

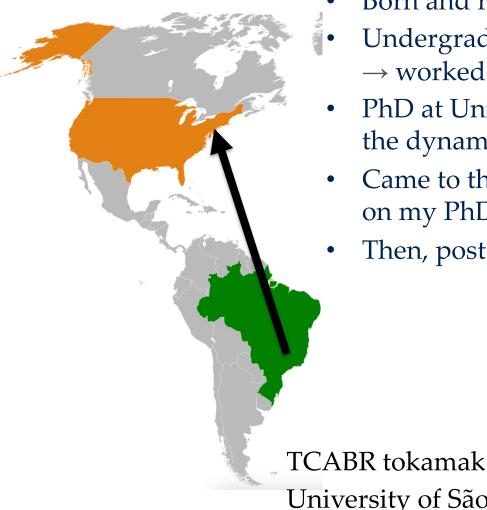
## Introduction to Plasma Physics

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SULI Introductory Plasma Physics Course June 11<sup>th</sup> 2024

#### A bit on my personal trajectory



- Born and raised in a small town in Brazil
- Undergrad and MSc at University of Campinas
   → worked on theory of current drive in plasmas
- PhD at University of São Paulo → worked on the dynamics of Alfvén waves in tokamaks
- Came to the US in 2014 on a fellowship to work on my PhD thesis topic at PPPL
- Then, postdoc and staff scientist at PPPL



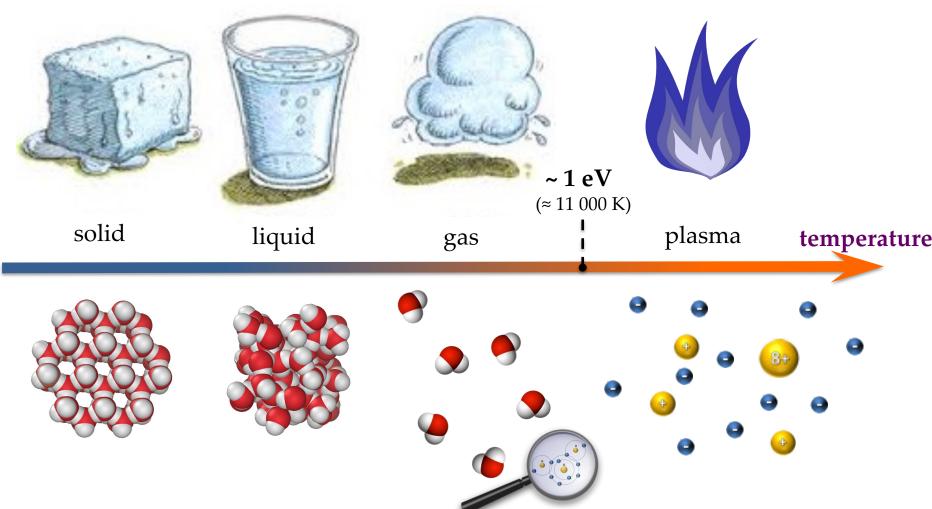
## Outline

- What is plasma?
  Occurrence and applications
  Criteria for plasmas
- A few key concepts
  - Debye length
  - Plasma oscillations
  - Gyrofrequency
  - Alfvén waves
- Summary

#### Outline

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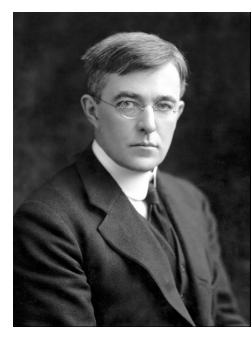
#### Matter exists in distinct forms



- Over 99% of the known Universe is in plasma form
- Modern telescopes suggest that the Universe is comprised by ~4.6% baryonic (ordinary) matter, ~26.8% dark matter, and ~68.3% dark energy

#### What is a plasma?

- In Ancient Greek, πλάσμα (plásma): 'moldable substance'
- The term "plasma" for an ionized gas was coined in 1927 by Irving Langmuir, because how an electrified fluid carried ions and electrons reminded him of how blood plasma carried red and white corpuscles.



Irving Langmuir (1881-1957);

Chemistry Nobel Prize 1932

## Definitions of plasmas

"Plasma is in some sense the natural, untamed state of matter..."

