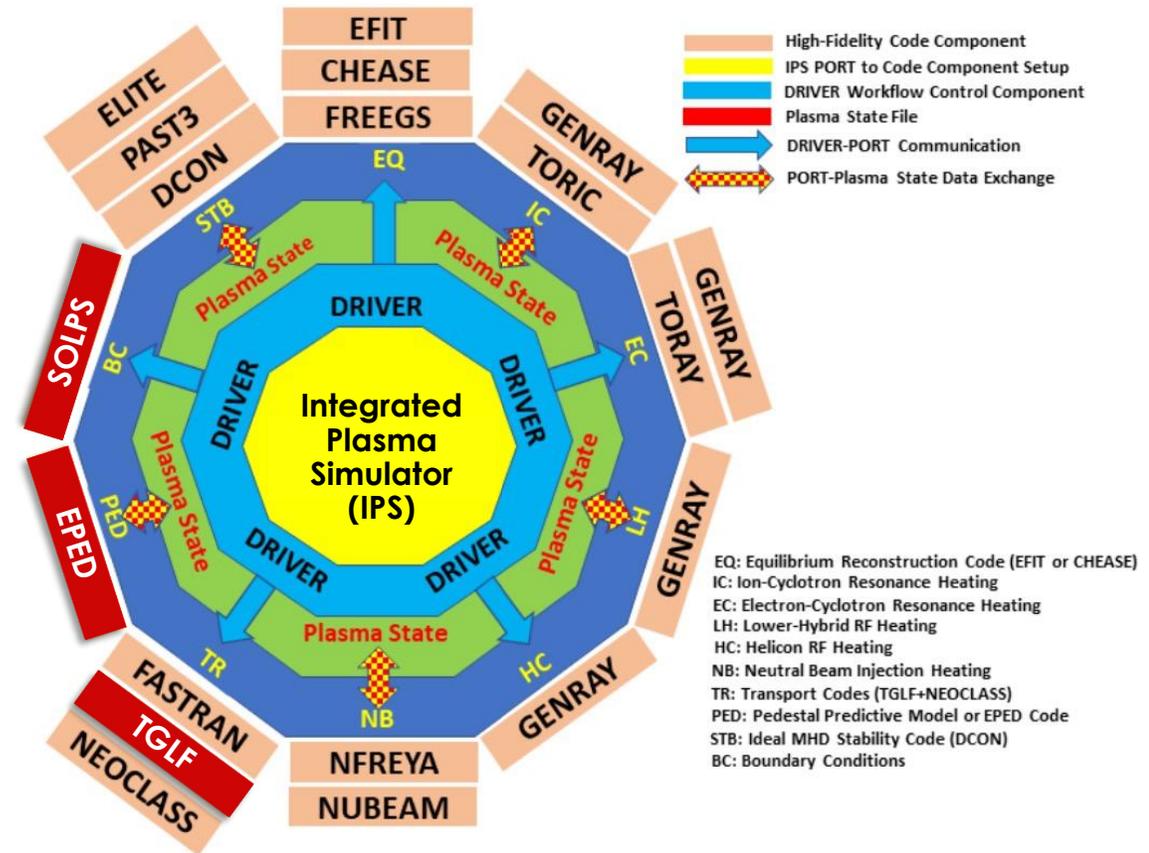
$$\frac{dy}{dt} = f\left(\sum_{i=1}^n x_i w_i + b_i\right)$$

The Integrated Plasma Simulator (IPS) Framework is Enabling Surrogate Model Development for Online Coupled Workflows

- **Accelerate** expensive simulation of core transport and pedestal stability within a coupled framework
- **Predict** plasma boundary response to core and pedestal conditions during detachment control
- **Quantify** uncertainty in machine learning surrogate models for decision-ready tokamak pulse simulation

These aims enable fast tokamak pulse simulation for:

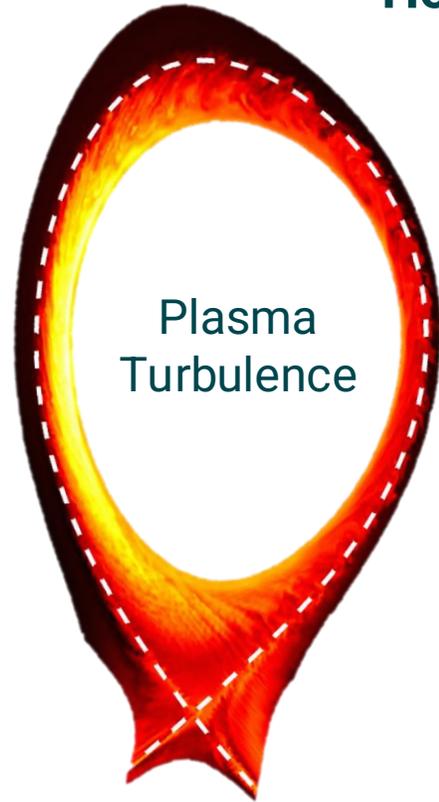
1. **Reactor design** leveraging physics informed and Bayesian inference neural networks to reduce dimensionality of coupled simulations
2. **Session planning** leveraging data assimilation and adaptive sampling to constrain predictions of operational space



Multi-Fidelity Model Coupling Framework

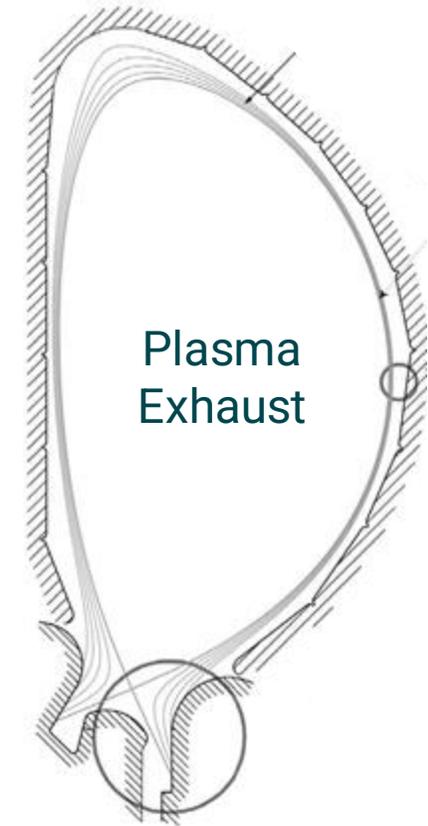
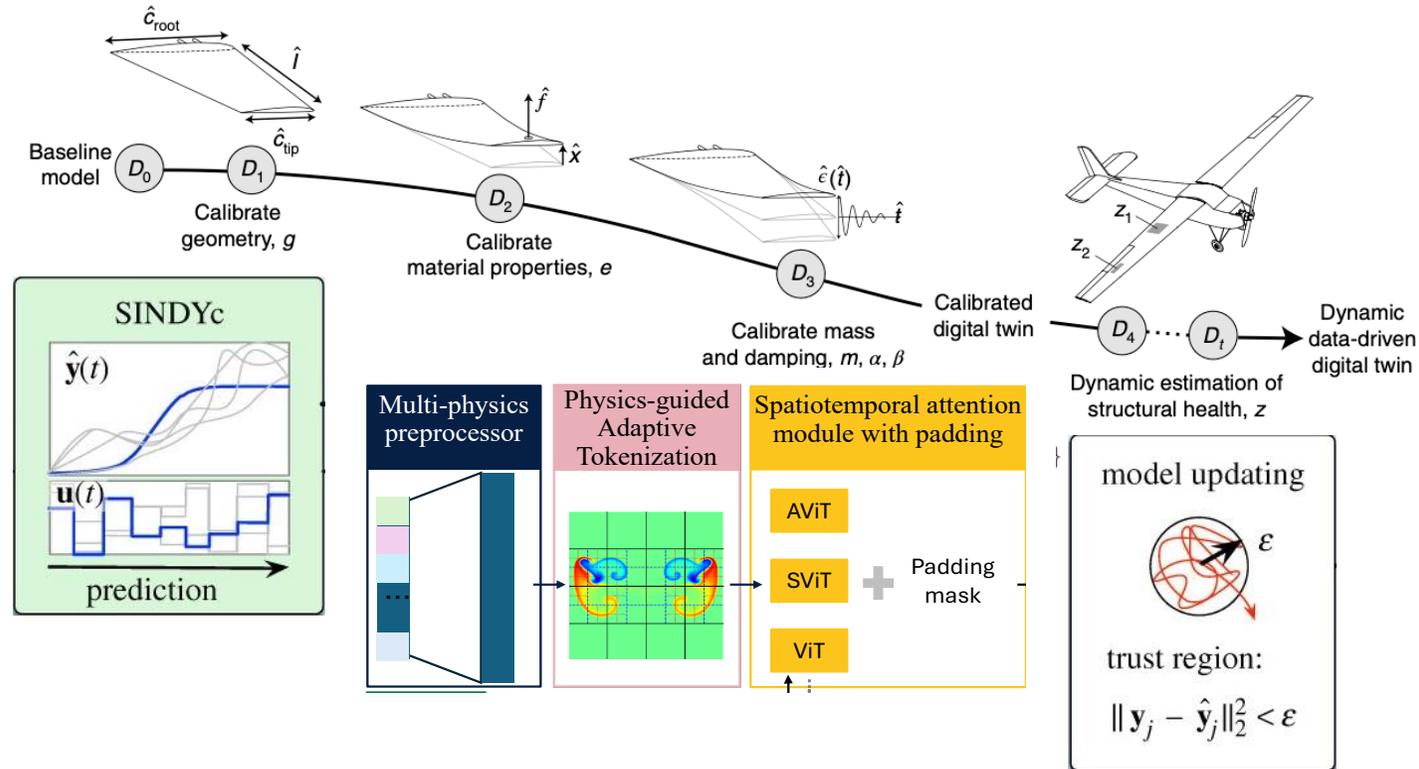
Code-Piloting Advanced Simulations for Reactor Digital Twins

