



Extreme Light-Matter Interactions Laboratory

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Relativistic Laser Plasmas

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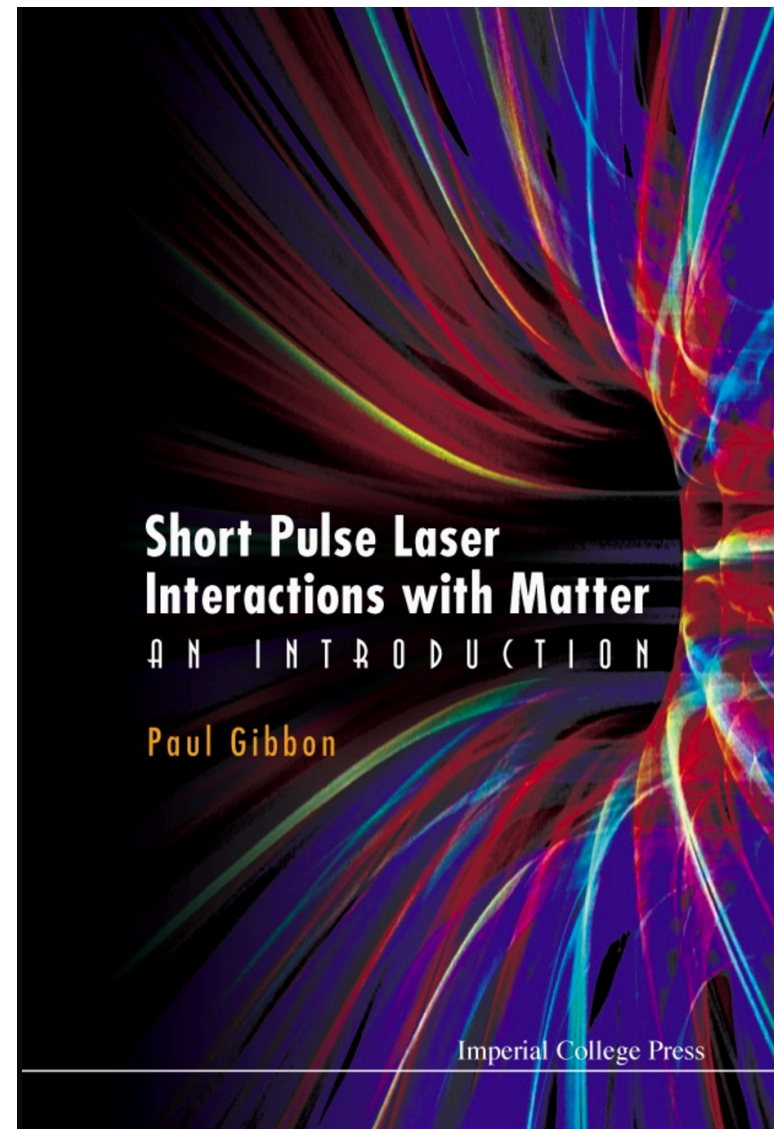
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PPPL, SULI lecture, June 15, 2017

Textbook

Paul Gibbon,
"Short Pulse Laser Interactions
with Matter", Imperial College
Press, London U.K. (2005)



[http://www.fz-juelich.de/ias/jsc/EN/AboutUs/Organisation/
ComputationalScience/Simlabs/slpp/Teaching/SPLIM/_node.html](http://www.fz-juelich.de/ias/jsc/EN/AboutUs/Organisation/ComputationalScience/Simlabs/slpp/Teaching/SPLIM/_node.html)

High-field science on a tabletop

Fundamental questions of high-field science and astrophysics

- How does matter behave in extreme fields?
- What is the physics of stellar atmospheres, relativistic astrophysical jets, supernovae, etc?
- Can we observe the effects of radiation reaction and nonlinear QED in plasma dynamics?
- Can we generate electron-positron pairs from the Dirac Sea in vacuum by intense photon interactions?

Applications

- Can we shrink kilometer-scale accelerator/synchrotron facilities to fit on a laboratory tabletop?
- Ultrafast sources of intense light and high-energy particles
 - Ultra-intense laser light (Raman/Brillouin amplification in plasmas)
 - Attosecond (10^{-18} s) x-ray pulses for time-resolved studies
 - High-energy (GeV) electron beams (Laser Wakefield Acceleration)
 - High-energy ion/proton beams (Solid foils)
- Laser Fusion

Laser fields

Lasers are the most intense sources of electromagnetic radiation available in laboratory conditions for high-field science experiments.

Power

$$P = \frac{\text{energy}}{\text{time}}$$

Irradiance or intensity

$$I = \frac{\text{energy}}{\text{time} \times \text{area}}$$

$$I_L = \frac{1}{2} \epsilon_0 c E_0^2$$

- SI units of I are W/m^2 , but W/cm^2 is used in laser science.
- Intensity defines the fields and e/m force.

$$I = 10^{20} \text{ W}/\text{cm}^2 \rightarrow E = 3 \times 10^{13} \text{ V}/\text{m}, B = 10^5 \text{ T}$$

E-field is 10^5 times higher than in conventional RF accelerators

Sunlight near Earth surface: $I \approx 0.1 \text{ W}/\text{cm}^2$

- This laser light converts cold target matter (gas, liquid, solid) almost instantaneously into plasma and drives huge currents.

- Lasers of similar powers may aim at very different experiments:

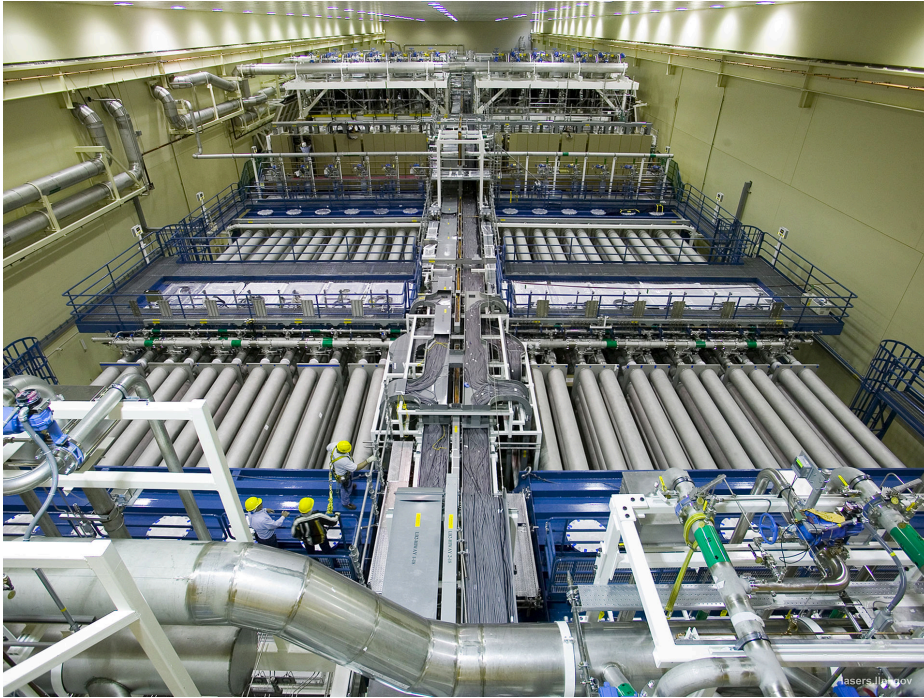


Image:LLNL

NIF: 1.8 MJ (192 beams at 351nm), ns, ~mm
 $P \sim 500\text{TW} \rightarrow I < 10^{16} \text{ W/cm}^2$

Rep.rate – one shot per day



Laser diode chip \approx less than 1 mm



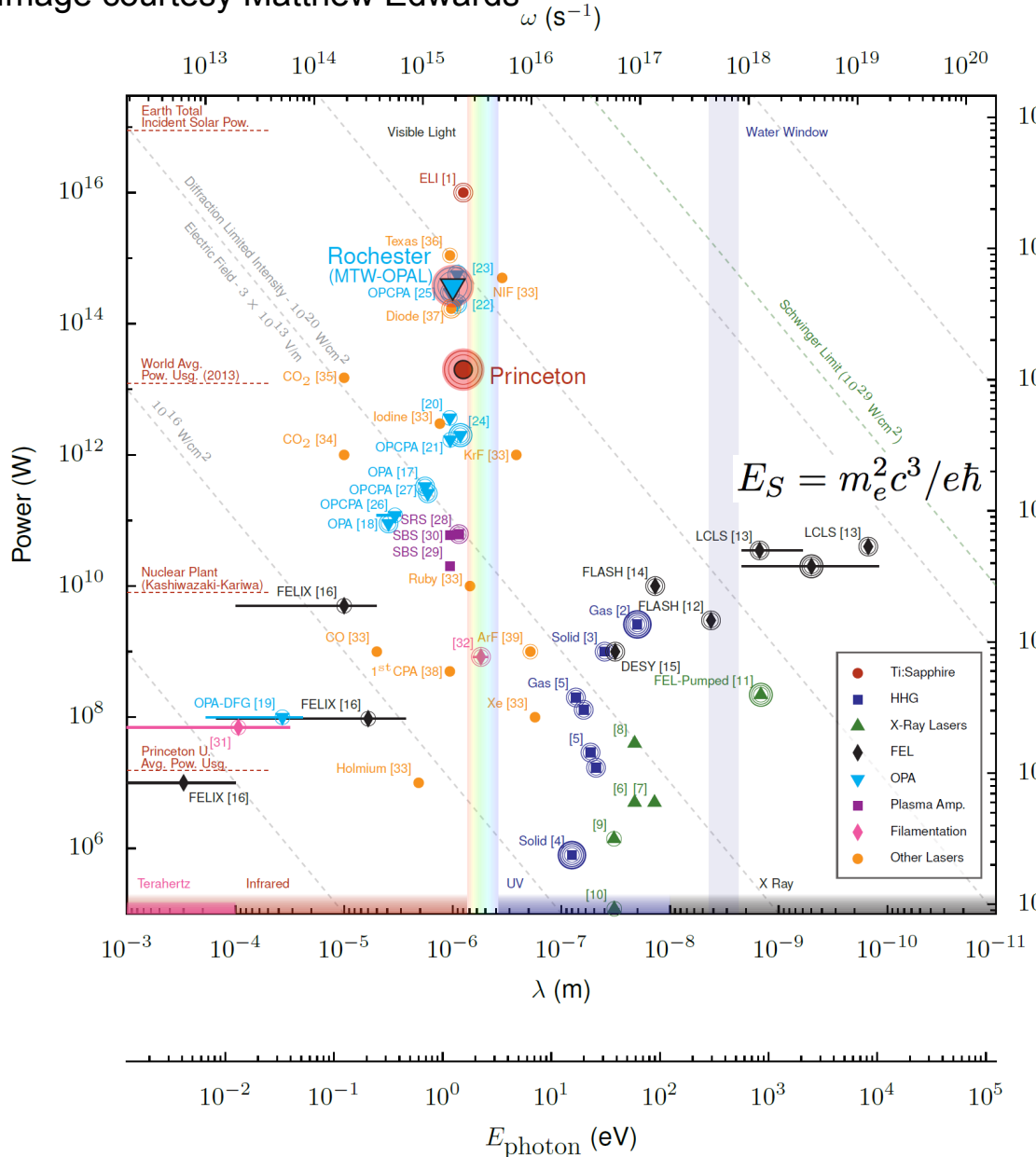
Princeton: 500 mJ (800nm), 25 fs, few μm spot size
 $P \sim 20 \text{ TW} \rightarrow I > 10^{19} \text{ W/cm}^2$

Rep.rate – 10 Hz

Ultrafast light sources

- < 1 ps
 - ◉ < 100 fs
 - ◐ < 10 fs
 - ◑ < 1 fs
- ps = 10^{-12} s
 fs = 10^{-15} s
 as = 10^{-18} s

Image courtesy Matthew Edwards



The shortest optical pulse generated is single-cycle. (tiny energy, yet to be demonstrated on a terawatt scale)

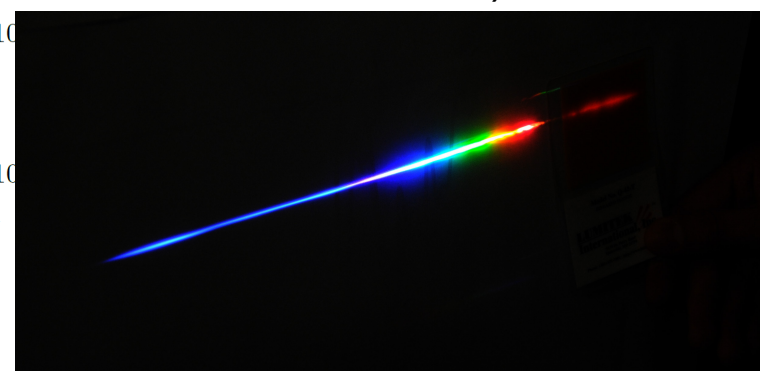


Image: MPQ

Ultrafast light sources

- < 1 ps
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Image courtesy Matthew Edwards