-Hazeltine and Waelbroeck, *The Framework of Plasma Physics* 

*"physical systems whose intrinsic properties are governed by collective interactions of large ensembles of free charged particles."* 

-NSF Basic Plasma Science and Engineering Website

A more formal definition will be given towards the end of this lecture

#### Star Birth - Eagle Nebula



Color pattern corresponds mostly to emissions from singly-ionized sulfur atoms (red), hydrogen (green) and doubly-ionized oxygen atoms (blue).

#### Aurora



Disturbances in the upper atmosphere caused by the solar wind (e.g., due to coronal mass ejections) lead to ionization and of atmospheric constituents that emit light of varying color and complexity <sup>9</sup>

#### Matter around black holes



As matter is drawn to a black hole, and its immense gravitational influence creates turbulent and violent conditions, heating gas and stripping electrons away from its constituent atoms. 10

#### Technological applications

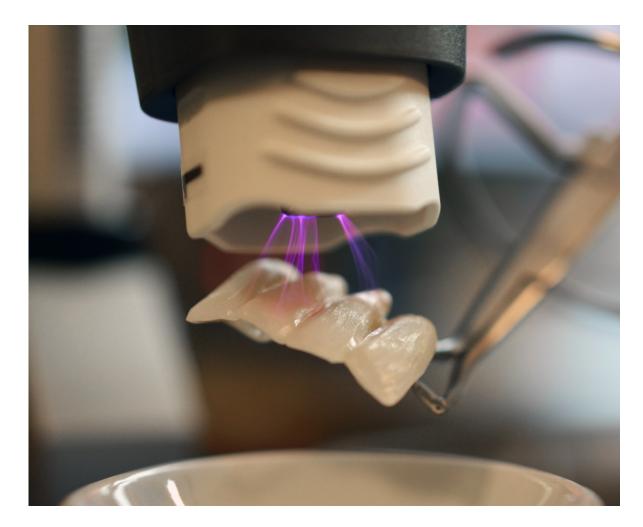
- Plasma pencil
- Plasma torch
- Plasma TV
- Fluorescent lamp
- Plasma thrusters for space travel
- Controlled thermonuclear fusion
- • •

#### Plasma torch



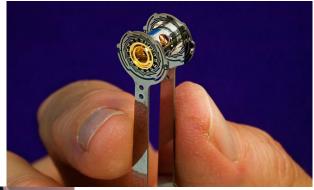
Useful in many applications, such as metal cutting, welding and waste disposal

#### Plasma pencil



Used to treat and sterilize irregular surfaces, making them appropriate for decontaminating dental cavities without drilling

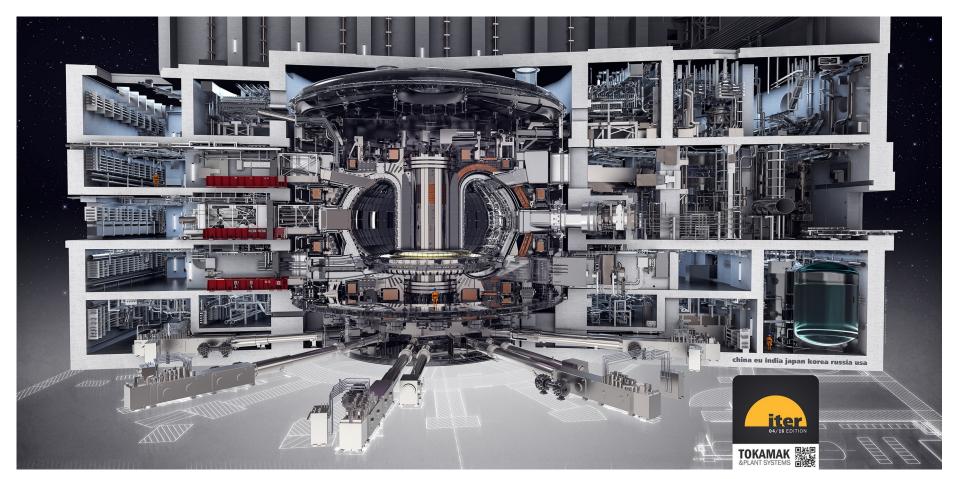
#### National Ignition Facility (NIF)