How can we predict anomalous transport in the plasma boundary?



Advanced Profile Prediction for Fusion Pilot Plant Design

PI: D. Ernst at MIT (FIRE 2024 to 2028)



Trustworthy Multiscale Probabilistic Turbulence Foundation Models

PI: P. Zhang at ORNL (LDRD 2023 to 2026)

$$\Gamma_{\perp} \propto D_{\perp} \chi_{\perp} \left(\frac{\partial}{\partial r} \right) n, T$$

About You – Where will your science story go from here?

“Remember to look up at the stars and not down at your feet. Try to make sense of what you see and wonder about what makes the universe exist. Be curious” – Stephen Hawking



ITER Tokamak



W7-X Stellarator

“In science, it is not speed that is most important. It is the dedication, the commitment, the interest and the will to know something and understand it” – Eugene Wigner



OAK RIDGE

National Laboratory



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Multi-Parameter Mapping of Input Space to Targeted Output Quantities for Reduced Simulation Scans of TGLF Sensitivity

Variable	Definition
r/a	Normalized minor radius
R/a	Normalized major radius
κ	Elongation
$r \frac{\partial \kappa}{\partial r}$	Elongation shear
δ	Triangularity
$\frac{\partial R}{\partial r}$	Shafranov shift
q	Safety factor
$\frac{q^2 a^2}{r^2} \frac{\partial q}{\partial r}$	Safety factor shear
β_e	Kinetic to magnetic pressure ratio
ν_{ie}/ac_s	Collision frequency
T_i/T_e	Ion to electron temperature ratio
n_D/n_e	Deuterium to electron density ratio
n_C/n_e	Carbon to electron density ratio
Z_{eff}	Effective ion charge
a/L_{Te}	Electron temperature scale length
a/L_{Ti}	Ion temperature scale length
a/L_{ne}	Electron density scale length
a/L_{nD}	Deuterium density scale length
a/L_{nC}	Carbon density scale length
$\frac{qa^2}{r} \frac{\partial p}{\partial r}$	Total pressure gradient
$\text{sign}(I_p) R \omega_{\text{tor}} \frac{a}{c_s}$	Parallel velocity
$-\text{sign}(I_p) R \frac{\partial \omega_{\text{tor}}}{\partial r} \frac{a}{c_s}$	Parallel velocity gradient
$-\text{sign}(I_p) \frac{r}{q} \frac{\partial \frac{V_{E \times B}}{R}}{\partial r} \frac{a}{c_s}$	$E \times B$ velocity shear