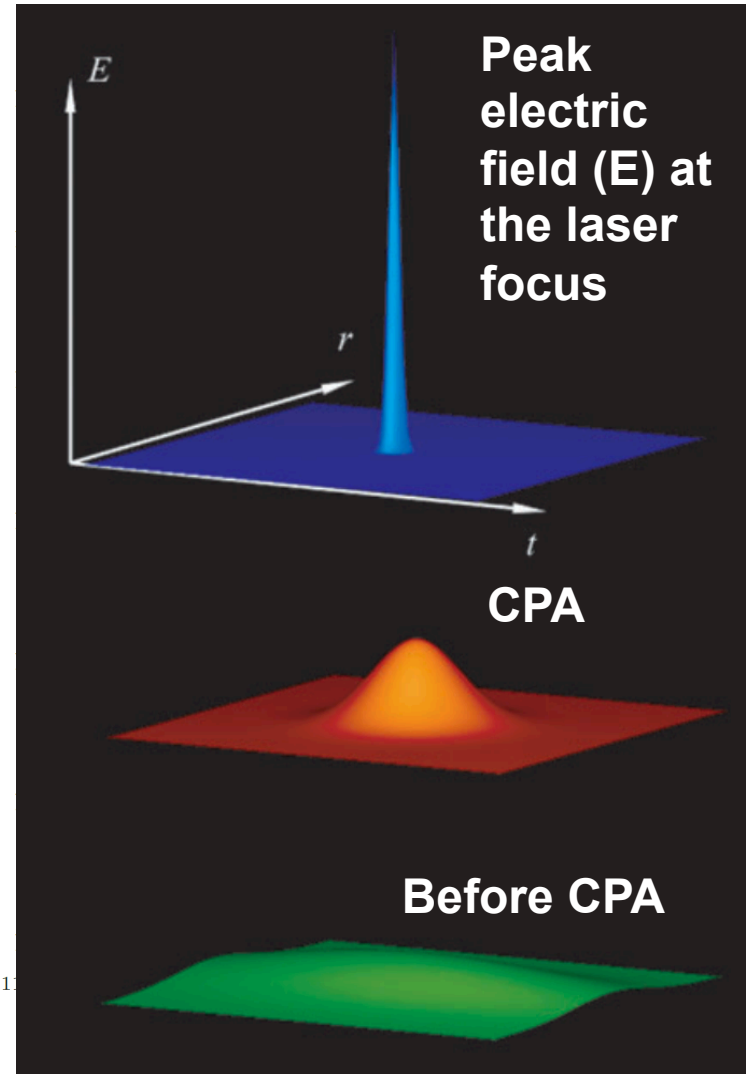
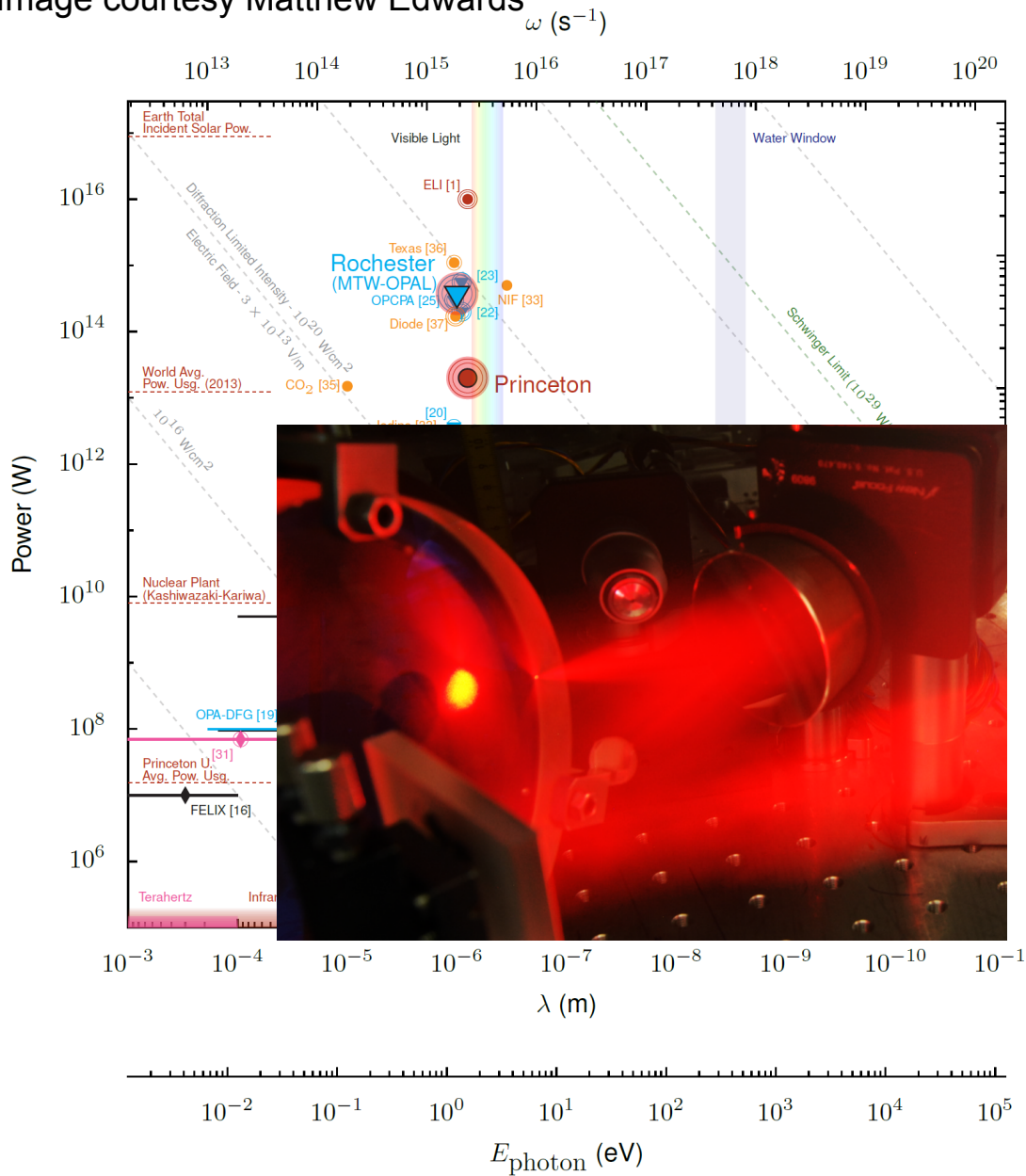


Image: D. Umstadter, J. Phys. D: Appl. Phys. 36, R151, 2003

High-Power Lasers

High power from high energy

- Flash-lamp-pumped glass-type, wavelength around $1\mu\text{m}$ (or 2w, 3w)
- Energies 10 - 1000 Joules
- Durations > 500 fs
- Titan (LLNL), Trident (LANL), MTW (LLE) and others
- Repetition rate $<$ few shots per day

High power from short pulse durations

- Laser-pumped Ti:Sapphire, wavelength = 800 nm
- OPCPA, broadband 650-1100nm
- Energies 0.1 – 100 Joules
- Durations $\sim 25 - 200$ fs
- Hercules (UMichigan), Scarlet (OSU), Callisto (LLNL), MTW-OPAL (LLE), Princeton, etc
- Repetition rate - up to 10 Hz

Chirped pulse amplification

- Invented by Gerard Mourou and Donna Strickland in 1985
- Way of increasing intensities beyond damage thresholds amplifying longer pulses (100 ps – 1 ns) and compressing them to <1ps after amplification.

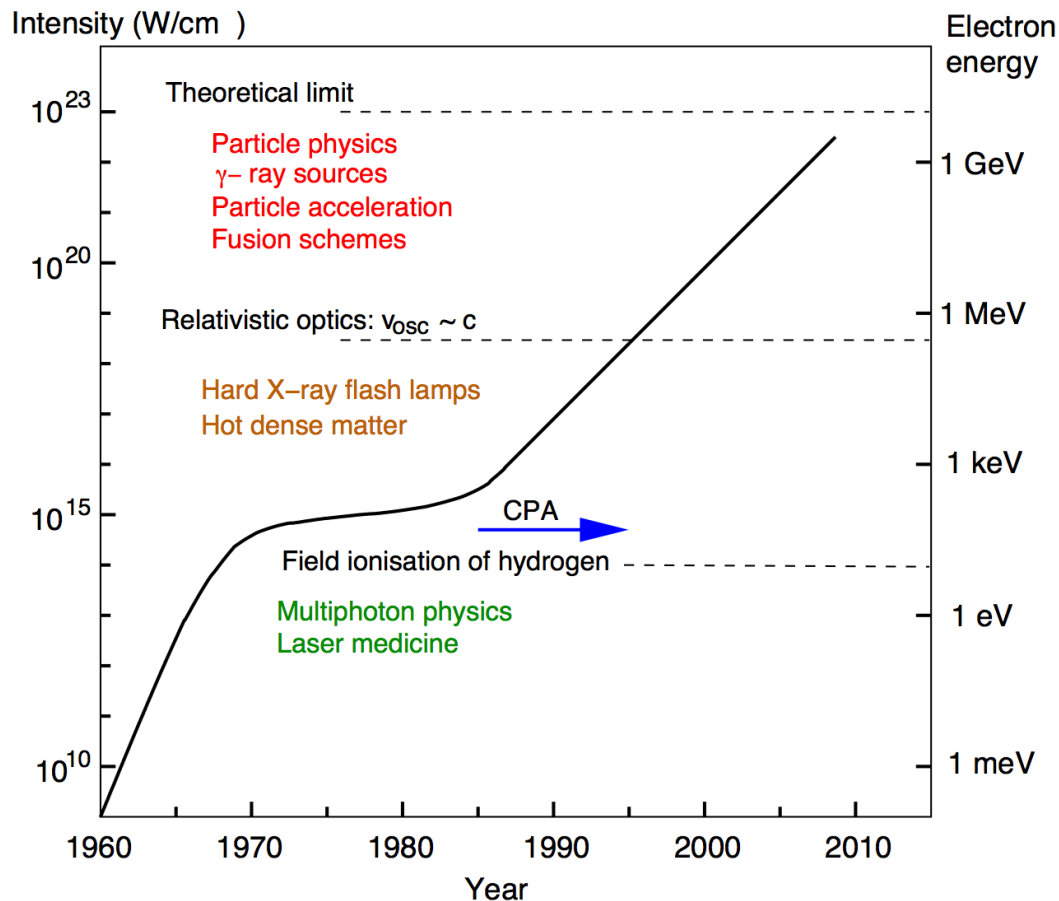
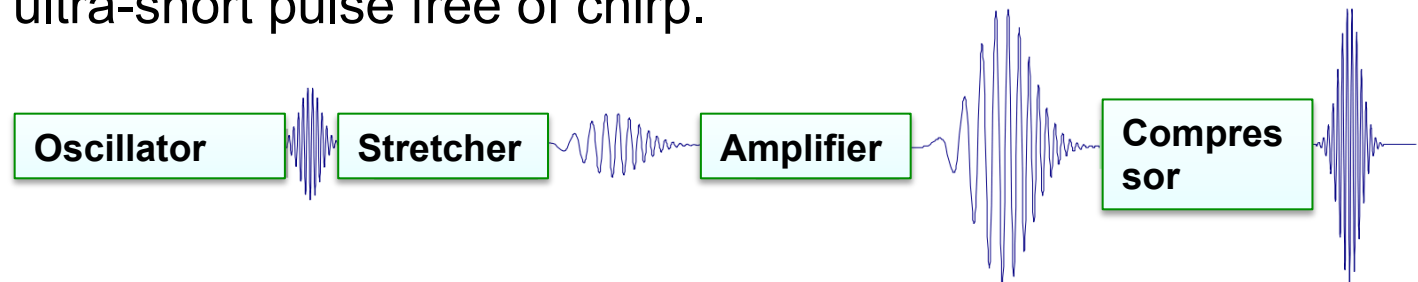


Figure: Paul Gibbon "Short Pulse Laser Interactions with Matter", Lectures

Chirped pulse amplification

- Invented by Gerard Mourou and Donna Strickland in 1985
- Way of increasing intensities beyond damage thresholds amplifying longer pulses (100 ps – 1 ns) and compressing them to <1ps after amplification.

1. **Oscillator:** generates fs, broadband, low-energy pulse
2. **Stretcher:** converts fs pulse to >100 ps. Creates different optical paths for each wavelength of the spectrum. Produces a “chirped” pulse.
3. **Amplifier:** increase the pulse energy by a factor of 10^7 – 10^{10}
4. **Compressor:** performs optical inverse of the stretcher to deliver an amplified fs pulse. The result is a high-intensity ultra-short pulse free of chirp.