#### Inertial fusion

#### International Thermonuclear Experimental Reactor (ITER)



#### Magnetic fusion

#### Plasmas occur within a wide range of parameters

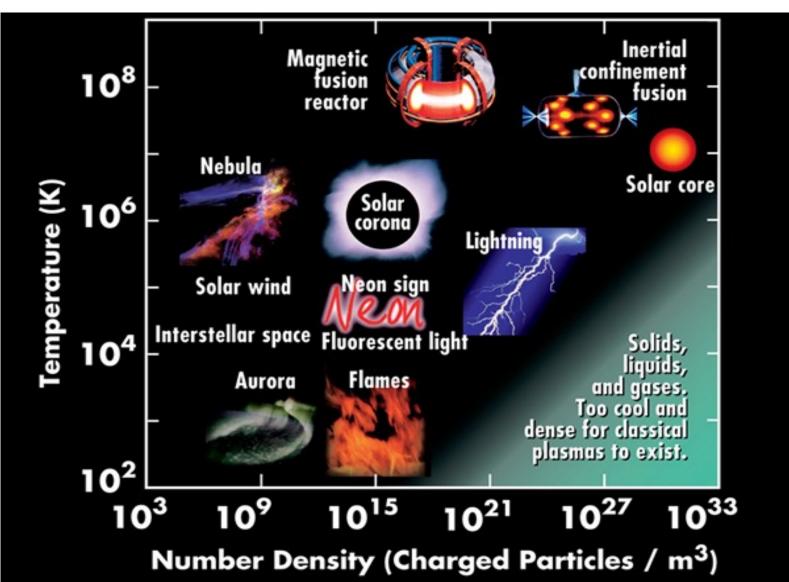


Image credit: National Ignition Facility, Lawrence Livermore Nat. Lab.

## Outline

# What is plasma? Occurrence and applications Criteria for plasmas

#### A few key concepts

- Debye length
- Plasma oscillations
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#### Studying plasmas involves many disciplines

- Electrodynamics
- Fluid mechanics
- Statistical physics
- Thermodynamics
- Quantum mechanics
- •

The electric and gravitational forces exerted on  $m_1$  by  $m_2$  are:

$$m_1 \vec{a} = \Sigma \vec{F} = \vec{F}_G + \vec{F}_E = \left[ -\frac{Gm_1m_2}{r_{1,2}^2} + \frac{q_1q_2}{4\pi\epsilon_0 r_{1,2}^2} \right] \vec{r}$$

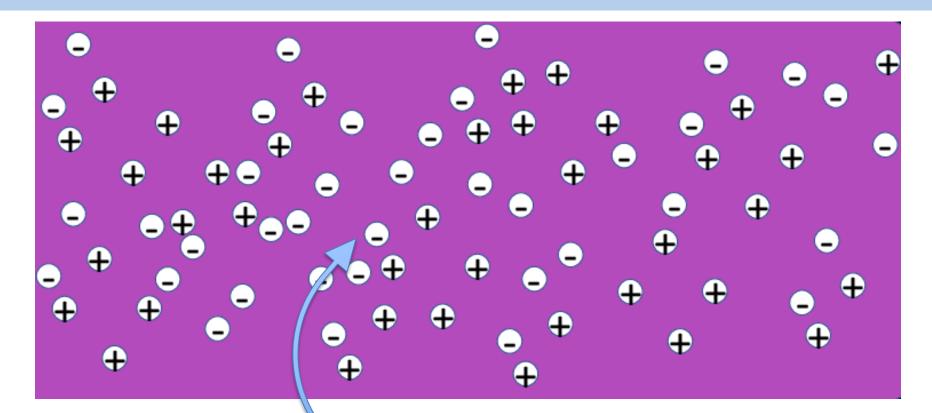
Assuming one is an ionized deuterium atom and the other is an electron:

$$\frac{F_E}{F_G} = 1.1 \times 10^{39}$$

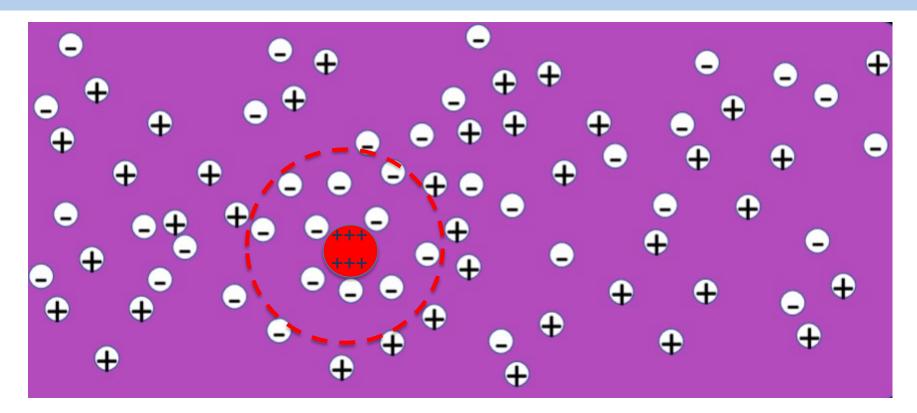
Gravity is irrelevant for lab plasmas (but not for astrophysical ones)

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Place a test charge into a quasi-neutral plasma



Key question: What is the radius of the sphere of influence of this extra charge? How far away do you have to be for the extra charge to be completely "shielded" by the plasma?

– Poisson's Equation

$$\nabla^2 \Phi = -\frac{\rho_q}{\epsilon_0}$$

$$\nabla^2 \Phi = -Ze\delta(\mathbf{x}) - \frac{e}{\epsilon_0}(n_0 - n_e)$$
$$n_e = n_0 e^{-\frac{e\Phi}{kT_e}}$$

$$\nabla^2 \Phi = -Ze\delta(\mathbf{x}) - \frac{en_0}{\epsilon_0}(1 - e^{-\frac{e\Phi}{kT_e}})$$

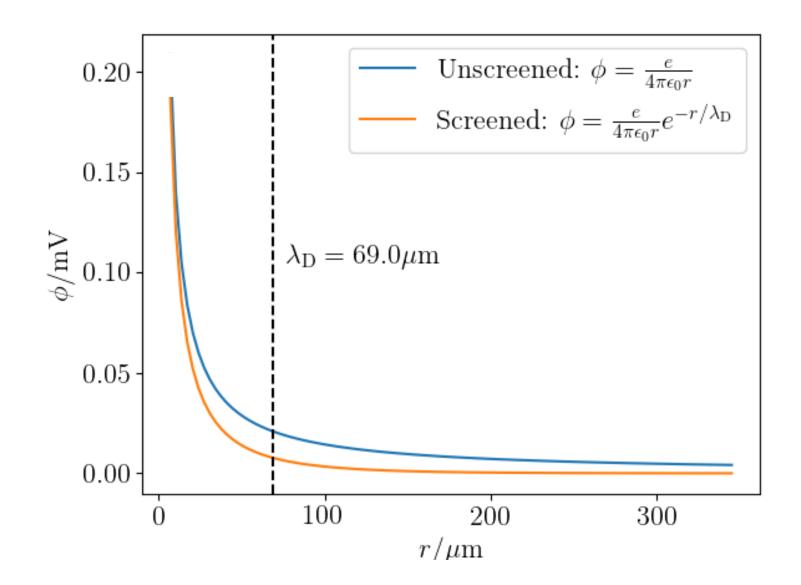
• Debye length Use  $e\phi \ll kT$  to linearize equation

$$\nabla^2 \Phi \approx -Ze\delta(\mathbf{x}) - \frac{en_0}{\epsilon_0} \left(1 - \left(1 - \frac{e\Phi}{kT_e}\right)\right)$$

$$\nabla^2 \Phi - \frac{1}{\lambda_D^2} \Phi \approx -Ze\delta(\mathbf{x}) \qquad \frac{1}{\lambda_D^2} = \frac{n_0 e^2}{\epsilon_0 k T_e}$$

$$\Phi(r) = \frac{Ze}{4\pi\epsilon_0 r} e^{\frac{-r}{\lambda_D}}$$

Typical Debye lengths: Solar core: 10<sup>-11</sup>m, Tokamak: 10<sup>-4</sup>m, Intergalactic medium: 10<sup>+5</sup>m <sup>24</sup>