24 Input Dimensions

Energy Flux 1 (6)

Energy Flux 2 (7)

Momentum Flux 2 (7+)

- VPAR_SHEAR_1
- BETAE
- RLNS_1/2
- RLTS_1/2

- VEXB+SHEAR
- KAPPA_LOC
- P_PRIME_LCO

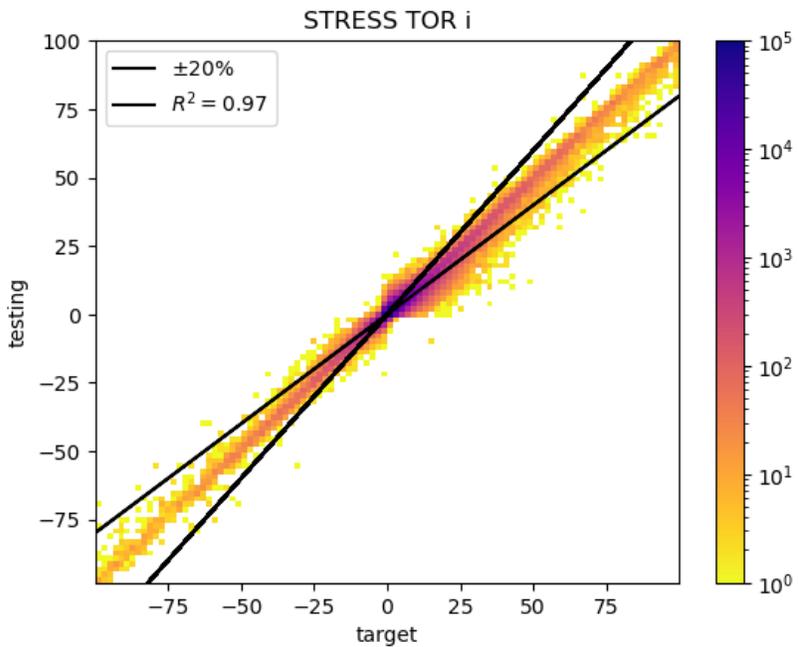
Particle Flux 1 (6)

Particle Flux 2 (7)

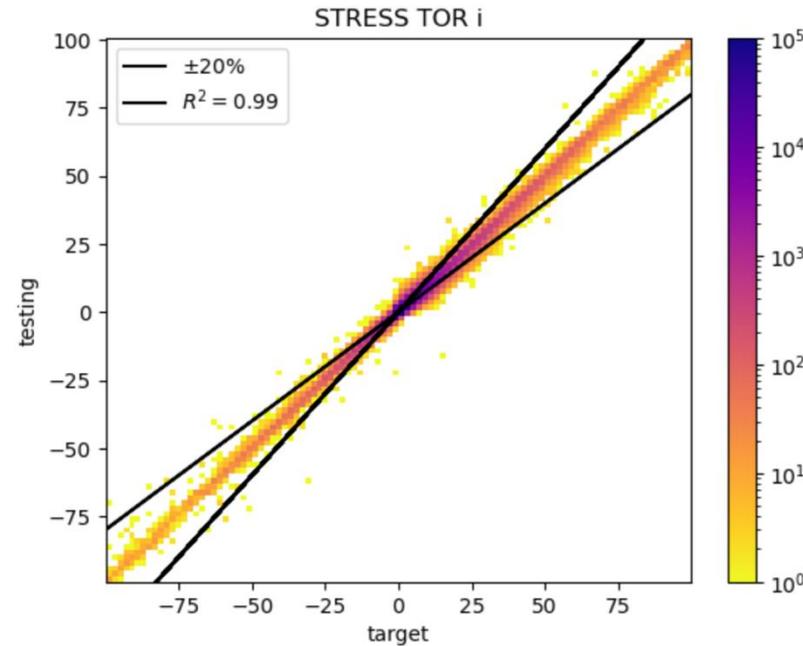
Particle Flux 3 (6)