Chirped pulse amplification

Typical femtosecond TW laser system

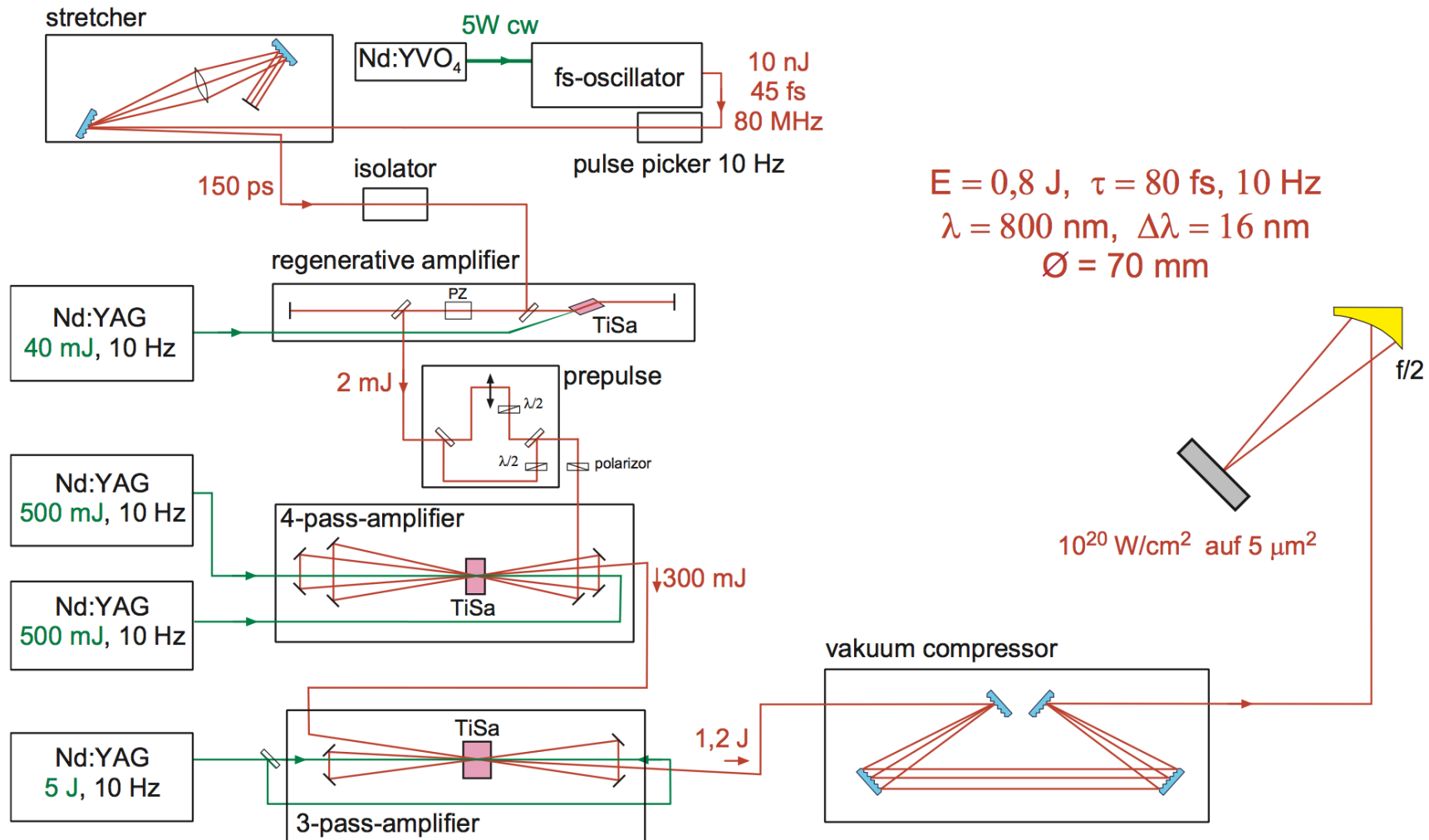


Image from Paul Gibbon "Short Pulse Laser Interactions with Matter", Lectures

Single electron in a plane wave

Assumptions:

- Electron is classical (no spin)
- Neglect back reaction force of electron's radiation
- Electron and wave are in vacuum, there are no other charges, potential electric field = 0

Relativistic momentum $\frac{dp}{dt} = -eE - \frac{e}{c} [\mathbf{v} \times \mathbf{H}]$

Charge e , Electric field E , Electron velocity \mathbf{v} , Magnetic field \mathbf{H} , Speed of light c

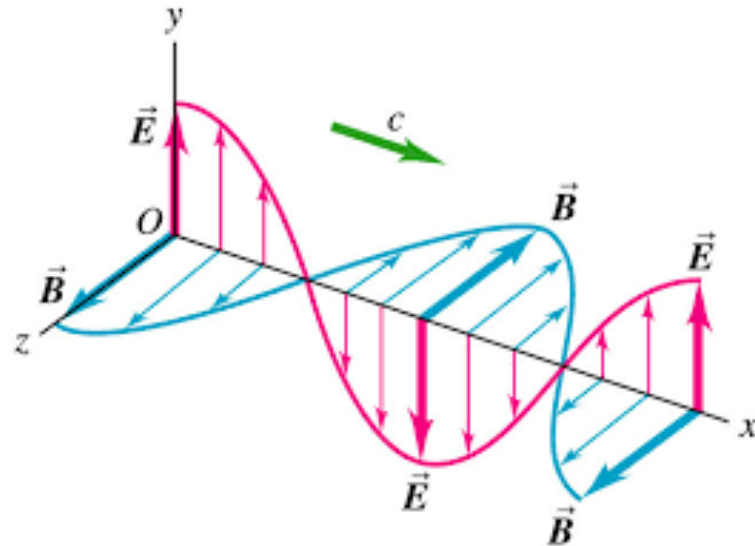
$\frac{dr}{dt} = \mathbf{v} = \frac{\mathbf{p}}{\gamma m}$

Lorentz factor $\gamma = 1 / \sqrt{1 - (v/c)^2} = \sqrt{1 + (p/mc)^2}$

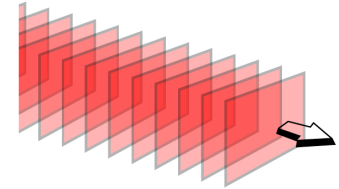
Rest mass m

Relativistic mass - the mass of an electron in motion

Plane electromagnetic wave

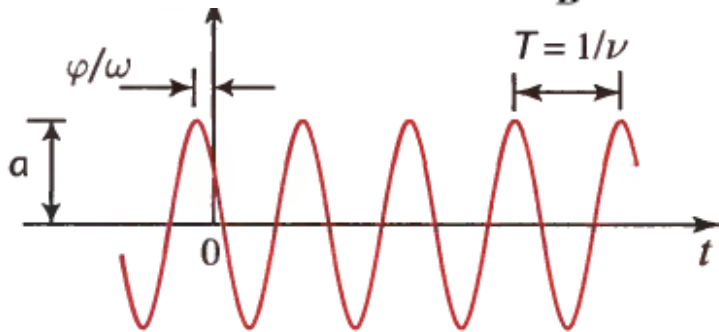


$$\frac{\partial^2 E}{\partial x^2} - \frac{1}{c^2} \frac{\partial^2 E}{\partial t^2} = 0$$



Sinusoidal plane wave is propagating along x-axis, its wavefronts are parallel planes.

$$E(x, t) = E(t - x/c) + E(t + x/c)$$



It is convenient to use the field vector potential:

$$A(t - x/c)$$

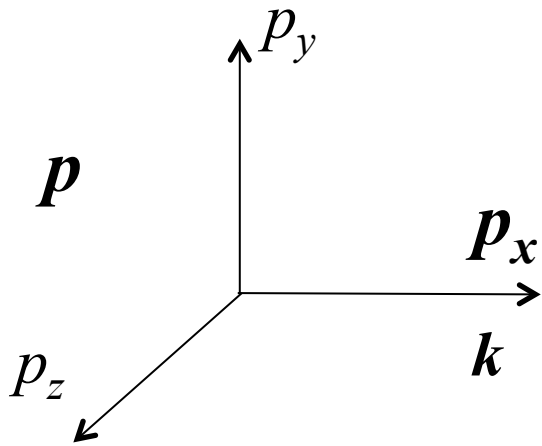
$$E = -\partial A / c \partial t \quad B = \nabla \times A$$

A plane wave is useful, because:

- Any function, which is smooth and rapidly decreasing in infinity, can be represented as a superposition of plane waves - Fourier transform in time and space.
- If a wavefront curvature radius of an arbitrary wave \gg than linear dimensions of a given volume, then inside this volume the wave is almost plane. E.g. atomic radius is about 0.1 nm \ll wavelength of visible light (400–700 nm).

Single electron in a plane wave.

Low field - Non-relativistic motion



Transverse motion

$$m \frac{d}{dt} \vec{v}^{(1)} = -e \vec{E}_0 \cos \omega t$$

$$\vec{v}^{(1)} = \frac{e \vec{E}_0 \sin \omega t}{m \omega}$$

$$\vec{r}^{(1)} = \frac{e \vec{E}_0 \cos \omega t}{m \omega^2}$$

$$\vec{E}(\vec{r}, t) = \vec{E}_0(\vec{r}) \cos \omega t$$

$$\frac{dp}{dt} = -eE - \frac{e}{c} [\cancel{v \times H}] \cos \omega t$$

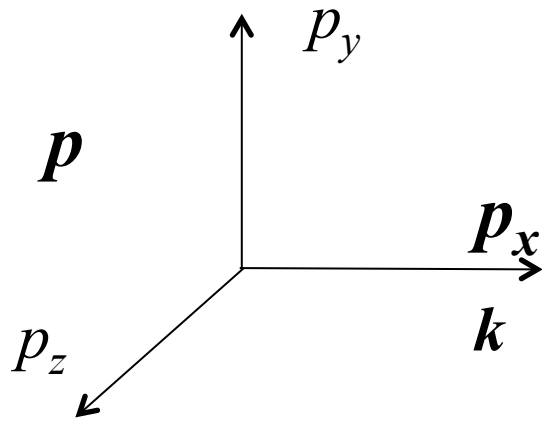
Longitudinal drift motion

There is no drift in the direction of laser propagation

Electron oscillates at the laser frequency in the direction parallel to the E field vector.

The amplitude of electron oscillation is less than the light wavelength.

Single electron in a plane wave. High field - Relativistic motion



$$\frac{dp}{dt} = -eE - \frac{e}{c}[\mathbf{v} \times \mathbf{H}]$$

$$A(t - x/c)$$

$$\mathbf{E} = -\partial A / c \partial t \quad \mathbf{B} = \nabla \times A$$

Transverse motion

$$\mathbf{p} = \mathbf{f} + eA/c$$

$\mathbf{f} = \text{const}$, $p_{x0} = \text{const}$, $E_0 = \text{const}$
are the initial electron momenta and
energy before the interaction with a wave.

Longitudinal drift motion

$$p_x = p_{x0} + \frac{feA}{E_0 - cp_{x0}} + \frac{e^2 A^2}{2c(E_0 - cp_{x0})}$$

This drift motion originates from

$$\mathbf{v} \times \mathbf{B} \propto E^2 \hat{\mathbf{k}}$$

Single electron in a plane wave. High field - Relativistic motion

Note:

Transverse motion

$$\mathbf{p} \sim \mathbf{A}$$

Longitudinal drift motion

$$p_x \sim A^2$$

For small A , $p \gg p_x$. For large A , $p_x \gg p$.

$$E = E_0 + c \frac{efA}{E_0 - cp_{x0}} + \frac{e^2 A^2}{2(E_0 - cp_{x0})}$$

When the field is vanishing ($A=0$), the electron momenta and energy return to their initial values (f, p_{x0}, E_0).