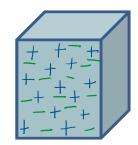


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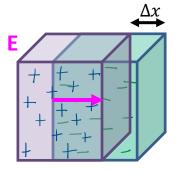
#### Plasma oscillations

• Plasma frequency



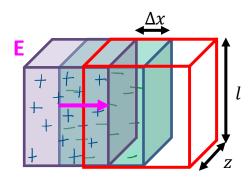
#### Plasma oscillations

- Plasma frequency
- Use Gauss' Law to find E
- Apply Newton's 2<sup>nd</sup> Law to find equation of motion



#### Plasma oscillations

• Plasma frequency



$$\int \vec{E} \cdot d\vec{A} = Q_{enc}/\epsilon_0 \qquad ma = F$$

$$Elz = en_e (\Delta x \ l \ z)/\epsilon_0 \qquad m_e \frac{d^2 \Delta x}{dt^2} = -eE$$

$$E = en_e \Delta x/\epsilon_0 \qquad \vec{x} = -\frac{n_e e^2}{m_e \epsilon_0} x$$

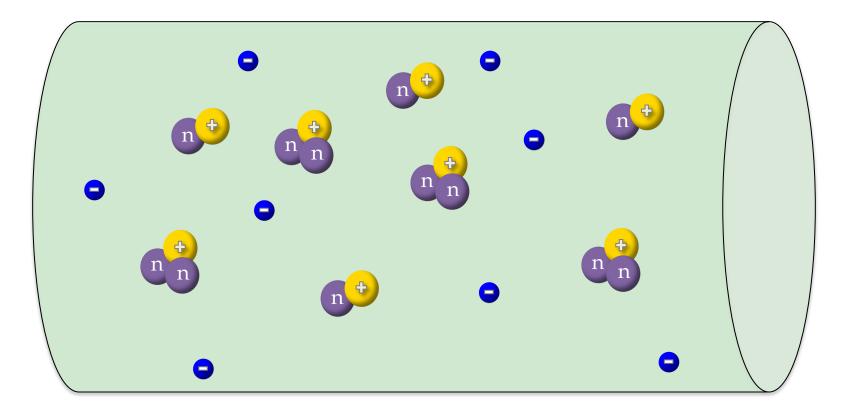
Compare with Hooke's Law  $\ddot{x} = -\omega^2 x$ 

$$\omega_{ps}^2 = \frac{n_s e^2}{\epsilon_0 m_s}$$

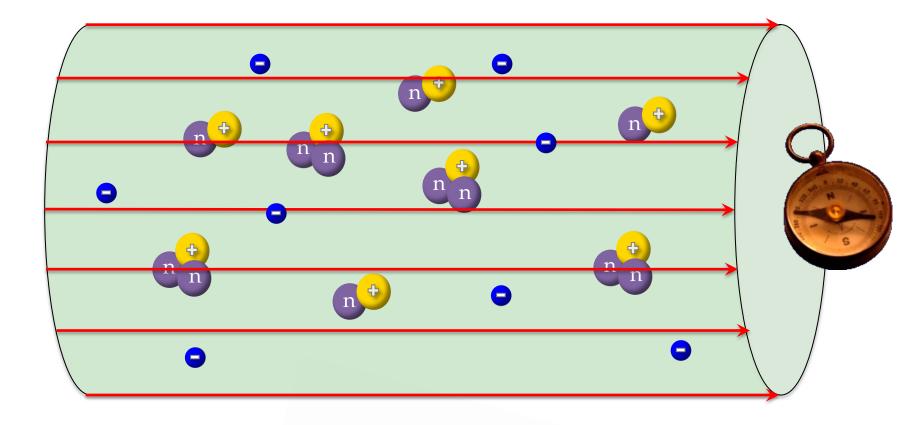
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#### How to confine a plasma?