- DRMAJDX_LOC
- VPAR_SHEAR_1
- RLNS_1/2/3
- AS_2/3

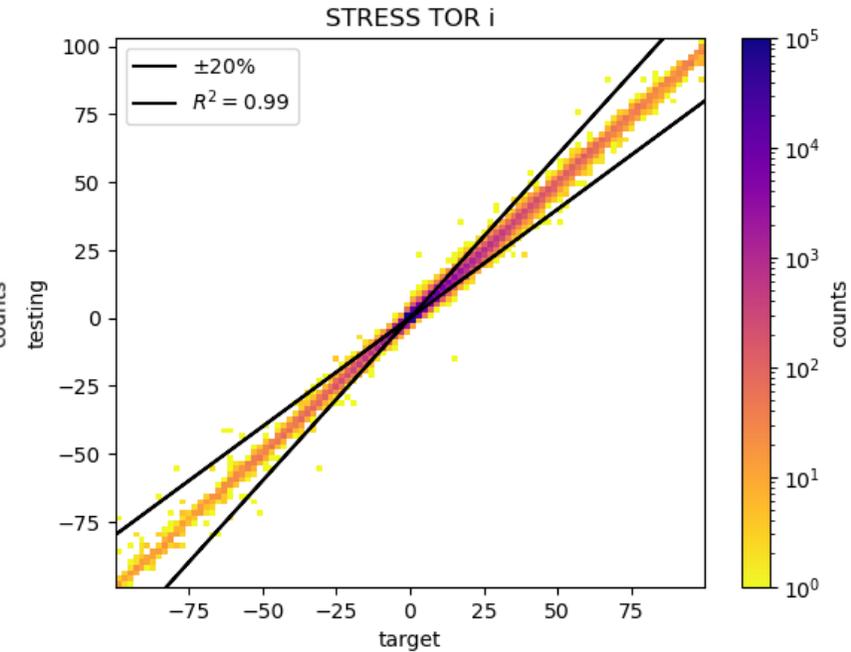
Particular Challenge in Capturing Momentum Flux Variability



7 Components



11 Components



17 Components

TGLF-NN-23D

Variable	Definition	Database	Output
r/a	Normalized minor radius	RMIN_LOC	OUT_tur_ENERGY_FLUX_1
R/a	Normalized major radius	RMAJ_LOC	<u>OUT_tur_ENERGY_FLUX_i</u>
κ	Elongation	KAPPA_LOC	OUT_tur_PARTICLE_FLUX_1
$r \frac{\partial \kappa}{\partial r}$	Elongation shear	S_KAPPA_LOC	<u>OUT_tur_STRESS_TOR_i</u>
δ	Triangularity	DELTA_LOC	OUT_tur_PARTICLE_FLUX_2
$\frac{\partial R}{\partial r}$	<u>Shafranov shift</u>	DRMAJDX_LOC	OUT_tur_PARTICLE_FLUX_3
q	Safety factor	Q_LOC	
$\frac{q^2 a^2}{r^2} \frac{\partial q}{\partial r}$	Safety factor shear	Q_PRIME_LOC	
β_e	Kinetic to magnetic pressure ratio	BETAE	
ν_{ie}/ac_s	Collision frequency	XNUE	
T_i/T_e	Ion to electron temperature ratio	TAUS_2	
n_D/n_e	Deuterium to electron density ratio	AS2	
n_C/n_e	Carbon to electron density ratio	AS3	

Z_{eff}	Effective ion charge	ZEFF	
a/L_{Te}	Electron temperature scale length	RLTS_1	
a/L_{Ti}	Ion temperature scale length	RLTS_2	
a/L_{ne}	Electron density scale length	RLNS_1	
a/L_{nD}	Deuterium density scale length	RLNS_2	
a/L_{nC}	Carbon density scale length	RLNS_3	
$\frac{qa^2}{rB_{unit}^2} \frac{\partial p}{\partial r}$	Total pressure gradient	P_PRIME_LOC	
$sign(I_p)R\omega_{tor} \frac{a}{c_s}$	Parallel velocity	VPAR_1	
$-sign(I_p)R \frac{\partial \omega_{tor}}{\partial r} \frac{a}{c_s}$	Parallel velocity gradient	VPAR_SHEAR_1	
$-sign(I_p) \frac{r}{q} \frac{\partial \frac{V_E \times B}{R}}{\partial r} \frac{a}{c_s}$	E x B velocity shear	VEXB_SHEAR	
λ_D	Debye length/gyroradius	DEBEYE	