Electron's coordinates can be found in the parametric form:

$$r(\xi) = \frac{c^2}{E_0 - cp_{x0}} \int (f + \frac{e}{c} A) d\xi$$

$$x(\xi) = \frac{c^2}{E_0 - cp_{x0}} \int p_x(\xi) d\xi$$

$$t(\xi) = \xi + x(\xi) / c$$

Special case

Electron at rest before the wave arrival: $f=0$, $p_{x0}=0$

Transverse motion

$$p = \frac{e}{c} A$$

$$v = \frac{1}{\gamma m} \frac{e}{c} A$$

Longitudinal drift motion

$$p_x = \frac{e^2 A^2}{2mc^3}$$

$$v_x = c \frac{\gamma - 1}{\gamma}$$

$$E = mc^2 + \frac{e^2 A^2}{2mc^2}$$

$$\gamma = 1 + \frac{e^2 A^2}{2m^2 c^4}$$

$$A_{rel} = mc^2 / e$$

$$A = a A_{rel}$$

$$\gamma = 1 + a^2 / 2$$

a is a dimensionless figure of merit of the laser vector potential

At $a = 1$, electron kinetic energy = $\frac{mc^2}{2}$

Relativistic intensity of light

$$a_0 = A_0/A_{rel} = E_0/E_{rel} = \sqrt{I/I_{rel}}$$

$$A_{rel} = mc^2 / e \quad E_{rel} = mc\omega / e$$

$$I_L = \frac{1}{2} \varepsilon_0 c E_0^2; \quad \lambda_L = \frac{2\pi c}{\omega}$$

$$a_0 \simeq 0.85 (I_{18} \lambda_\mu^2)^{1/2}, \quad I_{18} = \frac{I_L}{10^{18} \text{ Wcm}^{-2}}; \quad \lambda_\mu = \frac{\lambda_L}{\mu\text{m}}$$

$$I_{\text{W/cm}^2} = 1.37 \times 10^{18} \left(\frac{a_0}{\lambda_{\mu\text{m}}} \right)^2$$

$$800 \text{ nm}: I_{rel} = 2.14 \times 10^{18} \text{ W/cm}^2$$

$$10 \mu\text{m}: I_{rel} = 1.37 \times 10^{16} \text{ W/cm}^2$$

Special case

Electron at rest before the wave arrival: $f=0$, $p_{x0}=0$

$$A_{rel} = mc^2 / e$$

$$A = aA_{rel}$$

$$\gamma = 1 + a^2 / 2$$

Transverse motion

$$v = c \frac{a}{1 + a^2 / 2}$$

$$v \leq c / \sqrt{2}$$

$$v = c / \sqrt{2} \quad a = \sqrt{2}$$

Longitudinal drift motion

$$v_x = c \frac{a^2 / 2}{1 + a^2 / 2}$$

v_x becomes close to c
with increasing a

Special case

Electron at rest before the wave arrival: $f=0$, $p_{x0}=0$

Transverse motion

$$r = \frac{e}{mc} \int_{-\infty}^{\xi} A d\xi' + r_0$$

If $A = A_{amp} \sin \xi$, then

$$r \approx r_0 - r_{amp} \cos \xi$$

$$r_{amp} = \frac{\lambda}{2\pi} a_{amp}$$

Longitudinal drift motion

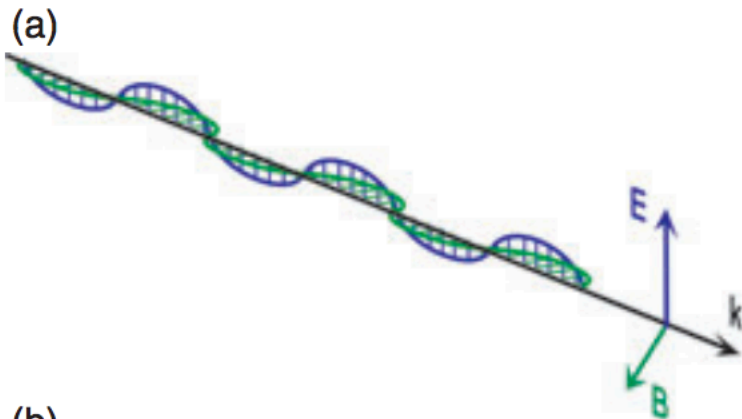
$$x = \frac{e^2}{2m^2 c^3} \int_{-\infty}^{\xi} A^2 d\xi' + x_0$$

$$t = \xi + \frac{e^2}{2m^2 c^4} \int_{-\infty}^{\xi} A^2 d\xi' + x_0 / c$$

$$a_{amp} = \frac{A_{amp}}{A_{rel}} = \frac{E_{amp}}{E_{rel}} = \sqrt{I / I_{rel}}$$

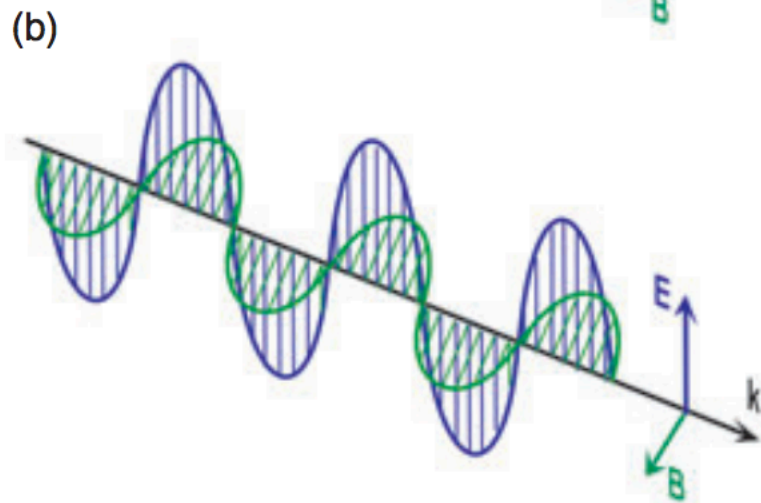
The amplitude of electron oscillation can be **much higher than the laser wavelength**.

Single electron in a plane wave



(a) **Non-relativistic case:** the wave amplitude is small, only the **E-field** acts on the electron, electron oscillates in the direction of the E field at the light's frequency, **velocity** $\ll c$, **displacement amplitude** $< \lambda$.

There is no displacement along the wave propagation direction.



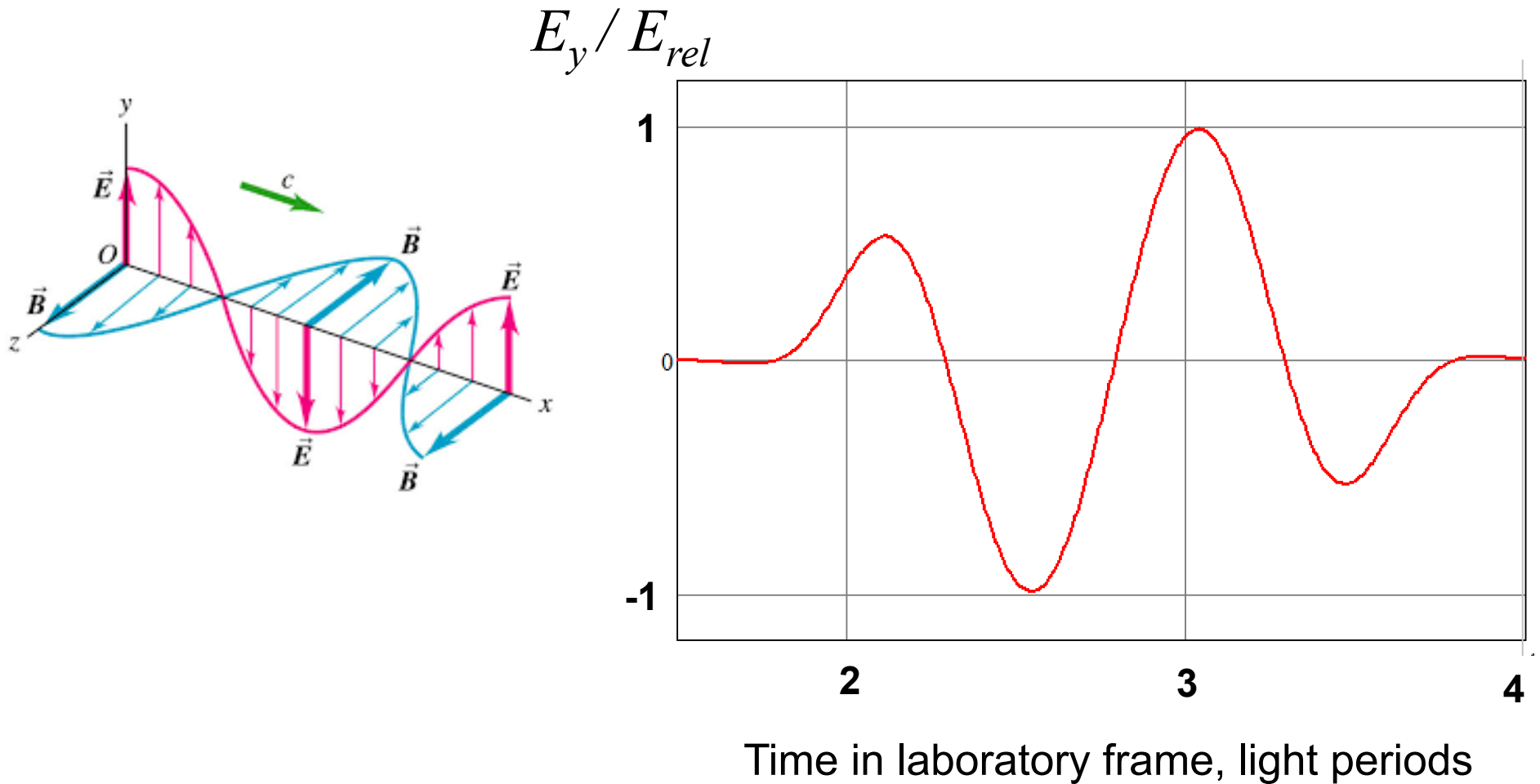
(b) **Relativistic case:** the wave amplitude is very large, **B-field** becomes important, the combined action of the E and B-fields pushes the electron forward. Transverse **velocity is limited by $c/\sqrt{2}$** , **longitudinal velocity is close to c** .

Image: D. Umstadter, J. Phys. D: Appl. Phys. 36, R151, 2003

Displacement amplitude $> \lambda$.

Thus, the plane wave approximation may fail for finite-size beams.

Single electron in a laser pulse



Plot the solution of the equation of motion for an electron, which was at rest before the laser pulse arrival.