#### Magnetic fields confine plasmas





#### Cyclotron frequency

$$F = ma$$

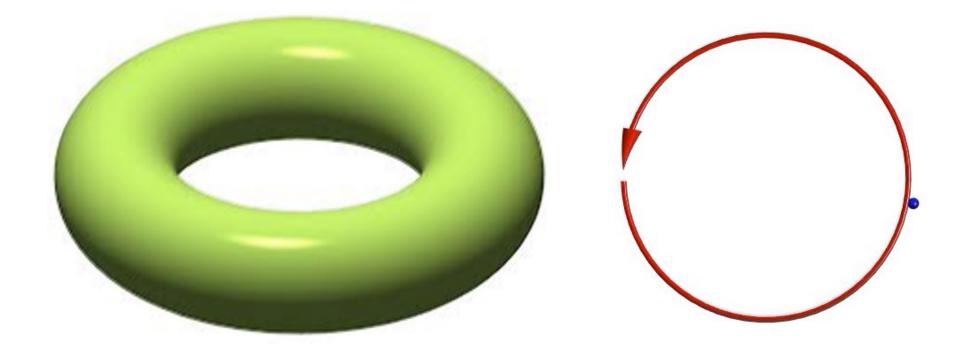
$$q(\vec{v} \times \vec{B}) = -mv^{2}/r$$

$$qvB = mv^{2}/r$$

$$v = qBr/m$$

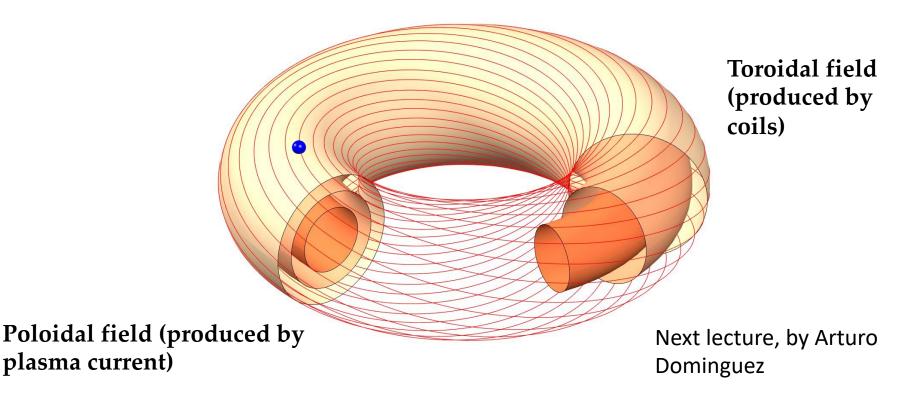
$$\omega = \frac{v}{r} = \frac{qB}{m}$$

## A torus is the simplest configuration needed to confine plasmas

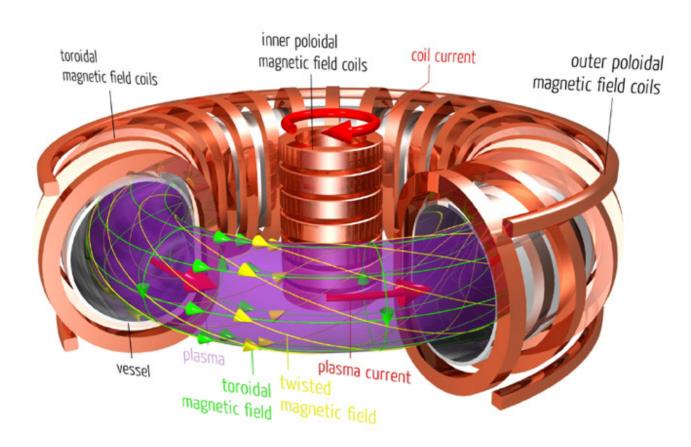


The picture is not simple: particle drifts are ubiquitous  $\rightarrow$  a combination of poloidal and toroidal fields is necessary [see next lecture by Arturo Dominguez]