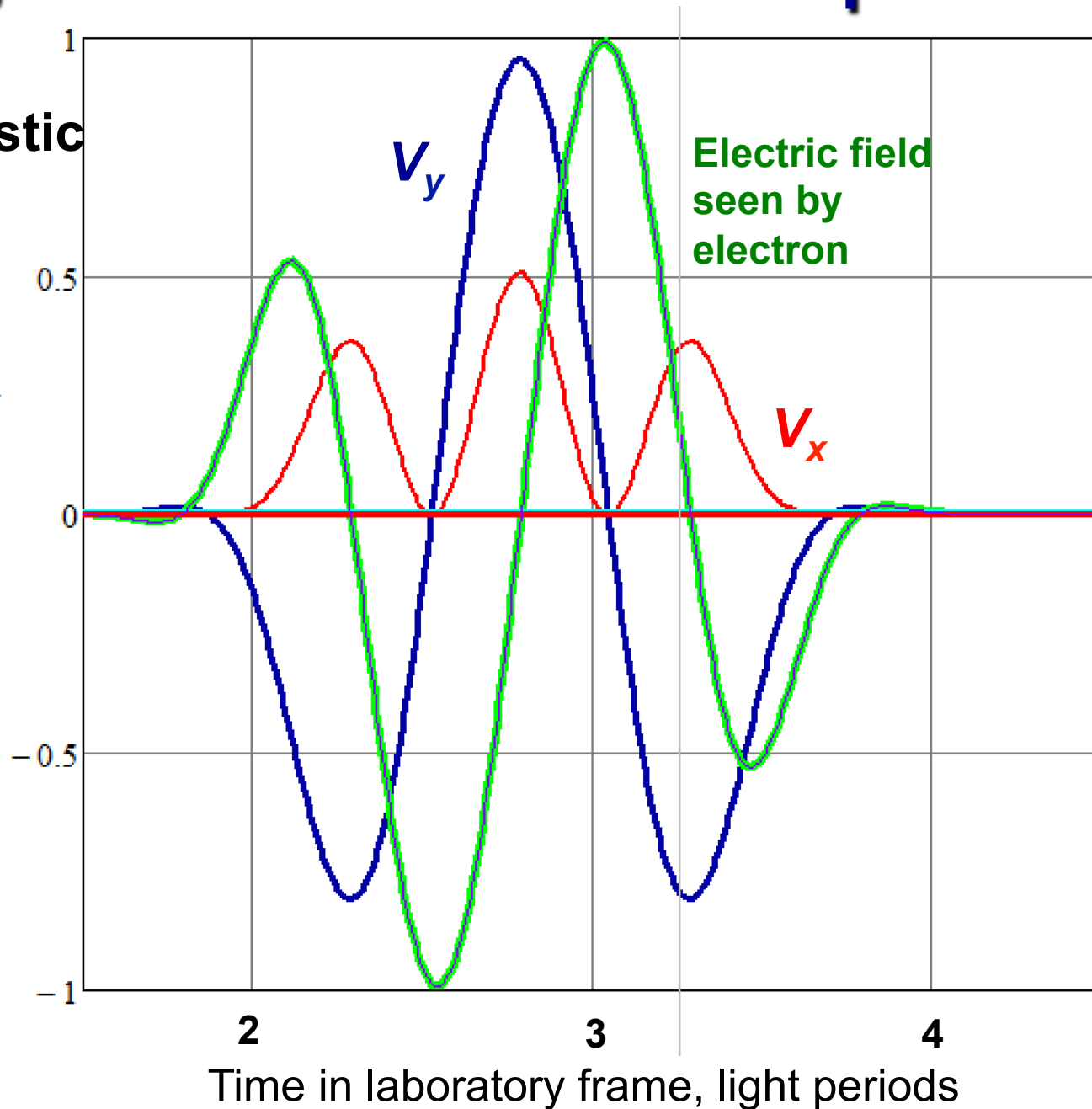
Single electron in a laser pulse

Non-Relativistic case

$$a_0 = 0.0001$$

$$(V_y / c) \times 10^4$$

$$(V_x / c) \times 10^8$$



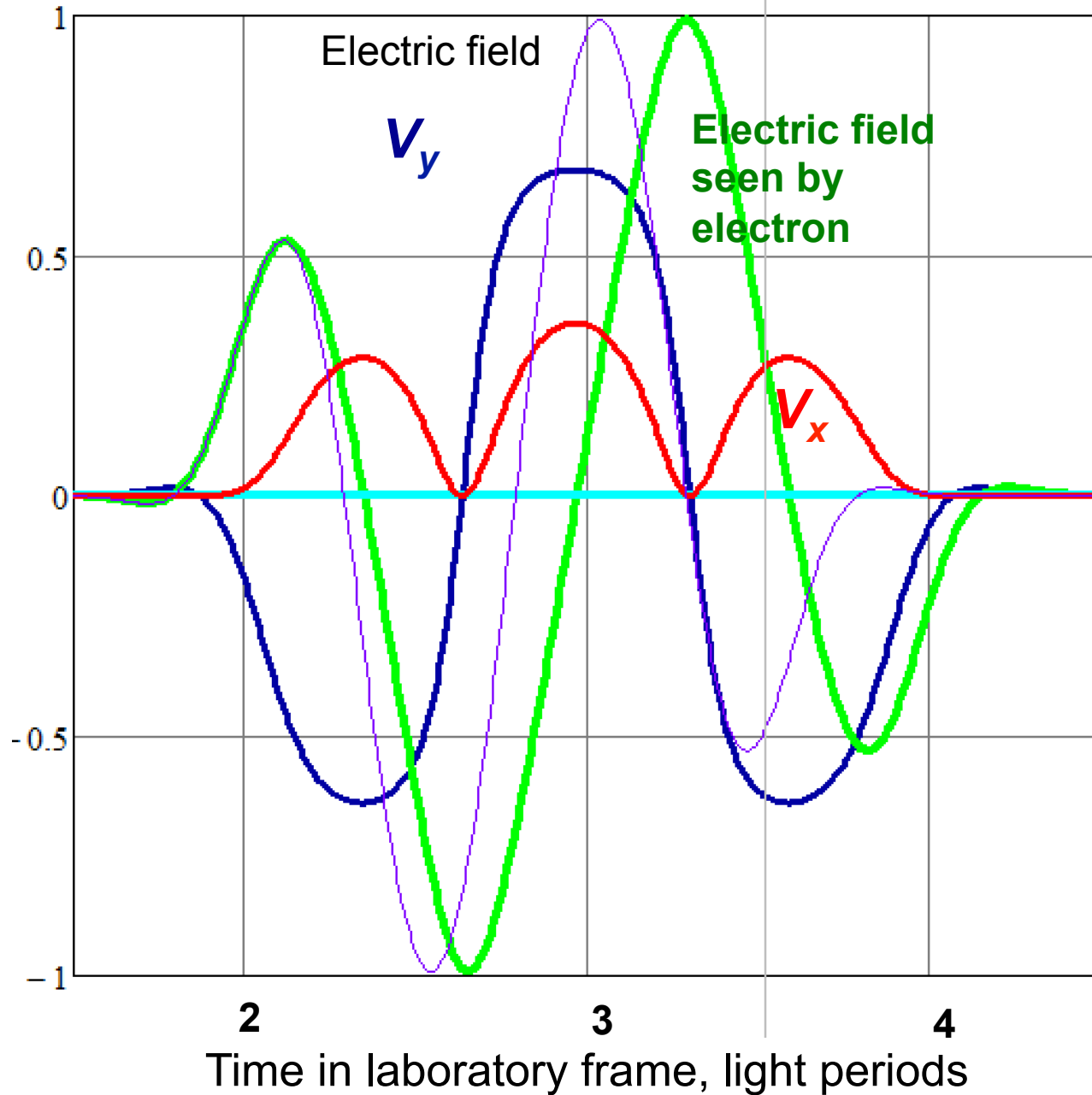
Single electron in a laser pulse

Onset of
Relativistic
effects

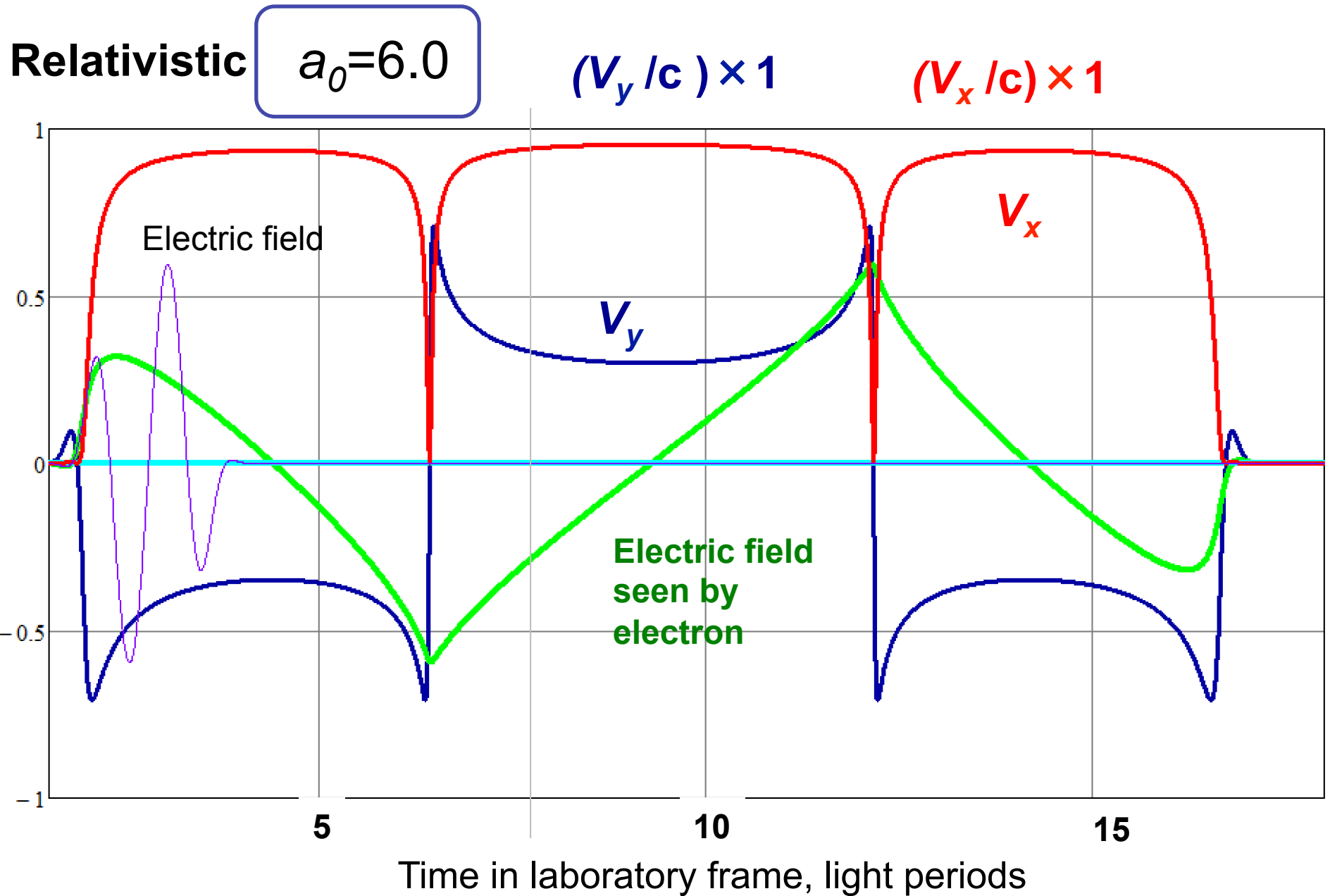
$$a_0 = 1.0$$

$$(V_y / c) \times 1$$

$$(V_x / c) \times 1$$



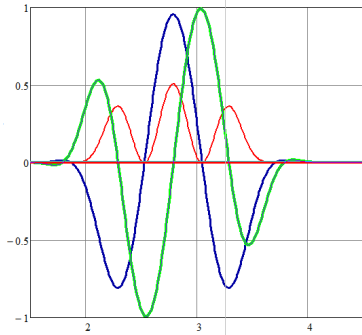
Single electron in a laser pulse



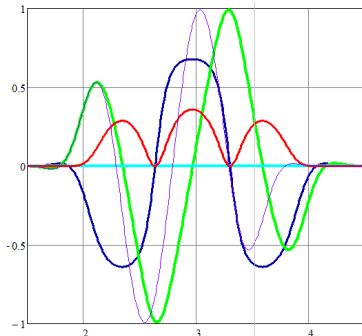
Single electron in a laser pulse

Non-Relativistic

$$a_0 = 0.0001$$

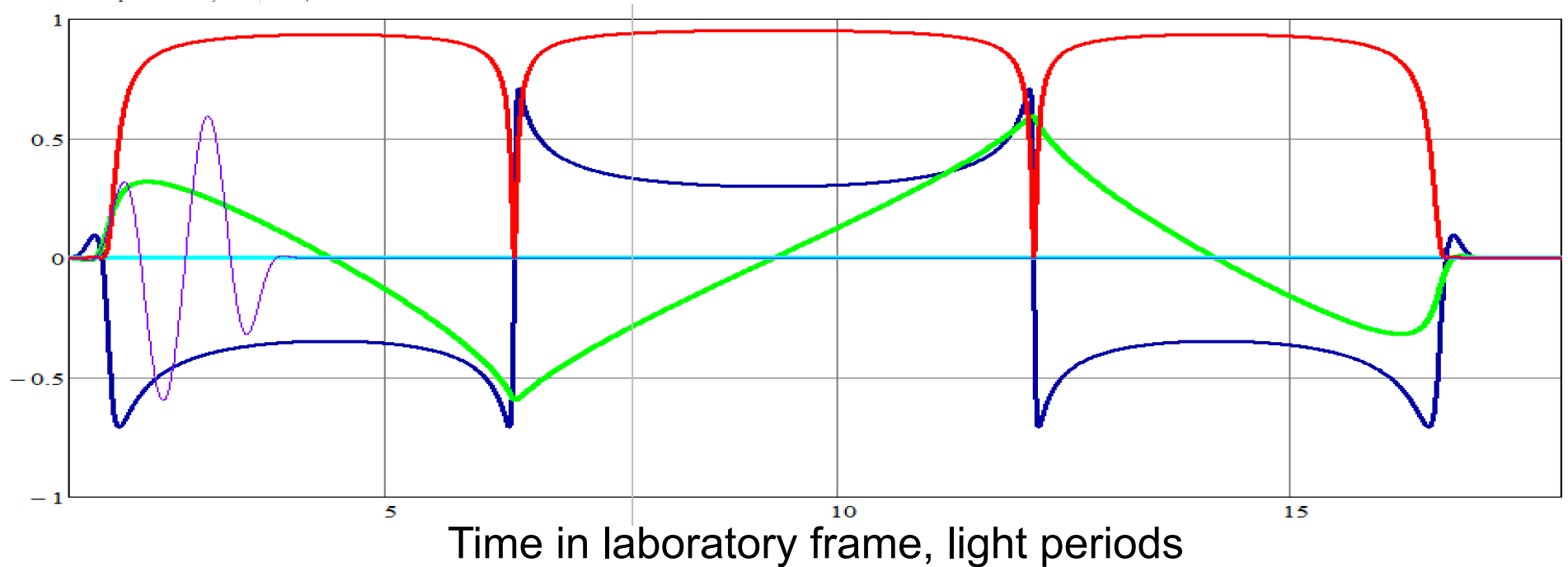


$$a_0 = 1.0$$



$$a_0 = 6.0$$

Relativistic



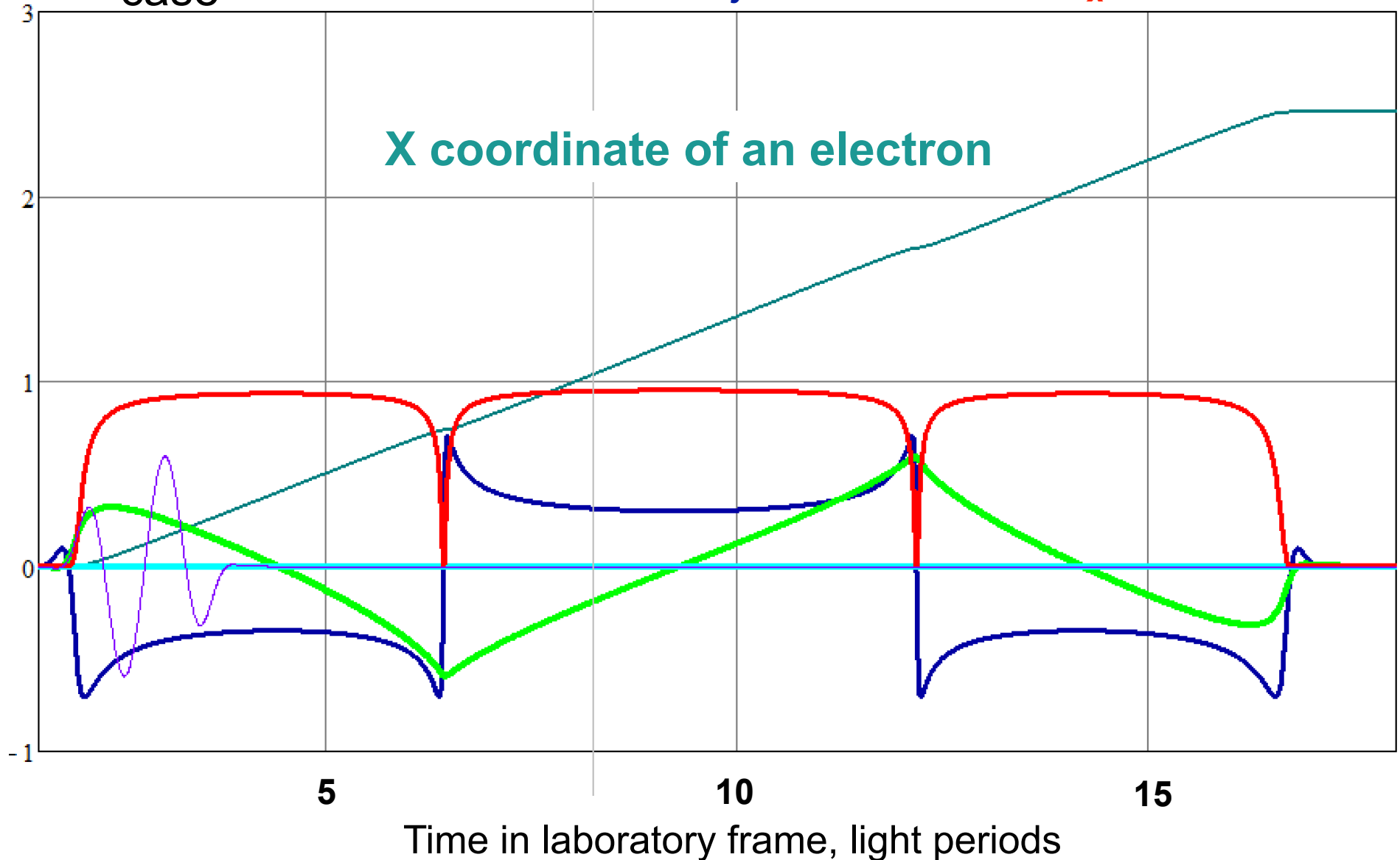
Single electron in a laser pulse

Relativistic
case

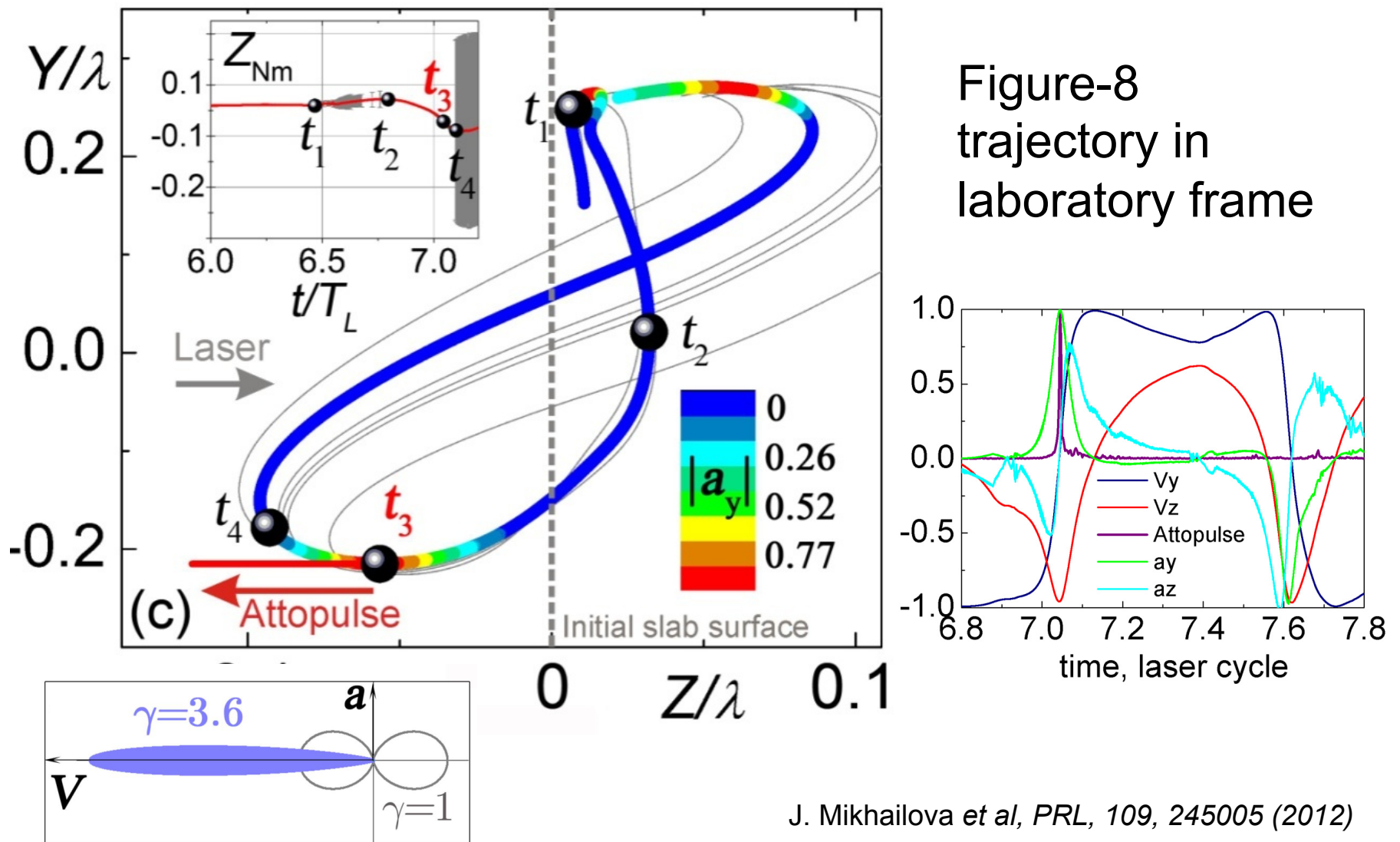
$$a_0 = 6.0$$

$$(V_y / c) \times 1$$

$$(V_x / c) \times 1$$



What if an electron sits in dense plasmas?



Relativistic Laser - Plasma Interactions

- Cold matter — solid, liquid or gas — is rapidly ionized when subjected to strong laser fields, far exceeding binding fields of atoms.

$$I_a = \frac{\epsilon_0 c E_a^2}{2}$$
$$\simeq 3.51 \times 10^{16} \text{ Wcm}^{-2}$$

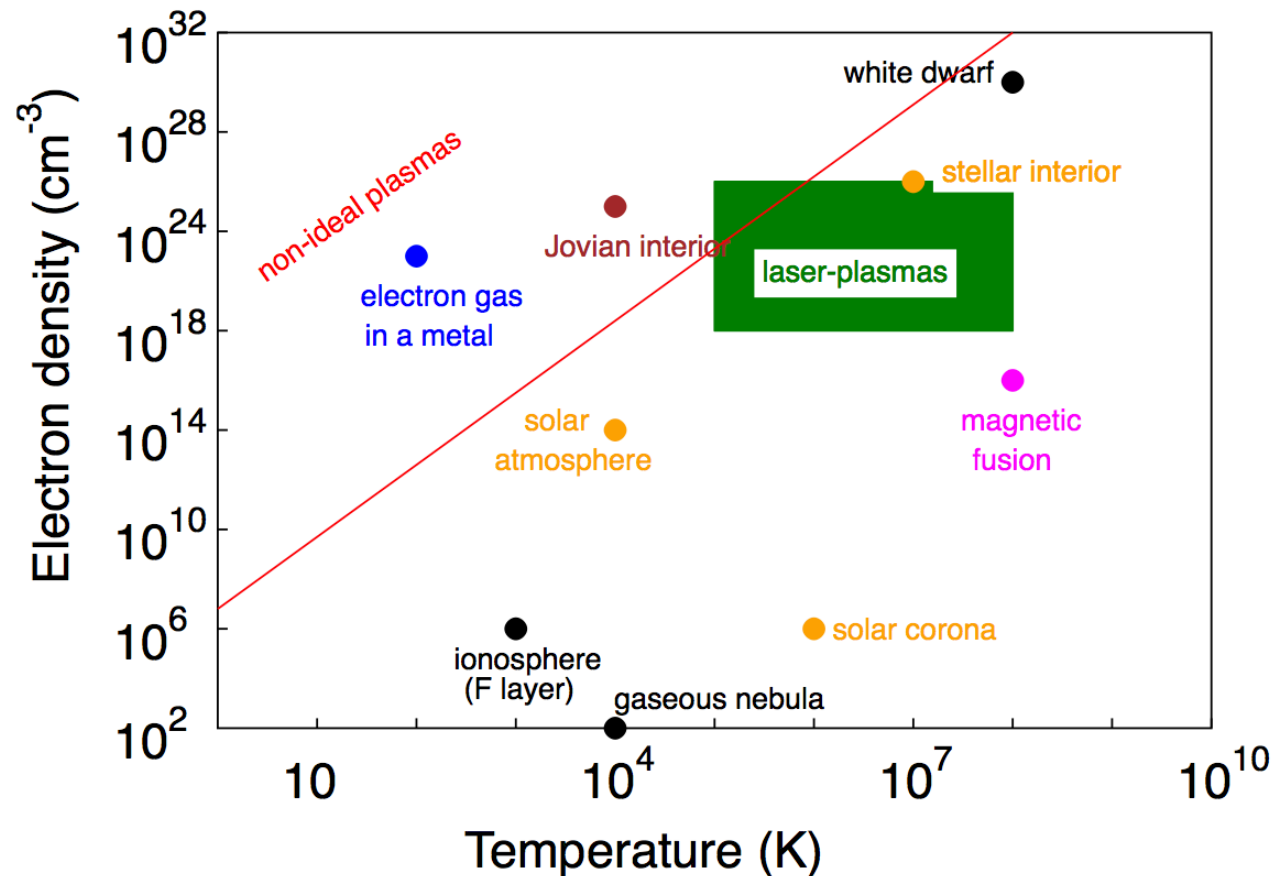
laser intensity of $I_L > I_a$ will *guarantee ionization* for any target material, though in fact this can occur well below this threshold value via other ionization mechanisms

- Electrons released are immediately caught in the laser field, oscillate and drift - field dominated physics.
- Capability for laser-based particle acceleration (GeV), and short-wavelength (x-ray) radiation sources.

Plasma classification: Ideal and Non-ideal

An *ideal* plasma has many particles per Debye sphere:

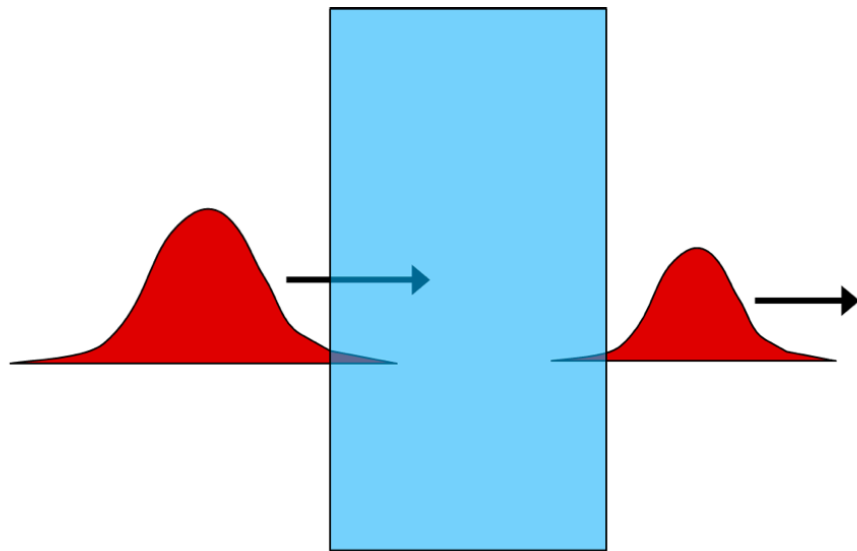
$$N_D \equiv n_e \frac{4\pi}{3} \lambda_D^3 \gg 1.$$