## A torus is the simplest configuration needed to confine plasmas

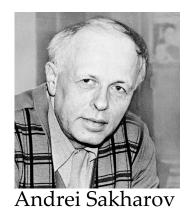


#### Tokamak: toroidal chamber with magnetic coils

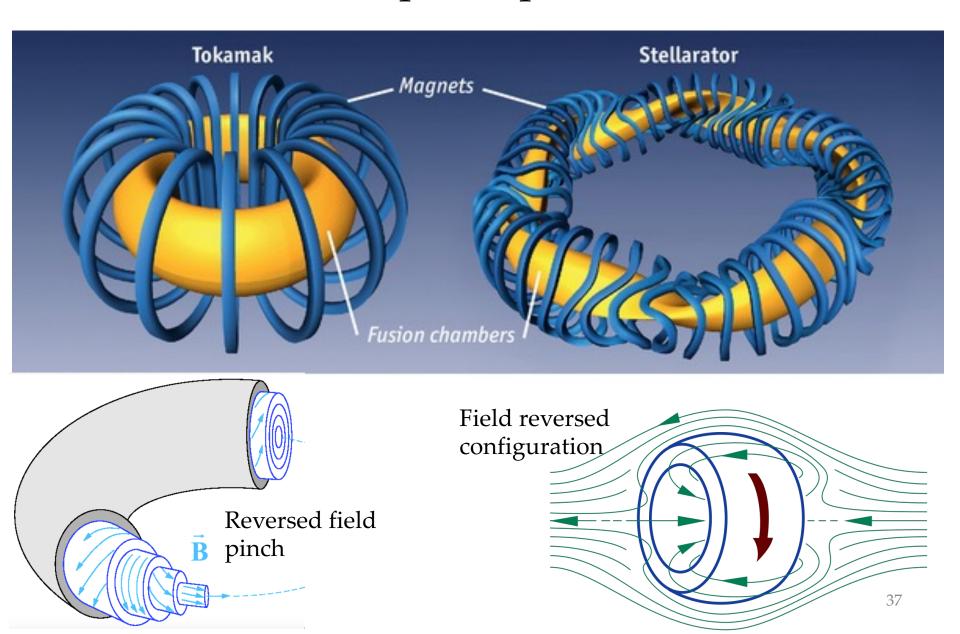




Igor Tamm



#### Alternative concepts of plasma confinement

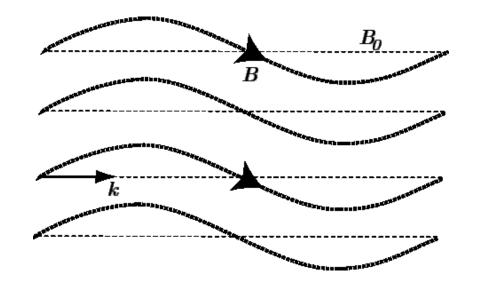


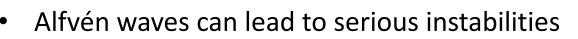
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#### Alfvén waves: a fundamental mode of oscillation in plasmas embedded in a magnetic field

Alfvén waves result from the coupling between fluid dynamics and electromagnetism  $\rightarrow$  birth of magnetohydrodynamics





• They might explain the solar corona heating



Hannes Alfvén (1908-1995) 1970 Physics Nobel Prize

H. Alfvén, *Nature* **150**, 405 (1942) <sup>39</sup>

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#### Properties of a plasma

1. Conducting medium, with many degrees of freedom

- 2. Shields electric fields
- 3. Supports many waves:
  - vacuum waves, such as light waves
  - gas waves, such as sound waves
  - a huge variety of new waves, based on electromagnetic coupling of constituent charged particles, and based on a variety of driving electric and magnetic fields

Definitions of plasmas (in the beginning of this lecture)

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A more formal definition will be given towards the end of this lecture

# A more formal definition of a plasma

1. Debye length << system characteristic length

2. Large number of particles in a Debye sphere

3. Plasma oscillation period << time between collisions

Plasmas are physical systems whose intrinsic properties are governed by collective interactions of large ensembles of free charged particles

## Take-aways

- Plasma phenomena appears in a variety of applications: (astrophysics, solar physics, plasma devices, nuclear fusion)
- Controlled fusion can be inertial or magnetic
- Basic time scale of plasma is the plasma oscillation period
- Basic space scale of plasma is the Debye shielding length

#### Further reading

## Introduction to Plasma Physics, F. F. Chen

Fundamentals of Plasma Physics,
 J. A. Bittencourt

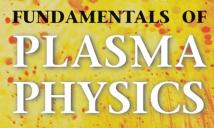
Francis F. Chen

Introduction to Plasma Physics and Controlled Fusion

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