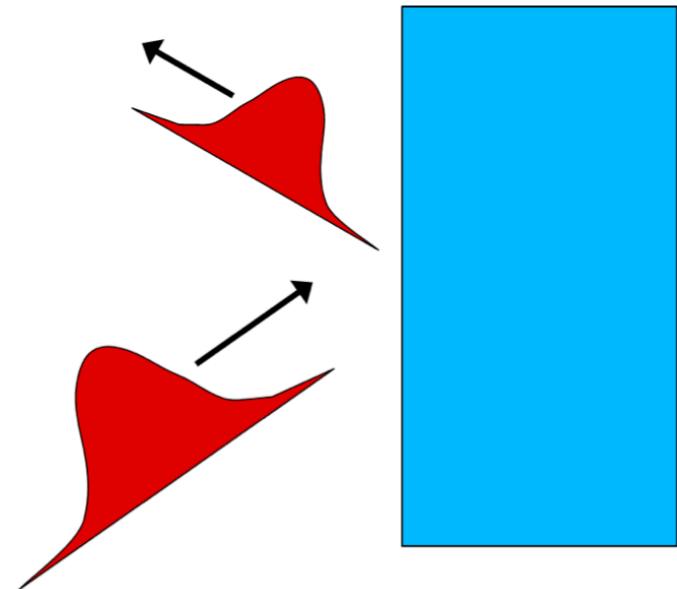
Underdense and Overdense

If the plasma response time is shorter than the period of an external electromagnetic field (such as a laser), then this radiation will be *shielded out*.

$$\omega_p \equiv \left(\frac{e^2 n_e}{\epsilon_0 m_e} \right)^{1/2} \simeq 5.6 \times 10^4 \left(\frac{n_e}{\text{cm}^{-3}} \right)^{1/2} \text{ s}^{-1}$$



Underdense, $\omega > \omega_p$:
plasma acts as nonlinear
refractive medium



Overdense, $\omega < \omega_p$:
plasma acts like mirror

Critical plasma density

Normalized target density

$$N = \frac{n_e}{n_c}$$

$$\frac{\omega_p^2}{\omega^2} = \frac{e^2 n_e}{\epsilon_0 m_e} \cdot \frac{\lambda^2}{4\pi^2 c^2} = 1$$

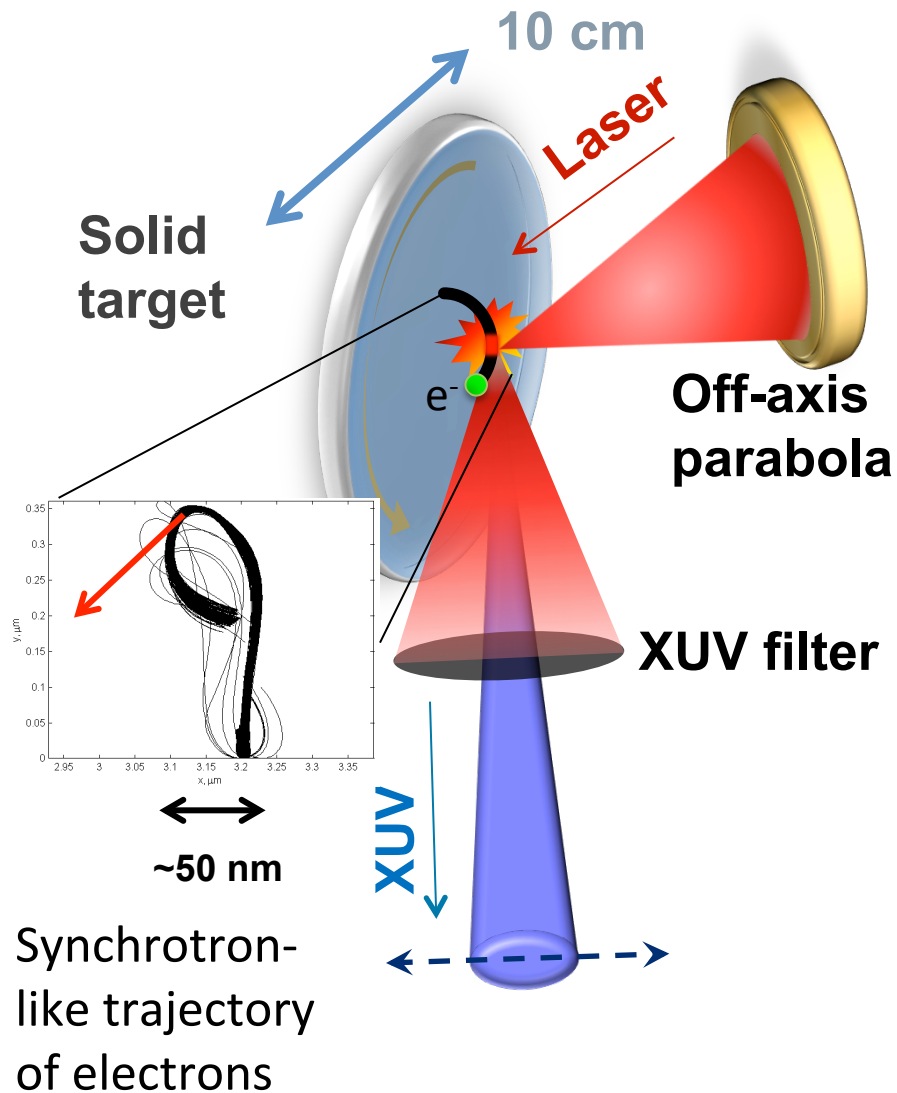
$$n_c \simeq 10^{21} \lambda_{\mu}^{-2} \text{ cm}^{-3}$$

$$800 \text{ nm: } n_c = 1.75 \times 10^{21} \text{ cm}^{-3}$$

$$10 \mu\text{m: } n_c = 1.12 \times 10^{19} \text{ cm}^{-3}$$

Target material	Electron density n_e (cm^{-3})	n_e/n_c (800nm)
Capillary discharge	10^{16}	10^{-5}
Gas jet	10^{18}	10^{-3}
Foam/aerogel	10^{21}	0.1 – 5
Frozen H	10^{22}	36
CH foil	5×10^{23}	600

High-Intensity Laser-Solid Interaction



- High intensity laser pulse hitting solid target creates a dense plasma and accelerates electrons at the surface to relativistic speeds.
- Relativistic electron motion is highly-nonlinear, allowing the generation of high harmonics of the incident laser.
- This emission can, according to theory, be phase-locked to produce attosecond x-ray pulses.

Vacuum

Plasma
(Electron Density)

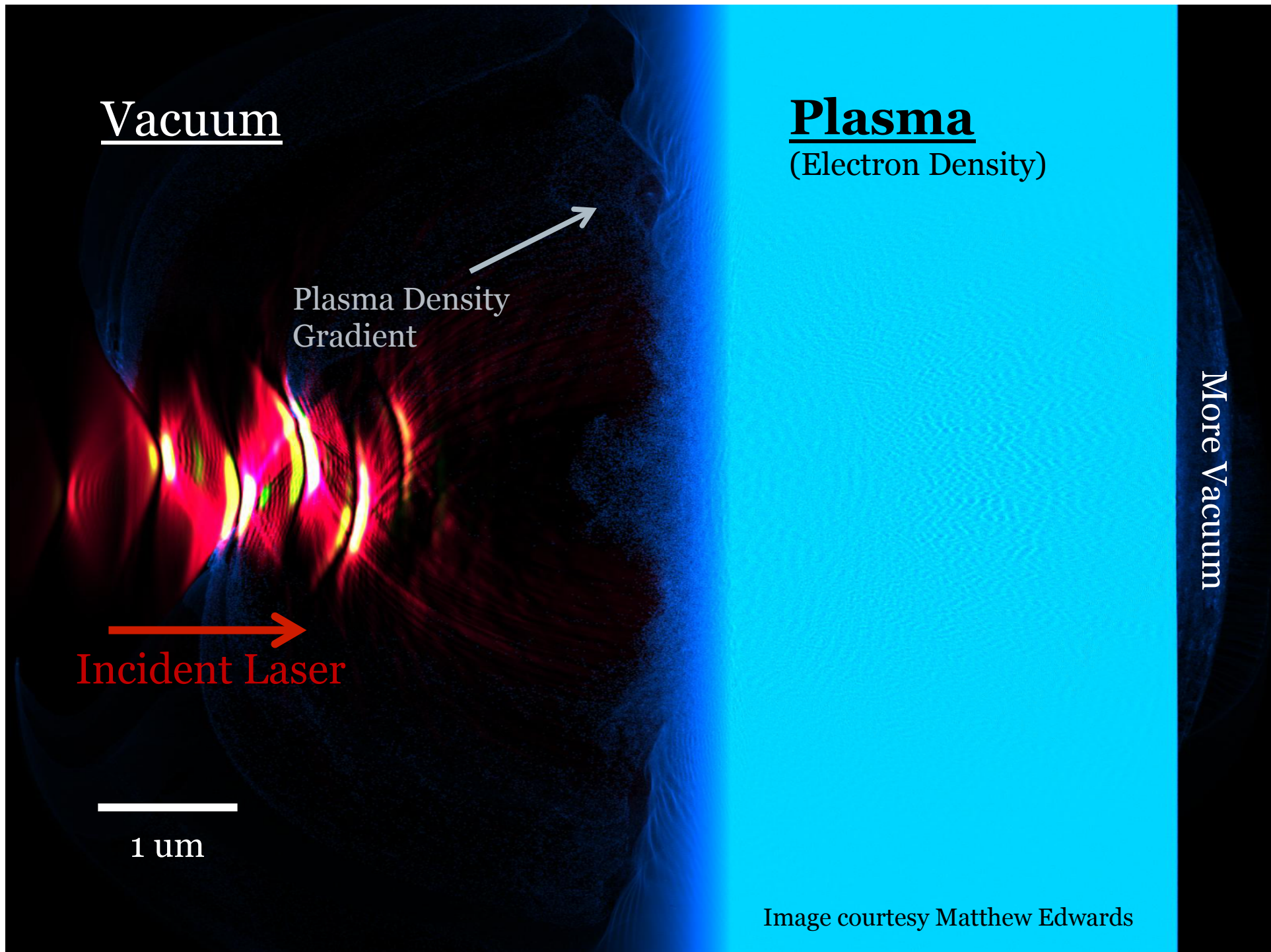
Plasma Density
Gradient

More Vacuum

Incident Laser

1 μm

Image courtesy Matthew Edwards



Accelerated
Electrons



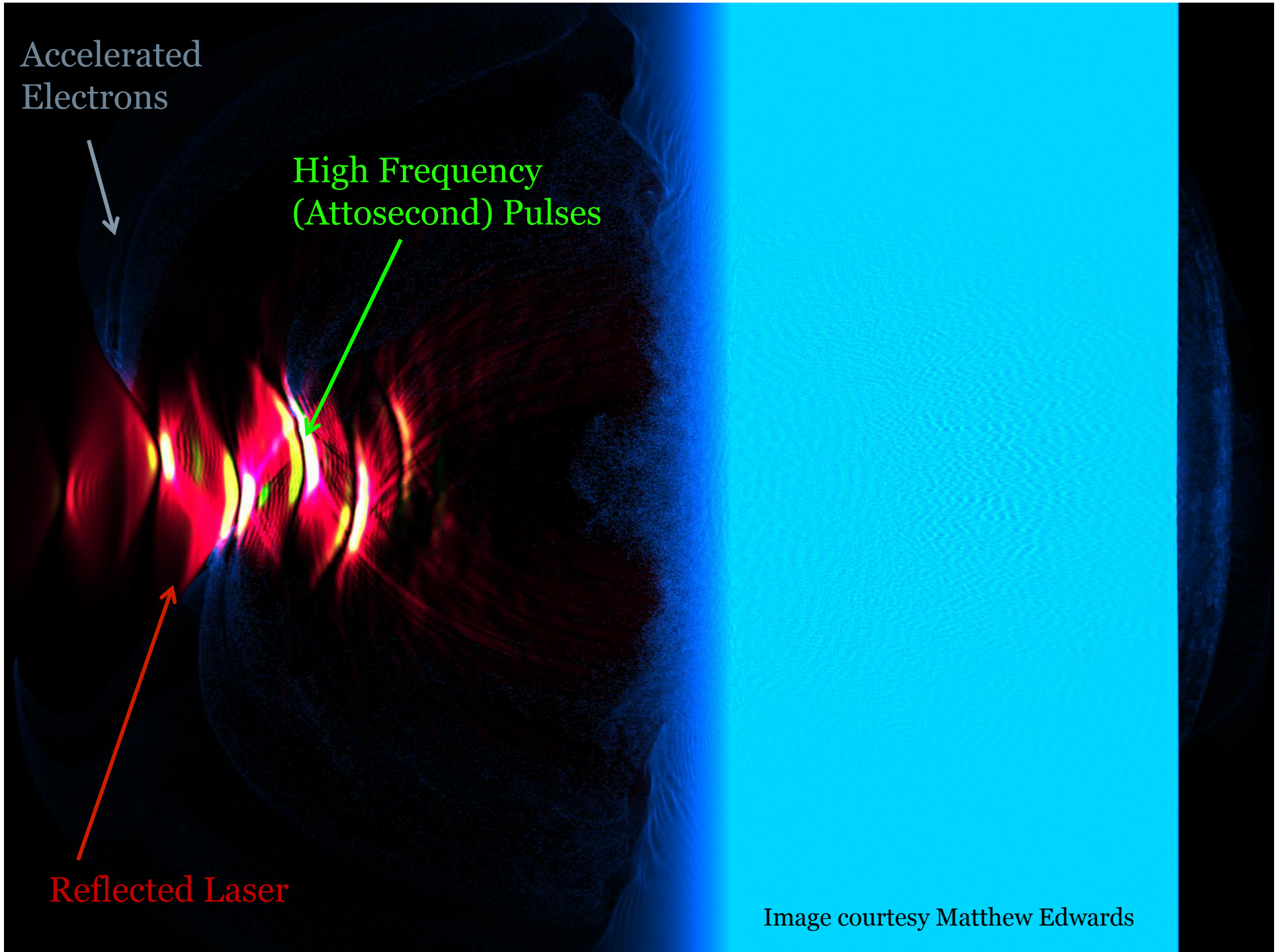
High Frequency
(Attosecond) Pulses

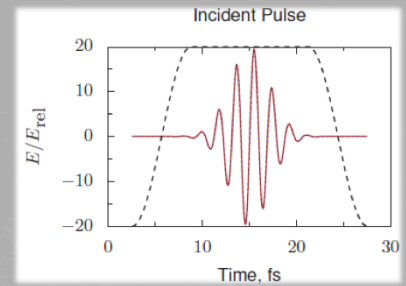
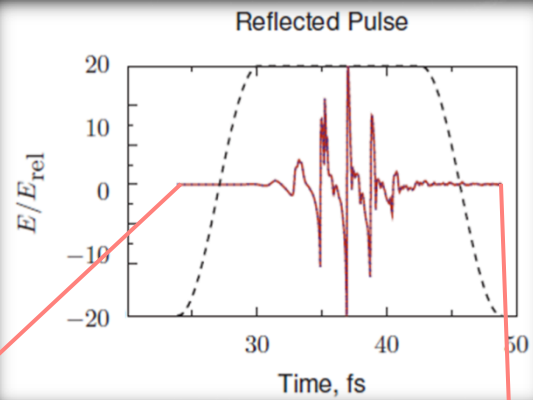


Reflected Laser

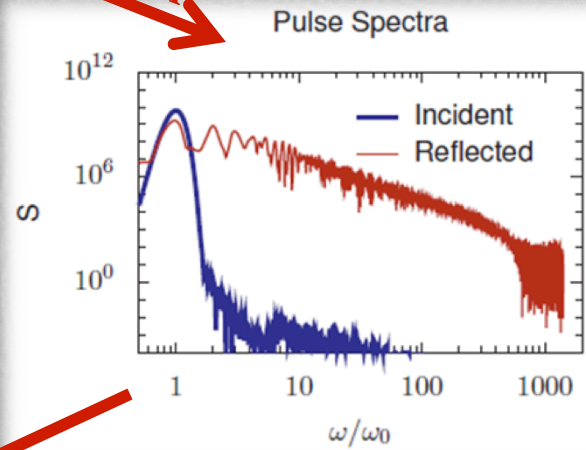


Image courtesy Matthew Edwards

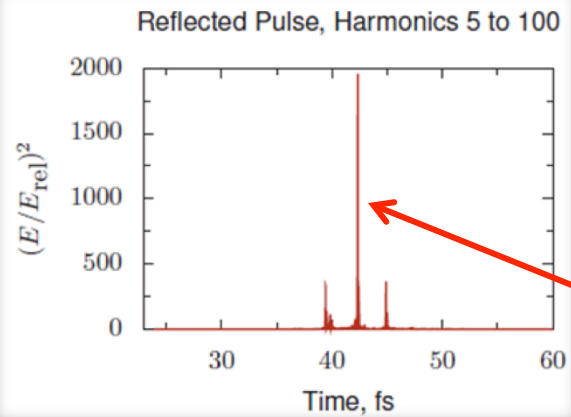




Fourier Transform



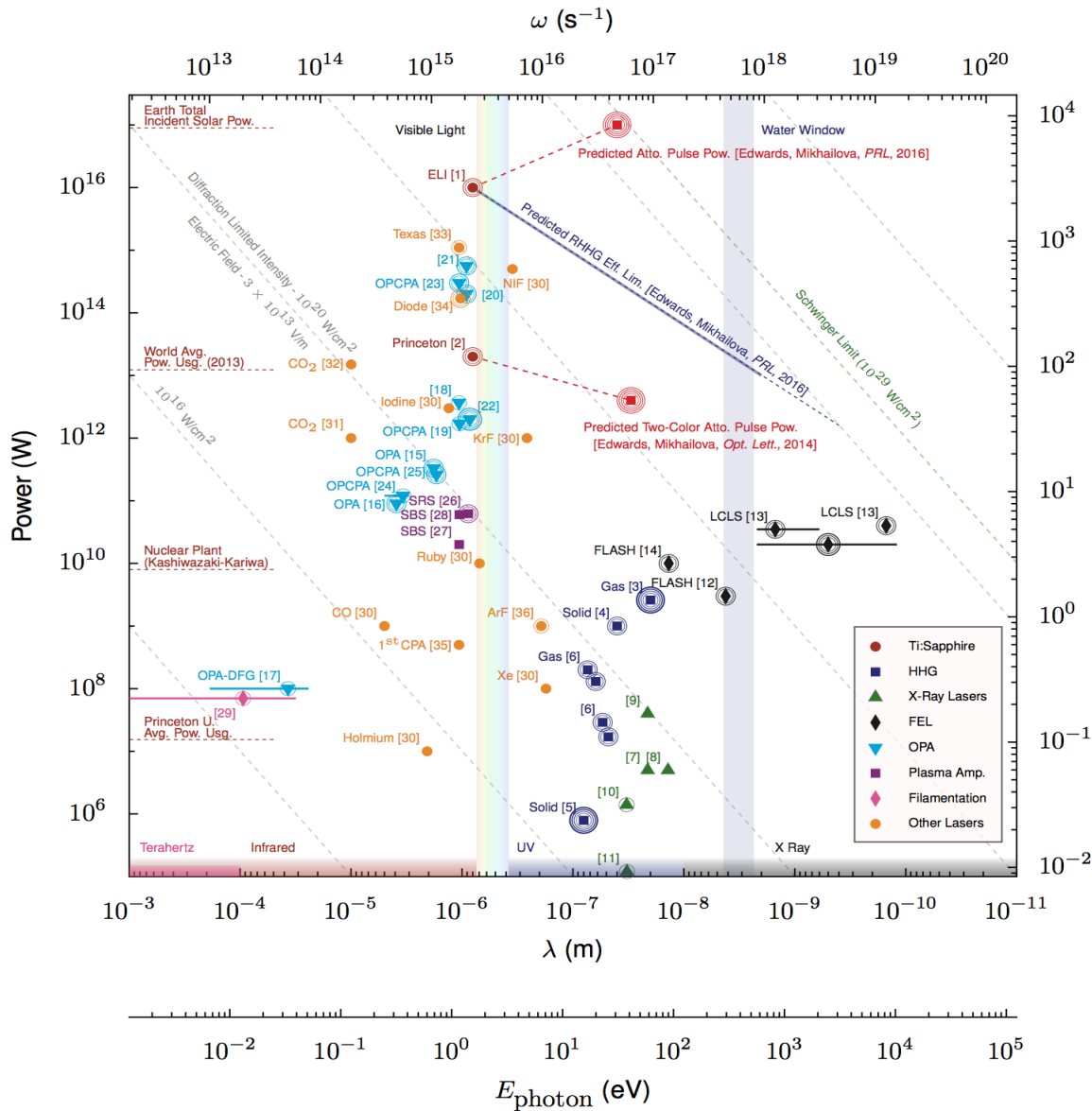
Filter out low frequencies



Attosecond Pulse

Image courtesy Matthew Edwards

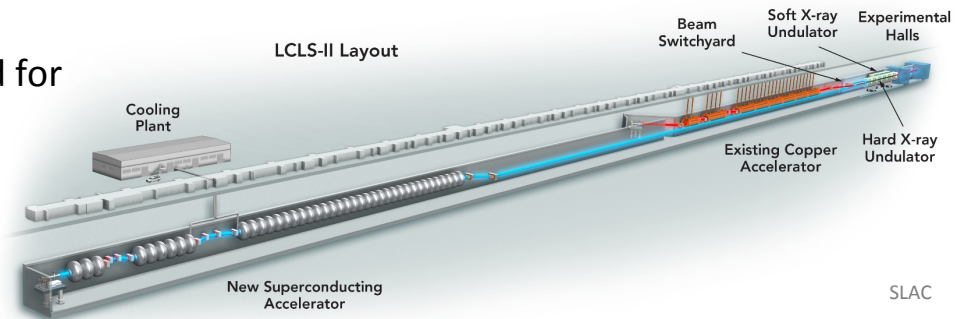
Ultrafast light sources



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X-ray Free Electron Lasers

- Capabilities (LCLS)
 - ~50 fs (pulse duration)
 - 40 GW (peak power)
 - 0.12 – 22 nm (wavelength)
 - Note that in principle FELs can be used for infrared through x-ray wavelengths.
 - 10^{18} W/cm² at 1 nm wavelength

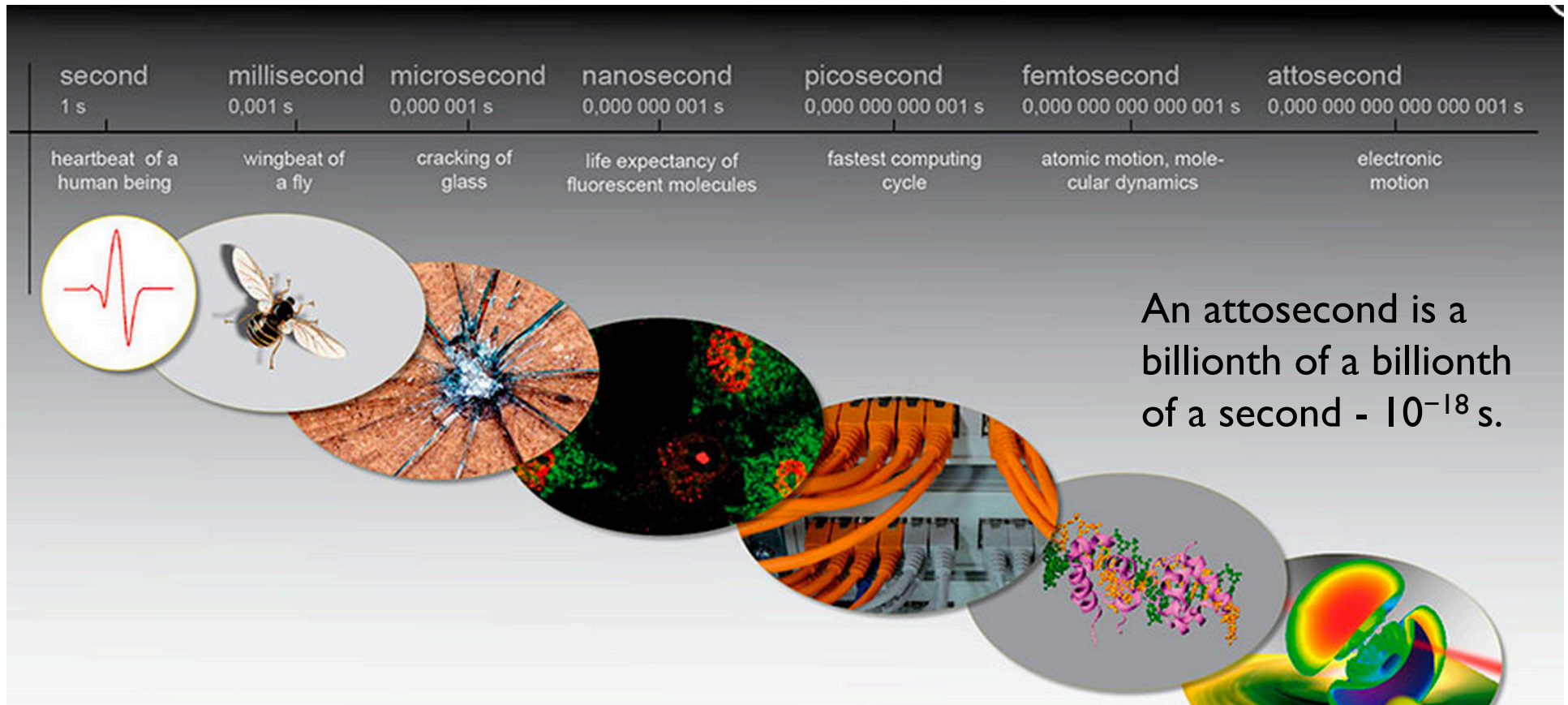


- Major Facilities
 - LCLS – SLAC National Accelerator Laboratory
 - FLASH – DESY (Hamburg)
 - European XFEL – DESY (Hamburg)
 - SACLA – RIKEN (Japan)



Development of compact sources of intense ultrafast x-rays

The time scale



An attosecond is a billionth of a billionth of a second - 10^{-18} s.



Age of the Universe is of the order of 10^{17} s



Human life is of the order of 10^9 s

Image: Ferenc Krausz, MPQ

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Thank you for your attention.